

Article

Synchrotron Radiation in Periodic Magnetic Fields of FEL Undulators—Theoretical Analysis for Experiments

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Received: 19 May 2020; Accepted: 3 July 2020; Published: 30 July 2020



Abstract: A theoretical study of the synchrotron radiation (SR) from electrons in periodic magnetic fields with non-periodic magnetic components is presented. It is applied to several free electron lasers (FELs) accounting for the real characteristics of their electron beams: finite sizes, energy spread, divergence etc. All the losses and off-axis effects are accounted analytically. Exact expressions for the harmonic radiation in multiperiodic magnetic fields with non-periodic components and off-axis effects are given in terms of the generalized Bessel and Airy-type functions. Their analytical forms clearly distinguish all contributions in each polarization of the undulator radiation (UR). The application to FELs is demonstrated with the help of the analytical model for FEL harmonic power evolution, which accounts for all major losses and has been verified with the results of well documented FEL experiments. The analysis of the off-axis effects for the odd and even harmonics is performed for SPRING8 Angstrom Compact free-electron LASER (SACLA) and Pohang Accelerator Laboratory (PAL-XFEL). The modelling describes theoretically the power levels of odd and even harmonics and the spectral line width and shape. The obtained theoretical results agree well with the available data for FEL experiments; where no data exist, we predict and explain the FEL radiation properties. The proposed theoretical approach is applicable to practically any FEL.

Keywords: undulator; free electron laser; harmonic generation; beam; off-axis effects

1. Introduction

The first theoretical results for the radiation of a charge on a circular orbit were obtained by Lienard in the end of the 19th century. A few years later in 1907 a complete theoretical study was performed by Schott. It was aimed on the description of the atomic spectra, but quantum mechanics was still unknown and this study failed to describe atoms. Instead it perfectly described the spectral and angular distribution of the radiation from electrons in a constant magnetic field. The results did not find application at that time and they were forgotten for almost half a century. The undulator radiation (UR) is a particular case of the synchrotron radiation in periodic magnetic field; it was proposed by Ginsburg [1] and first observed by Motz [2] in the middle of the 20th century. Ginsburg also claimed [1] that coherent radiation could be emitted by electrons in periodic magnetic fields; Ginsburg hypothesized that small groups of electrons, separated by the wavelength of the radiation, would emit coherent radiation. The real story of the synchrotron and undulator radiation began. Since then, the energy E of the electrons in accelerators has been shown to increase, so that now the relativistic factor $\gamma \gg 1$, $\gamma = E/mc^2 \sim 10^3 - 10^4$, where m is the electron mass and c is the speed of light. High demand for coherent radiation in X-ray band pushed towards building X-ray free electron lasers (FELs), which provide coherent radiation pulses in the Roentgen band with a duration of femtoseconds, which allow studies of ultra fast processes on nanoscale. To achieve this, the relativistic

beams and undulators must be aligned with micron scale tolerance and the whole installation must be manufactured with highest precision to ensure minimal off-axis deviation of the beam, maximum periodicity of the magnetic field and minimal influence of non-periodic magnetic constituents, which take electrons off their path in undulators and disrupt the coherence of their oscillations. Moreover, the relativistic electron beams must have low energy-spread and emittance, so that there could be efficient electron bunching. The latter occurs due to the interaction of the radiation with the beam; the electrons are grouped in microbunches, separated by the wavelength of the radiation and emit coherent radiation exactly as hypothesized by Ginzburg [1]. Modern FELs can generate gigawatt power in femtosecond pulses in the X-ray band [3–10].

Interest in theoretical studies of the synchrotron radiation in the 21st century is determined in great part by applications to FELs. Modeling of a FEL is usually done by numerical algorithms, which solve complex system of equations for electrons and radiation and describe the evolution of the electron beam and photon pulse. Such programs require computational facilities and qualified personnel for programming, specific for each FEL. While FEL physics is quite clear at the present stage, the particular reasons for the harmonic behavior in specific installations are not clarified by numerical simulations. In what follows we propose the analytical formalism, which is based on the exact solution for the radiation integral in complex multiperiodic magnetic fields with non-periodic constituents. The resulting integral forms of generalized special functions of Bessel- and Airy-types exactly describe the radiation in a given angle from electrons with initial off the axis position in the beam with finite emittance and energy spread. Moreover, the phenomenological model of a single-pass FEL, based on the logistic equation [11], has been redefined in the last decade and turned from a scholastic exercise to a real tool, capable of giving immediate analytical description of the power and bunching evolution in practically any FEL with arbitrary configuration of the undulator fields [12–17]. Its analytical formulation is rather simple and can be run by any user on a PC or even scientific calculator.

The novelty consists of a fully analytical description of the UR harmonic generation accounting for the off-axis effects, periodic and non-periodic magnetic field components, found in FEL undulators. The exact analytical results in terms of the generalized Bessel functions describe the effect of the periodic magnetic harmonics and of the angular effects and betatron oscillations in a wide beam; the Airy-type generalized functions describe the exact shape of the spectral line due to the non-periodic components. Coupled with the phenomenological FEL model, whose latest version is given in the Appendix A, we get the analytical description of the harmonic power evolution in FELs. Applications to a number of FEL experiments at SPRING8 Angstrom Compact free-electron LAsER (SACLA) and Pohang Accelerator Laboratory X-ray Free electron Laser (PAL-XFEL) are provided and discussed in the context of other facilities, such as at Linac Coherent Light Source (LCLS) etc.

2. Spontaneous UR intensity and Spectrum Distortions

The ideal undulator assumes pure sinusoidal magnetic field along the axis. In the ideal planar undulator only odd harmonics are radiated on the axis. However, the radiation spectrum of real undulators and FELs differs from the ideal: even harmonics appear in FEL experiments [18–22]. This is attributed to the non-ideally harmonic magnetic field and finite beam size. Theoretical estimation for even FEL harmonics in the experiment [21] at Advanced Photon Source's (APS) Low Energy Undulator Test Line (LEUTL) were based on the work [23]; however, they required the bunching values for the first and second harmonics, which in turn needed numerical simulations or theoretical calculations. Moreover, applying formulae of [21,23] to X-ray experiments, we systematically get harmonic powers ~25 times lower than those measured. To our best knowledge no convincing comprehensive theory has been provided thus far. In what follows we give analytical description of the synchrotron radiation from real electron beams in periodic magnetic fields of FEL undulators accounting for finite beam size and non-periodic magnetic components; the latter deviate electrons off the undulator axis. We demonstrate that this effect may exceed that of the finite size of the electron beam and that of the relevant betatron oscillations on the UR intensity. The UR, accounting for the magnetic field harmonics, has been recently

studied, for example, in [24,25]. It was concluded that reasonably strong field harmonics, ~30% of the main periodic field, still mostly influenced the saturation and gain lengths in FELs and had little effect on the FEL harmonic intensities. In what follows we will focus on the effect of weak non-periodic magnetic components $H_x = H_0\rho \ll 1$, $H_y = H_0\kappa \ll 1$. They are naturally caused by residual magnetic fields in undulators, magnetizing errors of constant magnets, by the field of the Earth, ~0.5 Gauss, and they are weak compared to the undulator field amplitude, $H_0 \sim 1$ Tesla. The proposed analytical formalism, however, is not limited to such weak fields $H_d/H_0 \sim 10^{-4}$, but allows arbitrary strengths.

As usual in classical electrodynamics, the calculations of the radiation intensity from an electron consist in the computation of the radiation integral:

$$\frac{d^2I}{d\omega d\Omega} = \frac{e^2}{4\pi^2 c} \left| \omega \int_{-\infty}^{\infty} dt \left[\vec{n} \times \left[\vec{n} \times \vec{\beta} \right] \right] \exp[i\omega(t - \vec{n} \cdot \vec{r}/c)] \right|^2, \quad (1)$$

where the notations are common to SR and UR theories: \vec{n} is the unit-vector from the electron to the observer, \vec{r} is the electron radius-vector, $\vec{\beta}$ is its velocity and c is the speed of light. For the sake of generality, we consider the radiation from an electron in the two-dimensional bi-harmonic multiperiodic field:

$$\vec{H} = H_0(\sin(k_\lambda z) + d \sin(pk_\lambda z), d_1 \sin(hk_\lambda z) + d_2 \cos(lk_\lambda z), 0), \quad (2)$$

$$k_\lambda = 2\pi/\lambda_{u,x}, \lambda_{u,x} \equiv \lambda_u, h, l, p \in \text{integers}, d, d_1, d_2 \in \text{reals}.$$

The account for the third field harmonic usually allows good reconstruction of the field for a given radiation pattern [26]. Not limiting ourselves to the third harmonic, the proposed analytical formalism allows field harmonics of arbitrary strength and order. The calculations go along the lines of [27]; the integrand and the exponential of the radiation integral (1) are expanded in series of the small parameter $1/\gamma \ll 1$, which naturally arises in the relativistic limit because high-energy electron beams are used in FEL installations. For the two-dimensional field (2) proper formulae are much more cumbersome than in [27], but the approach remains the same: the non-oscillating terms in the exponential in (1) yield the resonances of the UR; their resonant wavelengths are expressed as follows:

$$\lambda_n = \frac{\lambda_u}{2n\gamma^2} \left(1 + \frac{k^2}{2} \omega + (\gamma\Theta)^2 \right), \quad (3)$$

where the account for the multiple periods of the undulator is given by $\omega = 1 + \left(\frac{d}{p}\right)^2 + \left(\frac{d_1}{h}\right)^2 + \left(\frac{d_2}{l}\right)^2$, and the angle $\Theta^2 = \theta^2 + \theta_H^2 - \sqrt{3}\theta_H\theta \frac{\rho \sin\varphi - \kappa \cos\varphi}{\sqrt{k^2 + \rho^2}}$ includes the usual off-axis angle θ and the effective bending angle θ_H , which describes the effect of the non-periodic magnetic components. The purely periodic terms in the exponential of the radiation integral are collected and form the generalized Bessel-type functions $J_n^m(\xi_i)$, which naturally arise in the following form:

$$J_n^m(\xi_i) = \int_0^{2\pi} \frac{d\alpha}{2\pi} \cos \left[i \left(\begin{array}{l} n\alpha + \xi_1 \sin(h\alpha) + \xi_2 \cos(l\alpha) + \xi_3 \sin\alpha + \xi_4 \sin(2\alpha) \\ + \xi_5 \sin(2h\alpha) + \xi_6 \sin(2l\alpha) + \xi_7 \cos((l+h)\alpha) + \xi_8 \cos((l-h)\alpha) \\ + \xi_0 \sin(p\alpha) + \xi_9 \sin((p+1)\alpha) + \xi_{10} \sin((p-1)\alpha) + \xi_{11} \sin(2p\alpha) \end{array} \right) \right], \quad (4)$$

where:

$$\xi_0 = \xi_4 \frac{8d}{kp^2} \gamma\theta \sin\varphi, \xi_1 = \xi_4 \frac{8d_1}{kh^2} \gamma\theta \cos\varphi, \xi_2 = \xi_4 \frac{8d_2}{kl^2} \gamma\theta \cos\varphi, \xi_3 = \xi_4 \frac{8}{k} \gamma\theta \sin\varphi, \quad (5)$$

$$\xi_4 = \frac{mk^2/4}{1 + \frac{k^2}{2} \left(1 + \left(\frac{d}{p}\right)^2 + \left(\frac{d_1}{h}\right)^2 + \left(\frac{d_2}{l}\right)^2 \right) + \gamma^2 \theta^2}, \xi_5 = \frac{\xi_4 d_1^2}{h^3}, \xi_6 = -\frac{\xi_4 d_2^2}{l^3}, \quad (6)$$

$$\xi_7 = \frac{\xi_4 4d_1 d_2}{hl(l+h)}, \xi_8 = \frac{\xi_4 4d_1 d_2}{hl(l-h)}, \xi_9 = \frac{\xi_4 4d}{p(p+1)}, \xi_{10} = \frac{\xi_4 4d}{p(p-1)}, \xi_{11} = \frac{\xi_4 d^2}{p^3}, \quad (7)$$

$k \cong \lambda_u [cm] H_0 [kG] / 10.7$ is the main undulator parameter, θ is the off-axis angle and φ is the azimuthal angle. We assume multiple field harmonics, l, h, p in (2). The additional constant magnetic field components $H_x = H_0 \rho$, $H_y = H_0 \kappa$, $H_z = H_0 \varsigma$ can affect the undulator. In relativistic beams the longitudinal constant component $H_z = H_0 \varsigma$ can be neglected and the transversal field $H_d = H_0 \sqrt{\kappa^2 + \rho^2}$ plays major role. It gives rise to non-periodic components [28–30] in the exponential of the radiation integral, such as $\propto \left(\kappa \frac{\omega}{\omega_0} \frac{k}{\gamma} \theta \cos \varphi + \rho \frac{\omega}{\omega_0} \frac{k}{\gamma} \theta \sin \varphi \right) (\omega_0 t)^2$, $\propto \frac{\omega}{\omega_0} \left(\frac{k}{\gamma} \right)^2 (\kappa^2 + \rho^2) (\omega_0 t)^3$ etc., where $\omega_0 = \frac{2\pi c}{\lambda_u}$; similar terms appear for the field harmonics, involving $(p, h, l) \omega_0$. The effect of the field H_d is accumulated along the undulator length $L = \lambda_u N$, where N is the total number of periods, and quantified by the normalized bending angle θ_H :

$$\gamma \theta_H = \frac{2\pi}{\sqrt{3}} k N (\kappa^2 + \rho^2)^{1/2} \cong \frac{2\pi}{107 \sqrt{3}} L [m] H_d [G]. \quad (8)$$

Physically the constant magnetic field H_d bends the electron trajectory into the effective angle $\gamma \theta_H$ and causes synchrotron radiation from much wider curve, than that of the electron oscillations along the undulator periods. However, its effect in long undulators should not be underestimated, as we will show in what follows. The non-periodic magnetic components in the exponential of the radiation integral compose the following ad-hoc generalized Airy-type function in the integral form:

$$S(v_n, \eta, \beta) \equiv \int_0^1 d\tau e^{i(v_n \tau + \eta \tau^2 + \beta \tau^3)}, \quad (9)$$

where $v_n = 2\pi n N ((\omega/\omega_n) - 1)$ is the detuning parameter, describing the deviation of the frequency ω from the UR resonances $\omega_n = 2\pi c/\lambda_n$,

$$\beta \cong \frac{2\pi n N (\gamma \theta_H)^2}{1 + k^2/2}, \quad \eta = \frac{4\pi^2 n N (kN) \gamma \theta (\kappa \cos \varphi - \rho \sin \varphi)}{(1 + k^2/2)}; \quad (10)$$

where for practical evaluations we can use $kN \cong 0.934 L [cm] H_0 [Tesla]$. The special function S can be expressed as the action of the operational differential operators, also employed for the studies of Hermite and Laguerre families of orthogonal polynomials in [31–33]. The generalized multivariable Hermite polynomials:

$$H_n(x, y) = e^y \partial_x^2 x^n, H_n(x, y, z) = e^y \partial_x^2 + z \partial_x^3 x^n, \quad (11)$$

were studied operationally by Srivastava et al. (see, for example, [34–37]). Exponential differential operators provide the link between S function (9) and sincx function, which describes the shape of the ideal UR spectrum line. The generalization (9) in the non-periodic magnetic field can be given by the following operational relation:

$$S(x, y, z) \equiv \int_0^1 d\tau e^{i(x \tau + y \tau^2 + z \tau^3)} = e^{-i y \partial_x^2 - z \partial_x^3} \int_0^1 e^{i x \tau} d\tau = e^{-i y \partial_x^2 - z \partial_x^3} \left(\frac{\sin x/2}{x/2} e^{i \frac{x}{2}} \right), \quad (12)$$

$$S(v_n, 0, 0) = e^{i v_n / 2} \text{sinc}(v_n / 2). \quad (13)$$

The generalized Airy function S is related to the generalized three-variable Hermite polynomials as follows:

$$S(x, y, z) = \sum_{n=0}^{\infty} \frac{i^n H_n(x, -iy, -z)}{(n+1)!}; \quad (14)$$

$H_n(x, y, z)$ can be expressed as sums of the Hermite polynomials of two variables

$$H_n(x, y, z) = n! \sum_{r=0}^{\lfloor n/3 \rfloor} \frac{z^{n-3r}}{(n-3r)! r!} H_n(x, y); \quad (15)$$

$H_n(x, y)$ are just another form of writing for common Hermite polynomials:

$$H_n(x, y) = (-i)^n y^{n/2} H_n\left(\frac{ix}{2\sqrt{y}}\right) = i^n (2y)^{n/2} He_n\left(\frac{x}{i\sqrt{2y}}\right); \quad (16)$$

they have the following sum presentation [38]:

$$H_n(x, y) = n! \sum_{r=0}^{\lfloor n/2 \rfloor} \frac{x^{n-2r} y^r}{(n-2r)! r!} \quad (17)$$

Hermite polynomials $H_n(x, y, z)$ and $H_n(x, y)$ possess the generating exponents:

$$\sum_{n=0}^{\infty} \frac{t^n}{n!} H_n(x, y, z) = e^{xt+yt^2+zt^3}, \quad \sum_{n=0}^{\infty} \frac{t^n}{n!} H_n(x, y) = e^{xt+yt^2}. \quad (18)$$

On the undulator axis the second argument of the generalized Airy function S vanishes and $S \equiv S(v_n, \beta, \eta)$ simplifies:

$$S(x, y, z)|_{\text{on-axis}} = S(x, z) = \int_0^1 e^{i(x\tau + z\tau^3)} d\tau = \sum_{m=0}^{\infty} \frac{i^m H_m(x, -z)}{(m+1)!}. \quad (19)$$

The effect of the non-periodic magnetic field is quantified by the induced angle θ_H in β ; the dependence on the off-axis angle θ is in η . The maximum values $\max[S] = 1$ and $\max[\partial S / \partial v_n] = 0.5$ explain why the coefficient 2 is grouped with $\partial S / \partial v_n$ in (21). Upon computing the radiation integral we get the UR intensity:

$$\frac{d^2 I}{d\omega d\Omega} \cong \frac{e^2 N^2 \gamma^2 k^2}{c(1+k^2/2)^2} \sum_{n=-\infty}^{\infty} n^2 I_n^2, \quad (20)$$

where the intensity of the n -th UR harmonic reads as follows:

$$I_n = \sqrt{|S(f_n^1 + f_n^2)|^2 + |(2\partial S / \partial v_n) f_n^3|^2}. \quad (21)$$

The generalized Airy function $S \equiv S(v_n, \beta, \eta)$ describes the shape of the spectrum line, distorted by the non-periodic magnetic field; the shape of the line for odd UR harmonics is given by function S , of the even harmonics—by $\partial S / \partial v_n$. The Bessel coefficients $f_n^{1,2}$ give the amplitudes of proper UR harmonics x - and y -polarizations and are expressed in terms of the generalized Bessel functions $J_n^m \equiv J_n^m(\xi_i(m))$ (4) as follows:

$$f_{n;x}^1 = \frac{d_1}{h} (J_{n+h}^n + J_{n-h}^n) + i \frac{d_2}{l} (J_{n+l}^n - J_{n-l}^n), \quad f_{n;x}^2 = \frac{2}{k} \gamma \theta \cos \varphi J_n^n, \quad (22)$$

$$f_{n;y}^1 = (J_{n+1}^n + J_{n-1}^n) + \frac{d}{p} (J_{n+m}^n + J_{n-m}^n), \quad f_{n;y}^2 = \frac{2}{k} \gamma \theta \sin \varphi J_n^n. \quad (23)$$

$$f_n^3 = \sqrt{3} \gamma \theta_H J_n^n / k. \quad (24)$$

Formulas (22)–(24) account for the off-axis angle θ and for the non-periodic magnetic field H_d , written in terms of the bending angle θ_H . For the odd UR harmonics $n = 1, 3, 5, \dots$ mainly the Bessel coefficient f_n^1 determines the UR intensity (20). The resonance of the UR has an infrared shift respectively to the ideal value at $v_n = 0$:

$$\omega_n = 2n\omega_0\gamma^2 / \left(1 + \frac{k^2}{2}\omega + (\gamma\theta)^2 + (\gamma\theta_H)^2 - \gamma^2 \sqrt{3}\theta_H\theta \frac{\rho \sin \varphi - \kappa \cos \varphi}{\sqrt{\kappa^2 + \rho^2}} \right). \quad (25)$$

Interestingly, the effective angle θ_H and the off-axis angle θ can counteract each other's effect on the radiation. The best compensation occurs for:

$$v_n \approx -(\beta + \eta), \text{ in the range } v_n, \beta, \eta \in [-2\pi, 2\pi]. \quad (26)$$

It follows from (26) that $v_n = 0$ for $\theta = \frac{2\pi}{3} \frac{k}{\gamma} N \frac{\kappa^2 + \rho^2}{\rho \sin \varphi - \kappa \cos \varphi}$. In the simplest case of one-dimensional magnetic field $H_d = \kappa H_0$, for $\varphi = \pi$ we get the angle $\tilde{\theta}$, in which the infrared shift of the received radiation is compensated: $v_n = 0$ for $\tilde{\theta} = \frac{2\pi}{3} \frac{k}{\gamma} N \kappa = \sqrt{3}\theta_H$; proper UR resonances are $\omega_n \cong 2n\omega_0\gamma^2 / \left(1 + (\omega k^2/2) + 0.27(\gamma\theta_H)^2 \right)$. The examples of the UR lines for the PAL-XFEL [39] undulator with $N = 194$ periods, $k = 1.87$, period $\lambda_u = 2.57$ cm, length $L = 5$ m and the electron energy spread $\sigma_e = 1.8 \times 10^{-4}$, is shown in Figure 1a for $\gamma\theta = 0$ and Figure 1b for $\gamma\theta \neq 0$ in the presence of the non-periodic magnetic component H_d . In Figure 1a, the field H_d causes an infrared shift and broadens the spectrum line, viewed in zero angle $\gamma\theta = 0$. In Figure 1b, note that the same field H_d can improve the shape of the spectrum line viewed in the angle $\gamma\theta = 0.067$. Note in Figure 1b as the initial detuning, caused by the off-axis angle $\gamma\theta = 0.07$, reduces if the undulator is affected by the field $H_d = \kappa H_0 \sim 10^{-4} H_0$; a further increase of κ broadens the UR line. For the on the axis case, $\gamma\theta = 0$, the effect of H_d is purely detrimental (see Figure 1a).

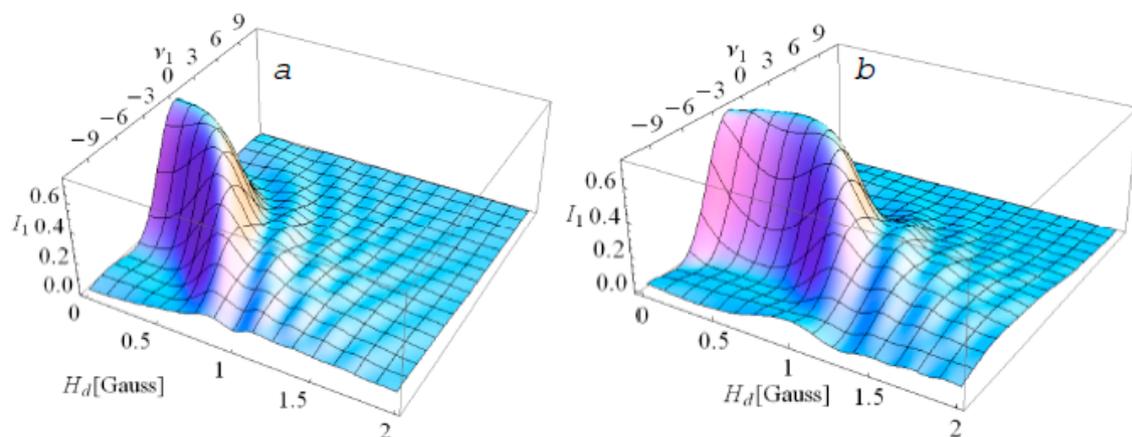


Figure 1. The PAL-XFEL fundamental UR line shape as the function of the constant magnetic field H_y : (a)—on-the-axis case $\gamma\theta = 0$, (b) off-axis angle $\gamma\theta = 0.067$; the undulator has $k = 1.87$, period $\lambda_u = 2.57$ cm, length $L = 5$ m, $N = 194$ periods.

The effect of the off-axis angles, the constant magnetic field H_d and their interplay are also shown for the fundamental harmonic in Figure 2. In Figure 2a, the symmetric ideal UR line of the fundamental tone is described by the $\text{sinc}(v_n/2)$ function for $\gamma\theta = 0$, $H_d = 0$; it shows a red shift, if viewed in the angle $\gamma\theta \neq 0$; the angles $\gamma\theta > 0.1$ cause a significant shift down from $v_n = 0$ and the intensity slightly decreases.

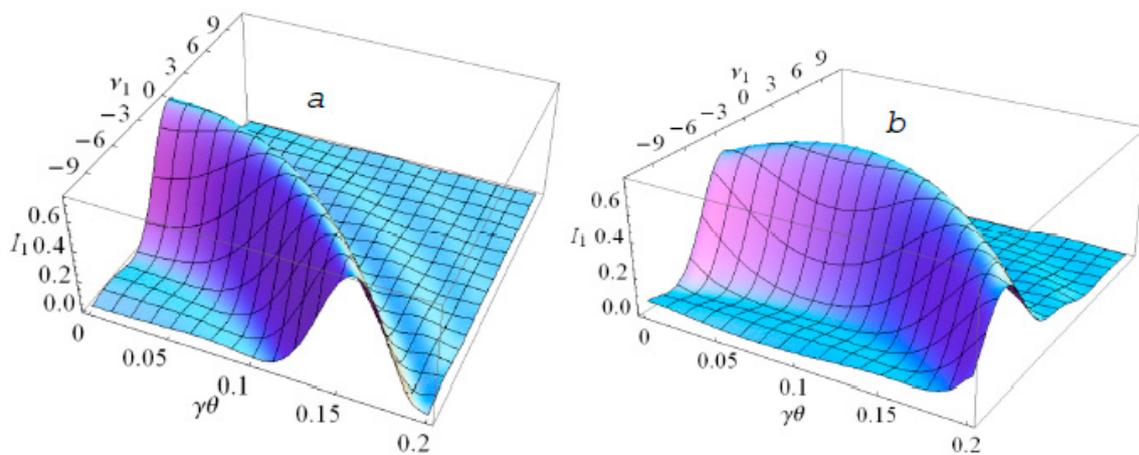


Figure 2. The PAL-XFEL fundamental UR line shape as the function of the off-axis angle $\gamma\theta$; (a)—no constant magnetic component, $\kappa = 0$, (b)—in the presence of the constant magnetic component $H_d = 0.5$ Gauss; the undulator has $k = 1.87$, period $\lambda_u = 2.57$ cm, length $L = 5$ m, $N = 194$ periods.

The radiation line of the fundamental harmonic $n = 1$, viewed in the angle $\gamma\theta \approx 0.1$, is slightly broadened and has red shift in -2π with respect to the resonance $\nu_n = 0$ (see Figure 2a). In Figure 2b we demonstrate the spectral line of the fundamental tone, broadened and red-shifted in $-\pi$ by the constant magnetic field $H \approx H_0 \times 10^{-4}$, if viewed on the axis, $\gamma\theta = 0$. In Figure 2b, the spectral line reassumes a more distinct shape with the increase of the off-axis angle $\gamma\theta$ from zero to ~ 0.1 . This demonstrates that the non-periodic magnetic component κH_0 and the off-axis angle $\gamma\theta$ can compensate each other's effect on the UR. In the presence of the field $H_d \approx 10^{-4} H_0$, the spectrum line gets narrower and the red shift is smaller for the same angles $\gamma\theta \approx 0.1$ as shown in Figure 2b.

Our theoretical analysis and (22)–(24) in particular allow the analytical study of even UR harmonics. For the even harmonics the contributions from $f_n^{1,2,3}$ can be of the same order of magnitude; we distinguish and analyze separately the terms $f_n^{1,2,3}$, factorized by S and $\partial S / \partial \nu_n$ in (21). Note the value $\max[\partial S / \partial \nu_n] = 0.5$. Upon the comparison of f_n^3 (24) with f_n^2 in (22), (23), we notice that the role of the angle θ_H in f_n^3 is formally the same as the role of the off-axis angle θ in f_n^2 , i.e., θ_H is involved in f_n^3 the same way as θ is involved in f_n^2 . Moreover, accounting for the factor $\sqrt{3}$ in (24), we get similar numeric factors in the Bessel coefficients f_n^3 and f_n^2 : $f_n^3 = \frac{\sqrt{3}}{k} \gamma \theta_H \frac{2\partial S}{\partial \nu_n} J_n^n \approx \frac{1.73}{k} \gamma \theta_H J_n^n$ vs. $f_n^2 \approx \frac{2}{k} \gamma \theta J_n^n$. The latter is a typical off-axis term in (22), (23). The angle θ_H is induced by the field $H[kG] \approx \frac{3\gamma\theta_H}{\lambda_u[cm]N}$. In order to have noticeable effect, there must be $\gamma\theta_H \approx 0.05$ – 0.1 . In a long undulator, this angle can be generated by rather weak magnetic field H_d : for example, the bending angle $\gamma\theta_H = 0.05$ can have noticeable effect on the FEL performance at LCLS, where the off-axis target deviation was $5 \mu\text{m}$. LCLS undulator length, $L = 3.4$ m, is translated to the field strength $H = 0.44$ Gauss. The need to screen out such fields was pointed out in [18,19]. Moreover, at the PAL-XFEL the field of the Earth, ~ 0.5 Gauss, can induce the angle $\gamma\theta_H \approx 0.08$ in the 5 m long undulator, and cause even stronger deviation of the electron trajectories.

The contributions to the intensity of the second harmonic of PAL-XFEL undulator is demonstrated in Figures 3–5, where we have assumed the off-axis angle $\gamma\theta = 0.1$ and low energy spread $\sigma_e = 10^{-4}$. The contribution to the second harmonic intensity due to the constant field κH_0 (24) accounting for the off-axis angle θ is shown in Figure 3. The shape of the spectrum line is given by $\partial S / \partial \nu_n$. The term f_n^3 increases with the increase of the bending field H_d . The interplay with the angles $\gamma\theta$ (see Figure 3b) limits the increase of f_n^3 and determines its behavior for the stronger field $H_d > 1.5 \times 10^{-4}$ Gauss. This latter value depends on $\gamma\theta$ and on the undulator parameters; for $\theta = 0$, the Bessel coefficient f_n^3 grows further for increasing κH_0 (see Figure 3a).

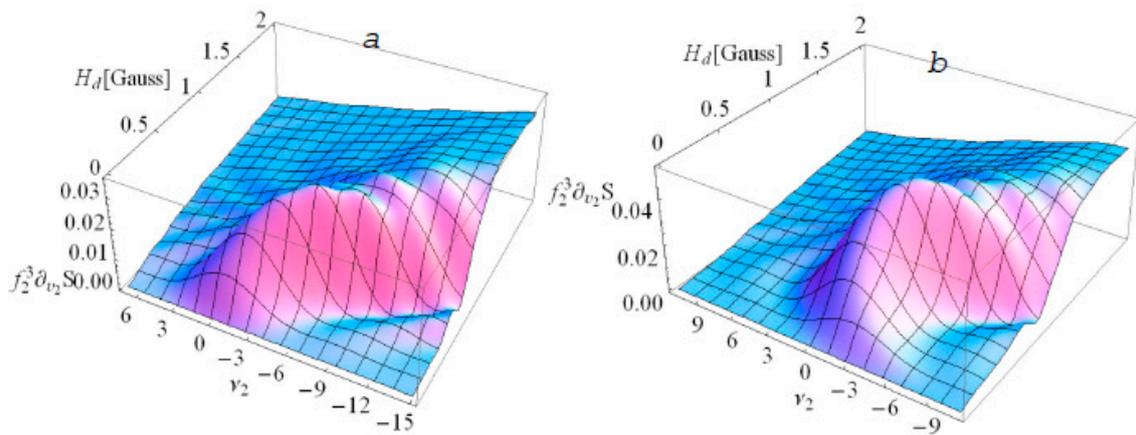


Figure 3. The contribution of the constant magnetic field term f_2^3 to the UR line of the second UR harmonic, $n = 2$, of the PAL-XFEL undulator with $k = 1.87$, period $\lambda_u = 2.57$ cm, length $L = 5$ m, $N = 194$ periods: (a)—on the axis case, $\gamma\theta = 0$, (b)—the off-axis angle $\gamma\theta = 0.067$.

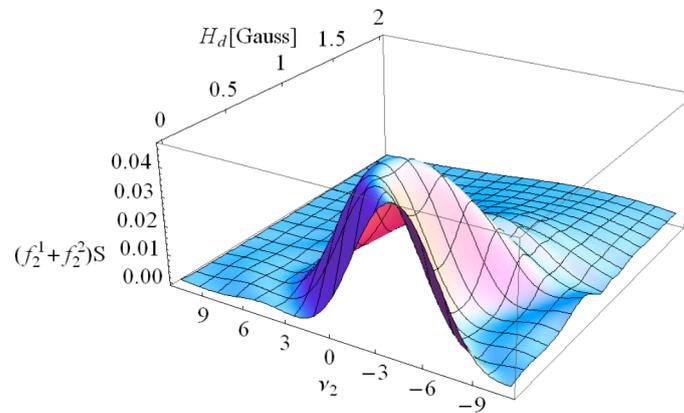


Figure 4. The contribution of the terms f_2^1 and f_2^2 to the UR line of the second UR harmonic, $n = 2$, of the PAL-XFEL for the undulator with $k = 1.87$ period 2.57 cm, length 5 m, $N = 194$ periods, off-axis angle $\gamma\theta = 0.067$.

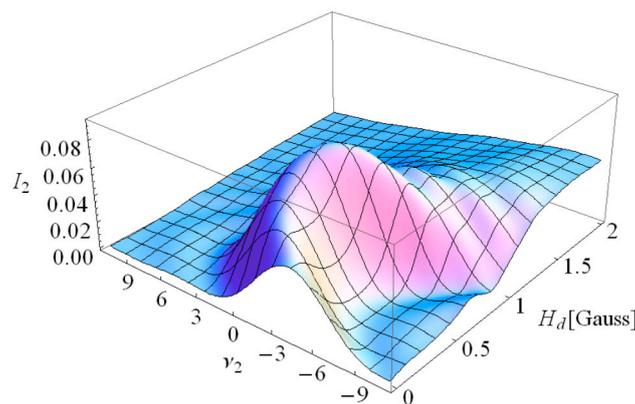


Figure 5. The total contribution of all terms $f_2^{1,2,3}$ to the normalized intensity I_2 of the harmonic $n = 2$ accounting for the functions S and $\partial S / \partial \nu_n$, giving the line shape to 2nd UR harmonic of the PAL-XFEL undulator with $k = 1.87$, period $\lambda_u = 2.57$ cm, length $L = 5$ m, $N = 194$ periods for the off-axis angle $\gamma\theta = 0.067$.

The contributions of the terms $f_n^{1,2}$ for the second UR harmonic for $\gamma\theta = 0.067$ are shown in Figure 4; they decrease with the increase of the constant field strength κH_0 ; the associated spectrum line shape is described by the function S . The comprehensive contribution of all terms to the normalized intensity

of the second harmonic, I_2 , is shown in Figure 5; the maximum intensity is at $H_d \approx 0.7 \times 10^{-4}$ Gauss. The decrease of the UR intensity, caused by $\gamma\theta \sim 0.067$, is compensated by the field $H_d \approx 0.7 \times 10^{-4}$ Gauss, and at this point, the second harmonic intensity is at its maximum.

The electron energy spread in the beam, σ_e , is accounted for by the convolution $\int_{-\infty}^{\infty} \frac{d^2 I(v_n + 4\pi n N \epsilon)}{d\omega d\Omega \sqrt{2\pi\sigma_e}} e^{-\frac{\epsilon^2}{2\sigma_e^2}} d\epsilon$. The effect of the energy spread on the UR harmonics is purely detrimental and it causes symmetric broadening of the spectral lines. In this context it is important to underline that high radiation harmonics are more sensitive to the energy spread and to other loss factors, than the fundamental harmonic. Weak, but detectable at low energy spread, $\sigma_e \approx 10^{-4}$, FEL harmonics can be almost totally suppressed, if the energy spread increases to $\sigma_e \approx 10^{-3}$. The relevant example of SACLAR radiation will be considered in what follows.

Eventually, let us evaluate the effect of the betatron oscillations in the finite width of the electron beam, where the electrons enter the undulator off the undulator axis. This topic has been in focus of researchers' attention since the first accelerators were built in the middle of the 20th century. It is well described in various articles and books (see, for example, [9,40–44]). The field between the arrays of the planar undulator magnets is better approximated by the magnetic components $H_x = H_0 \sin(k_\lambda z) \cosh(k_\lambda y)$, $H_z = H_0 \cos(k_\lambda z) \sinh(k_\lambda y)$, which satisfy Maxwell equations in the whole gap between the magnets. The radiation in a two-frequency planar undulator with proper field was considered in [44]; rigorous calculations were explicitly presented there. For the multiharmonic undulator field (2), we follow the approach of [44]; cumbersome calculations do not differ in principle from those in [44]. The transversal oscillations of the electron in the finite sized beam are described by the betatron frequency in its usual form [44]:

$$\omega_\beta = \frac{\sqrt{2}\pi c k \delta}{\lambda_n n \gamma} = \frac{2\sqrt{2}\pi c \gamma k \delta}{\lambda_u (1 + (\omega k^2/2))}, \quad (27)$$

where $\delta = 1$ for the common planar undulator, $\delta = \sqrt{1 + d^2}$ for the bi-harmonic field $H_y = H_0(\sin(k_\lambda z) + d \sin(pk_\lambda z))$ and for the multiharmonic undulator field (2) we get $\delta = \sqrt{1 + d^2 + d_1^2 + d_2^2}$. The betatron frequency ω_β (27) is much lower than the UR frequency $\omega_n \cong \frac{4\pi c n \gamma^2}{\lambda_u (1 + (k^2/2))}$; their ratio is roughly the inverse of the relativistic factor: $\frac{\omega_\beta}{\omega_n} \cong \frac{k\delta}{\sqrt{2}n\gamma} \propto \frac{1}{\gamma}$. This explains the high interest to this topic already in early SR and UR experiments, where relatively low-energy electron beams were used and the contribution of the betatron oscillations was considerable. For the intensity of the UR harmonic n , accounting for the betatron oscillations, we get the following expression:

$$I_n = \sqrt{\sum_{p=-\infty}^{\infty} \left\{ \tilde{J}_p(\xi, \zeta) \left(|S(f_n^1 + f_n^2)|^2 + |(2\partial S/\partial v_n) f_n^3|^2 \right) + |S f_{n,p;y}^4|^2 \right\}}, \quad (28)$$

where $f_n^{1,2,3}$ are given by (22)–(24), and the Bessel functions:

$$\tilde{J}_p(\xi, \zeta) = \frac{1}{2\pi} \int_0^{2\pi} \cos(pq - \zeta \sin q - \xi \sin 2q) dq, \quad (29)$$

depend on the arguments:

$$\xi = \frac{\pi^2 y_0^2 k \delta}{2n\gamma \sqrt{2}\lambda_u \lambda_n} = \frac{\pi^2 \gamma y_0^2 k \delta}{\sqrt{2}\lambda_u^2 (1 + (k^2/2))}, \quad \zeta = \frac{2\pi\theta y_0}{n\lambda_n} = \frac{4\pi\theta y_0 \gamma^2}{\lambda_u (1 + (k^2/2))}, \quad (30)$$

where y_0 is the off-axis position of the electron in the beam and θ is off-axis angle. The summation series $\sum_{p=-\infty}^{+\infty}$ over p describe the account for all subharmonics p of the harmonic n . In real devices, finite

number q of the subharmonics contribute: $\sqrt{\sum_{p=-q}^{+q} \tilde{J}_p^2} \cong 1$; where q describes the degree of the split of the harmonic n and depends on the beam parameters; it varies strongly from one installation to another. Some examples will be considered in the following section, where we model some FEL experiments. The subharmonics are distant at the betatron frequency ω_β . In the relativistic beams, $\gamma \gg 1$, this split of the UR lines due to the betatron oscillations is small: $\omega_\beta \propto \omega_n / \gamma$. The even UR harmonics appear on the undulator axis due to the betatron oscillations [9,40–44]; proper Bessel coefficient expectably differs from that in [44] only in Bessel functions due to different undulator field (2):

$$f_{n,p;y}^4 \cong \frac{\sqrt{2}\pi y_0 \delta}{\lambda_u} (\tilde{J}_{p+1}(\xi, \zeta) - \tilde{J}_{p-1}(\xi, \zeta)) J_n^4(\xi_i), \quad (31)$$

where n is the number of the UR harmonic and p is the number of the betatron subharmonic. The physics and the approach with regard to the betatron oscillations remain the same for any undulator. For the bi-harmonic planar undulator $d_1 = d_2 = 0$, and the result (31) reduces to that in [44] in different notations. For the common planar undulator with single field harmonic $H_0 \sin(k_\lambda z)$, $d = d_1 = d_2 = 0$, the Bessel functions $J_n^4(\xi_i)$ in (31) reduce to $\tilde{J}_n(\bar{\xi}, \bar{\zeta})$ (29), whose arguments are $\bar{\zeta} = \frac{8\bar{\xi}\gamma\theta}{k}$, $\bar{\xi} = -\frac{1}{8} \frac{k^2 \lambda_u}{\gamma^2 \lambda_n} = -\frac{1}{4} \frac{nk^2}{1+(k^2/2)}$. Quantitative evaluation of the Bessel coefficients shows that the contribution of the betatron oscillations is usually small: $f_{n,p;y}^4 \sim 10^{-2}$, in comparison with other Bessel coefficients: ~ 0.15 – 0.8 in (22), (23); usually $f_{n,p;y}^4$ do not exceed $f_{n=1,3,5}^{1,2}$. However, the split of the spectrum lines due to the betatron oscillations can be considerable and it strongly depends on the parameters of the installation and on the beam. Some examples are given in the context of the modeling of FELs in the following section. Beam sizes vary from ~ 0.2 mm to ~ 20 μm in modern FELs; beam deviations from the axis are usually small; for example, they are ~ 5 – 25 μm on one gain length, $L_g = 1.6$ – 3.5 m, in the LCLS FEL experiments [18–20]. However, the off-axis deviation of ultrarelativistic electrons in just ~ 10 μm in one undulator section length, $\cong 3$ m, can cause the effective angle $\gamma\theta \sim 0.1$ and noticeable effects. In what follows we will analyze in detail the harmonic generation in SACLA and PAL-XFELs and compare them with some other user facilities, such as LCLS.

3. Analysis of the Harmonic Generation in Some FEL Experiments

To calculate the harmonics in high gain FELs, we use the above obtained analytical results and the phenomenological FEL model (see the Appendix A). The latter is based on the Pierce parameter ρ [45–50], and it significantly develops the first approach in [51–53], applied with some modifications in [54–60]. The present formulation in contrast with [51–53] describes all losses separately for each FEL harmonic; it also includes the gradual dual-stage saturation and reproduces the oscillations of the saturated power, differently from that in [54–60]. Moreover, formulae (4)–(7), (22), (23) above in Section 2 in the limiting cases of the planar 2-frequency undulator correct the misprinted results in [54–60].

3.1. SACLA FEL Experiment

The SACLA facility first produced coherent radiation with 10 keV photons in 2011 [61]. User operations began in 2012; hard X-ray line BL2 was installed in 2014; soft X-ray line BL1 [62] and the dedicated accelerator SCSS+ were installed in 2014; further upgrades [63] followed, a new BL3 line was installed in 2017 for multiple beamline operation [64] with two-color XFEL and self-seeded XFEL [65]. However, contrary to exhaustive description of user facilities and generic specifications of the range of the parameters, in which SACLA operates, there has been little theoretical analysis of the FEL radiation in the experiments. Moreover, the information available on some instances of the operation of this FEL is incomplete and even controversial. In particular, this regards the electron beam energy spread, which is not stated for the hard X-ray SACLA setup in the papers [64–68] describing the facility. Rather complete data are available only for the undulators for the hard X-rays, commissioned [69] in 2012.

As we understand, the facility was upgraded several times since then, but the data for the beam characteristics were not clearly reported. Moreover, for the same year 2012 we find the energy spread 10^{-4} and betatron value 30 m in [69], while in [61] we find the energy spread (in projection) $<10^{-3}$ and $\beta_{x,y} = 22$ m. One order of magnitude difference in the energy spread together with the change in β from 30 m to 22 m has very strong effect of the FEL radiation: the saturation length can vary from ~ 20 m to ~ 60 m, the saturated powers change etc.

Consider first the soft X-ray FEL; it operates with three undulator sections with variable parameter $k \in [0.5 - 2.1]$ and a total of 777 periods, each $\lambda_u = 1.8$ cm long. The coherent radiation is generated at the fundamental $\lambda \sim 1-12$ nm; the details are available in [62]. Despite explicit description, [62] does not contain any data on the power evolution, saturation and gain lengths, although the beam and radiation characteristics are well specified. We have studied the instance of this experiment with the maximum possible value of $k = 2.1$, the electron energy $E = 780$ MeV and beta-functions $\beta_x = 6$ m, $\beta_y = 4$ m [62]. The current $I = 300$ A was calculated by the authors of [62] for the bunch charge 0.23 nC and the bunch length $\tau_e \approx 0.7$ ps (we get 330 A though). There is a great deal of uncertainty with regard to the values of the energy spread and the emittance; the energy spread per slice is not given, the projected value, $\sigma_e^{projected} = 0.6\%$, is well too high as compared with other installations, such as LCLS [19] etc.; this lack of definite data for experiments also includes the emittance $\epsilon^n_{x,y}$: the reported data vary between 0.5–3 mm \times mrad [62]. We suppose that most of it is in the projection that is due to transverse centroid shifts along the bunch and the time-sliced values after the injector are well preserved. Reassuming [62], we adopt the data simulation in Table 1, which yields the FEL power evolution, demonstrated in Figure 6.

Table 1. Some simulation data for SACLA FEL experiment at $\lambda = 12.4$ nm at SPRING-8.

Beam parameters:					
relativistic factor $\gamma = 1526$, beam power $P_E = 234$ GW, current $I_0 = 300$ A, current density $J = 2.9 \times 10^{10}$ A/m ² , beam section $\Sigma = 1.03 \times 10^{-8}$ m ² , emittances $\gamma \epsilon_{x,y} \approx 0.5$ μ m, $\beta_x = 6$ m, $\beta_y = 4$ m, beam size $\sigma_{x,y} \approx 40$ μ m, divergence $\theta_{div} \approx 8$ μ rad, $\theta = \sigma_{photon}/L_{gain} \sim 40$ μ rad, $\gamma\theta \approx 0.06$, energy spread (per slice) $\sigma_e = 1.6 \times 10^{-3}$					
Undulator parameters: $k = 2.1$, $\lambda_u = 1.8$ cm, $N = 259$, section length 4.66 m					
Calculated FEL properties: saturated length $L_s = 13$ m, gain length $L_{gain} = 1.1$ m, radiation beam size $\sigma_{photon} \approx \sqrt{\sigma_{x,y} \sqrt{\lambda_1 L_g / 4\pi}} \approx 36$ μ m					
Harmonic number	$n = 1$	$n = 2$	$n = 3$	$n = 4$	$n = 5$
Bessel coefficient f_n	0.79	0.09	0.32	0.09	0.18
Pierce parameter $\tilde{\rho}_n$	0.0015	0.0003	0.0008	0.0003	0.0006
Harmonic wavelength λ_n , nm	12.4	6.2	4.1	3.1	2.5
Saturated power $P_{E,n}$, W	1.9×10^8	—	6×10^5	—	3×10^4

The saturation occurs at the end of the final third undulator section of 4.5 m; we show the simulated FEL harmonic power, the measured energy of the fundamental tone after the third undulator in terms of power and the contribution of the third FEL harmonic, estimated at $\sim 0.3\%$ of the fundamental [62]. We have obtained the simulated value of the saturated fundamental FEL power (red line in Figure 6) $P_1 \approx 0.2$ GW, coinciding with the value $P_{max} = E_\gamma / \tau_{rad} = 0.2$ GW, for the measured fundamental energy [62] $E_\gamma \approx 0.1$ mJ $\pm 13\%$ for the FEL radiation pulse duration $\tau_{rad} = \tau_e \sqrt{2\pi L_g / L_s} = 0.5$ ps, emitted from the electron bunch with the root mean square r.m.s. length $\tau_e = 0.7$ ps. For the third FEL harmonic (see green dashed line in Figure 6) we get $\approx 0.3\%$ power of the fundamental in agreement with [62]. Note that the second harmonic level is very low. The simulated gain length is $L_g \approx 1$ m and the saturated length is $L_s \approx 13$ m; other data are collected in Table 1.

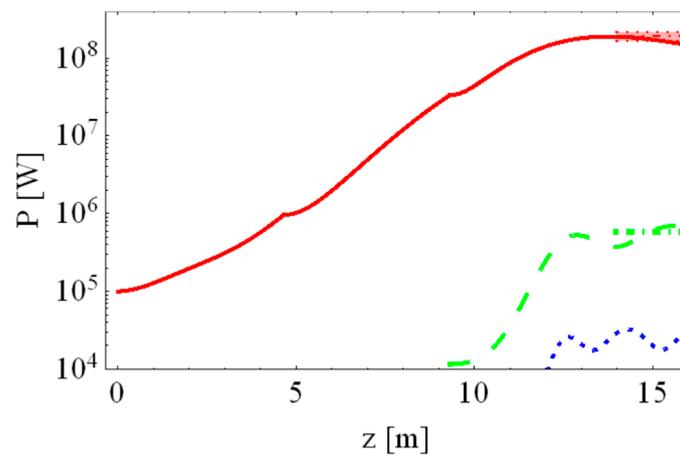


Figure 6. Evolution of the harmonic power in the SACLA experiment for $E = 780$ MeV, $\lambda_1 = 12.4$ nm, $\sigma_e = 1.6 \times 10^{-3}$, $I_0 = 300$ A. The harmonics are color coded: $n = 1$ —red solid, $n = 3$ —green dashed, $n = 5$ —blue dotted. The experimental values of the harmonic powers are denoted by the colored dot-dashed lines on the right.

The spectrum line contains only few subharmonics with $p = -1, 0, +1$; their contribution is shown in Figure 7 for off the axis distance.

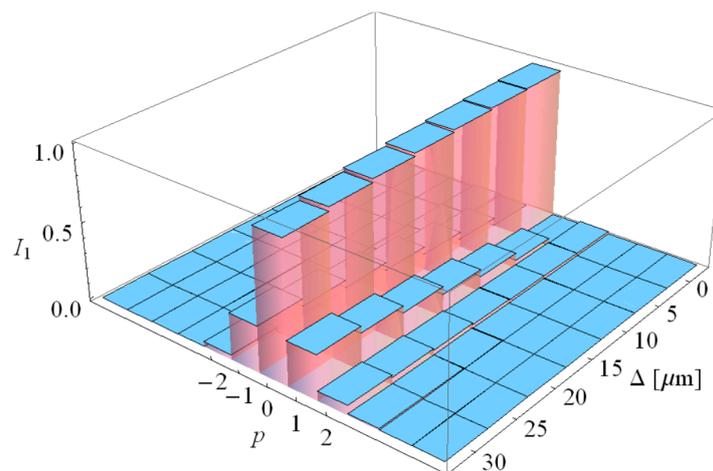


Figure 7. Split of the fundamental radiation line $\lambda = 12.4$ nm for SACLA experiment, as a function of the distance Δ from the electron beam axis.

The main contribution evidently comes from three subharmonics $p = [-1, 0, 1]$; the respective radiation line width is $\Delta\lambda/\lambda \sim 2 \times 10^{-3}$, comparable with the natural UR line width $1/2N \approx 2 \times 10^{-3}$. Theoretical estimation of the relative radiation line bandwidth in SASE FEL after the gain-narrowing in the exponential growth yields similar value $\Delta\lambda/\lambda \approx \sqrt{\rho\lambda_u/L_s} \approx 0.15\%$, close to the FEL scaling parameter $\rho \approx 0.0016$ (see Table 1). Superposition of the randomly distributed over the length of the electron bunch wave trains with the coherence length $l_c = \lambda^2/\Delta\lambda \sim 6$ μm gives the coherence time $t_c = \lambda^2/(c\Delta\lambda) \sim 0.02$ ps. The number of the coherence regions in the radiation pulse is therefore $\tau_{rad}/t_c \sim 20$.

We also modeled SACLA FEL radiation in a hard X-ray band. The results for the recent installation setup [64] are shown in Figures 8 and 9. The electron energy was 10 times higher than in the soft X-ray experiment, and the radiation at the wavelength $\lambda = 0.124$ nm was generated. Some data for the simulations of SACLA FEL in hard X-ray region are given in Table 2.

Table 2. Some simulation data for SACLA FEL experiment at $\lambda = 0.124$ nm at SACLA.

Beam parameters:					
relativistic factor $\gamma = 15264$, beam power $P_E = 78$ TW, current $I_0 = 10$ kA, current density $J = 3.04 \times 10^{12}$ A/m ² , beam section $\Sigma = 3.29 \times 10^{-9}$ m ² , emittances $\gamma \varepsilon_{x,y} \approx 0.4 \mu\text{m}$, $\beta_{x,y} = 20\text{m}$, beam size $\sigma_{x,y} \approx 22 \mu\text{m}$, divergence $\theta_{div} \approx 1.1 \mu\text{rad}$, $\theta = \sigma_{photon}/L_{gain} \sim 9 \mu\text{rad}$, $\gamma\theta \approx 0.14$, energy spread $\sigma_e = 0.926 \times 10^{-3}$					
Undulator parameters: $k = 2.1$, $\lambda_u = 1.8$ cm, $N = 277$, section length 4.66 m					
Calculated FEL properties: saturated length $L_s = 38$ m, gain length $L_{gain} = 2.6$ m, radiation beam size $\sigma_{photon} \approx \sqrt{\sigma_{x,y} \sqrt{\lambda_1 L_g / 4\pi}} \cong 11 \mu\text{m}$					
Harmonic number	$n = 1$	$n = 2$	$n = 3$	$n = 4$	$n = 5$
Bessel coefficient f_n	0.79	0.19	0.27	0.19	0.11
Pierce parameter $\tilde{\rho}_n$	0.00075	0.0003	0.00037	0.0003	0.0002
Harmonic wavelength λ_n , nm	12.4	6.2	4.1	3.1	2.5
Saturated power $P_{F,n}$, W	1.9×10^{10}	9×10^6	5×10^7	5×10^6	1.6×10^5

Omitting the details of the experiments and installation, which are described in [63–68], we note only that the SACLA facility has been continuously upgraded and the data on specific FEL experiments are incomplete (except for the early experiment [61]). For example, the electron beam energy spread for later SACLA setup and hard X-ray experiments is not mentioned in major papers [64–68]; we assumed for the simulation $\sigma_e = 0.0926\%$, following [63], where the upgraded RF system of SPRING 8 was described. Considering that the beam was alternatively sourced to BL3 and BL2 undulator lines, and the above spread was reported for $E = 6$ GeV, we assume that the spread should not increase in the experiment with the energy $E = 7.8$ GeV on BL3 line and 10 keV photons. Of course, it depends on the spreader, the optics, on whether the dispersion was closed etc., thus, the experimental conditions can be different; however, in the absence of explicitly reported data, we have to assume the first approximation of the only available data from [63,65]. Our simulation results are collected in Table 2; the computed saturated power is compared with that obtained from the measured in this experiment photon energy, $E_\gamma = 0.4$ – 0.5 mJ, reported in [64] (see Figures 2c and 3 in [64]).

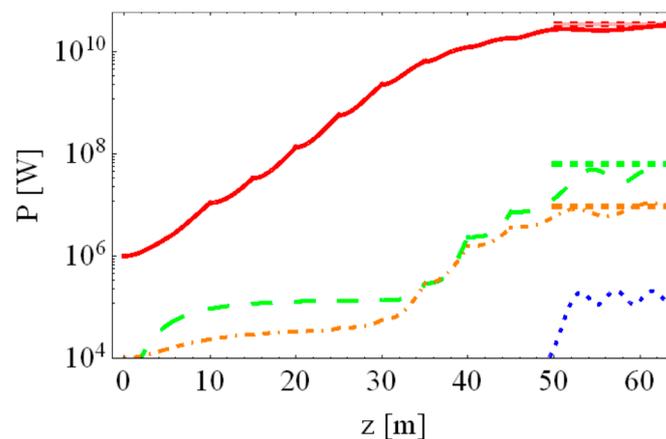


Figure 8. Evolution of the harmonic power in the SACLA FEL experiment for $E = 7800$ MeV, $\lambda_1 = 0.124$ nm, $\sigma_e = 9.26 \times 10^{-4}$, $I_0 = 10$ kA. The harmonics are color coded: $n = 1$ —red solid, $n = 2$ —orange dot-dashed, $n = 3$ —green dashed, $n = 5$ —blue dotted. The experimental values of the harmonic powers are denoted by the colored dotted lines on the right.

With regards to the bunch length $\tau_e = 20$ fs and charge $Q = 0.2$ nC, the current $I = 10$ kA, and other data [64], we get the photon pulse duration $\tau_\gamma \approx 13$ fs, and the saturated powers of the fundamental and

third harmonics as shown by the dashed lines after 50 m in Figure 8; they agree with our theoretical simulations. Horizontal dashed green and orange lines in the saturation region in Figure 8 trace the values 0.2% for the third and 0.03% for the second harmonics. Variation of the emittance, $\pm 1 \mu\text{m}$, and of energy spread, 0.08–0.1%, influences the gain and the saturation lengths and the third harmonic power; the fundamental power is less sensitive to it.

The radiation spectrum line is split in many subharmonics only at the extremities of the beam (see Figures 9 and 10). The account for the subharmonics $p = -5 \dots +5$ is sufficient everywhere, but for the maximum angles of electron–photon interaction at the beam edges (see Figure 10). Accounting for the split of the spectrum line in 11 subharmonics, $p = -5 \dots +5$, we get the total contribution of the latter, $\sqrt{\sum_{p=-5}^{+5} J_p^2}$, after averaging across the beam for all electron-photon interaction angles, close to unity: 0.97. Accounting for $p = -6 \dots +6$ or more subharmonics yields even more precise results. The respective spectral width is 0.06–0.1%, the Pierce parameter $\rho \approx 0.07\%$ and the coherence time $\tau_c \sim 0.7$ fs, which means that less than 20 coherence regions are in $\tau_\gamma \approx 13$ fs photon pulse.

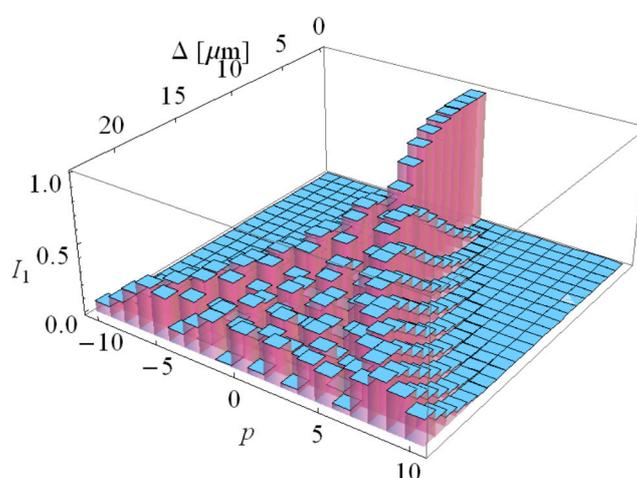


Figure 9. Split of the fundamental spectrum line $\lambda = 0.124$ nm for SACLA BL3 line, as a function of the distance Δ from the electron beam axis.

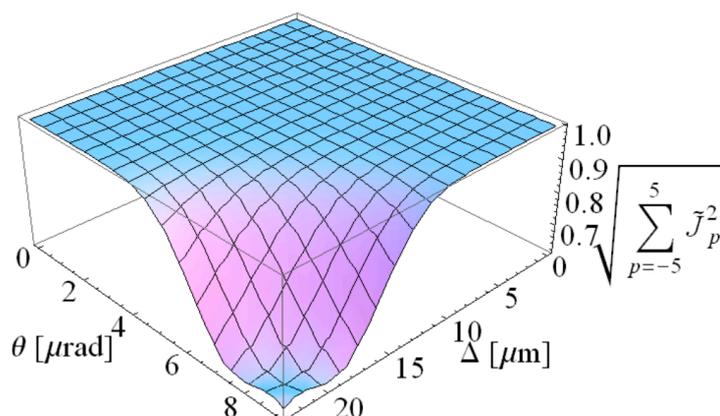


Figure 10. Total contribution of 11 subharmonics with $p = -5 \dots +5$, factorizing the Bessel factors f_n as a function of the distance Δ from the electron beam axis and electron-photon interaction angle θ .

More data are available for similar experiment [61], conducted earlier at SACLA with the electron energy 7 GeV and the undulator with $k = 1.8$. The harmonic power evolution was clearly traced along the undulators and the harmonic saturated powers were measured. Omitting the details, we provide in Figure 11 the comparison between our analytical results and the measured data as reported in [61]. The third harmonic content was $\sim 0.3\%$ of the fundamental. The saturation began after 45 m and was

obvious after ~50 m (see Figure 3 in [61]). The energy spread and emittance influence the gain and saturation lengths. Genuine simulations in [61] agreed fairly well with the experiment: the discrepancy in the harmonic powers at 25–55 m reached one order of magnitude, dependently on the assumed values for the simulation. We computed the saturation beginning at ~45 m, but the process of the saturation seems very gradual; we get full saturation at 55 m. Our analytical results arguably have an even better match with the experiment than the simulations of the authors in [61] and we reproduced the saturated power oscillations as shown in Figure 11. The third harmonic content also fits the measured range ~0.3% (see Figure 11).

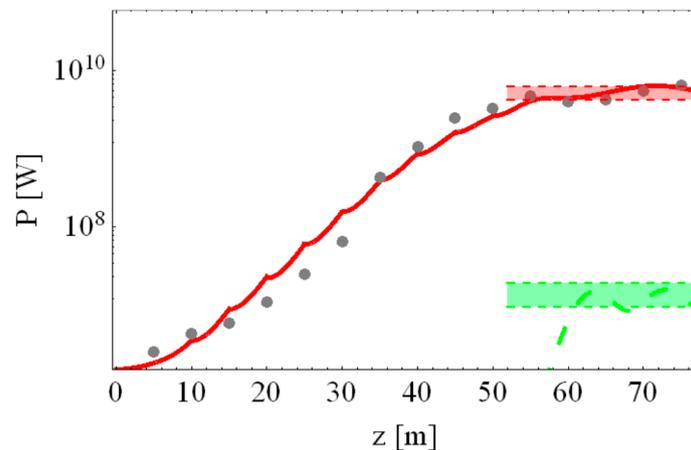


Figure 11. Evolution of the harmonic power in the SACLA experiment for $E = 7$ GeV, $\lambda_1 = 0.124$ nm, $\sigma_e = 8.7 \times 10^{-4}$, $I_0 = 3.5$ kA, $\beta_{x,y} = 22$ m, $\epsilon^n = 0.6\pi$ mm \times mrad. The harmonics are color coded: $n = 1$ —red solid, $n = 3$ —green dashed. The experimental values of the harmonic powers are denoted by the dots and by colored areas on the right.

3.2. POHANG FEL X-ray Experiments

FEL experiments for soft and hard X-ray radiation were conducted at PAL-XFEL facility [39]; the fundamental wavelengths at $\lambda = 1.52$ nm and $\lambda = 0.144$ nm were generated. The experiments were well documented; among other data in [39] the harmonic power evolution was reported. We have analyzed the harmonic generation in both soft and hard X-ray experiments. Some modeling data are collected in Tables 3 and 4. The results are presented in Figures 12–14 and discussed below. The PAL-XFEL resembles in many aspects the LCLS FEL [19]. There is difference in higher energy spread in PAL-XFEL LINAC, also the undulator parameter $k = 2$ in the PAL-XFEL experiments was lower than $k = 3.5$ at LCLS; the electron beam had lower energy for the same generated wavelength. The soft X-ray radiation at $\lambda_1 = 1.52$ nm was produced by the electrons with the energy $E = 3$ GeV and the energy spread $\sigma_e^{soft} = 0.05\%$ (~five times higher than in LCLS), in the undulators with the total pure length ~40 m; the undulator parameter was $k = 2$. The hard X-ray radiation at $\lambda_1 = 0.144$ nm was generated by the electrons with the energy $E = 8$ GeV (vs. $E \sim 13$ GeV in LCLS) with the energy spread $\sigma_e^{hard} = 0.018\%$ (~two times higher than in LCLS) in the undulators with the deflection parameter $k = 1.87$ (vs. $k = 3.5$ in LCLS) of the total pure length 100 m. The undulator sections were 5 m long.

Table 3. Some simulation data for PAL-FEL experiment for soft X-rays, $\lambda = 1.52$ nm, $E = 3$ GeV.

Beam parameters:					
$\gamma = 5870$, beam power $P_E = 6.60$ TW, current $I_0 = 2.2$ kA, current density $J = 1.246 \times 10^{11}$ A/m ² , beam section $\Sigma = 1.766 \times 10^{-8}$ m ² , emittances $\gamma\epsilon_{x,y} = 0.55$ μ m, $\beta = 30$ m, beam size $\sigma_{x,y} = 53$ μ m, divergence ≈ 1.8 μ rad, $\theta = \sigma_{\text{photon}}/L_{\text{gain}} \approx 15$ μ rad, energy spread $\sigma_e = 0.5 \times 10^{-3}$					
Undulator parameters: $k = 2$, $\lambda_u = 3.5$ cm, section length 5 m					
Calculated FEL properties: saturated length $L_s = 31$ m, gain length $L_{\text{gain}} = 2.0$ m, radiation beam size $\sigma_{\text{photon}} \approx \sqrt{\sigma_{x,y} \sqrt{\lambda_1 L_g / 4\pi}} \approx 0.29$ mm					
Harmonic number	$n = 1$	$n = 2$	$n = 3$	$n = 4$	$n = 5$
Bessel coefficient f_n	0.80	0.13	0.32	0.13	0.16
Pierce parameter $\tilde{\rho}_n$	0.0010	0.0003	0.0005	0.0003	0.0003
Harmonic wavelength λ_n , nm	1.52	0.76	0.51	0.38	0.30
Saturated power $P_{F,n}$, W	8.2×10^9	3.2×10^6	5.4×10^7	1.6×10^6	2.0×10^6

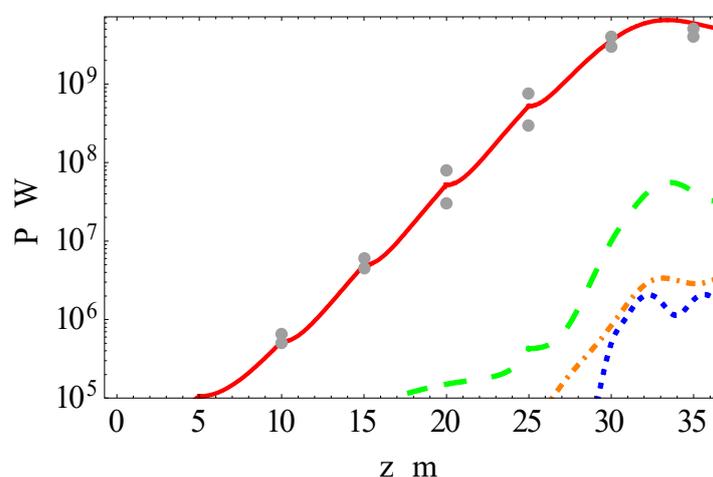


Figure 12. The harmonic power evolution along the undulators at PAL-XFEL for soft X-rays, $\lambda_1 = 1.52$ nm. The experimental average values are shown by dots, following the data in [39]. The harmonics are color coded: $n = 1$ —red solid, $n = 2$ —orange dot-dashed, $n = 3$ —green dashed, $n = 5$ —blue dotted.

The UR in long undulators can be distorted due to non-periodic magnetic fields. The relevant study is presented in the previous section. Using the data from the experimental setup [39], we analytically obtained the evolution of the FEL power for the harmonics, as shown in Figure 12 for soft and Figure 13 for hard X-rays accounting for the beam size, divergences and other data. The results are compared with the measurements of the fundamental harmonic power in [39]. Our analytical modeling gives good match with the experiments (see Figures 12 and 13). The agreement with the experiment for soft X-rays in the exponential growth is even better than that of the authors of [39]. Our analytical results for hard X-ray experiment agree fairly well with the measurements; however, the agreement is marginally better than that of the three-dimensional (3D) numerical simulations in [39]. This evidences the correct analytical account for all underlying physical phenomena.

Table 4. Some simulation data for PAL-XFEL experiment for hard X-rays, $\lambda = 0.144$ nm, $E = 8$ GeV.

Beam parameters:					
$\gamma = 15,660$, beam power $P_E = 20.0$ TW, current $I_0 = 2,5$ kA, current density $J = 3.16 \times 10^{11}$ A/m ² , beam section $\Sigma = 7.91 \times 10^{-9}$ m ² , emittances $\gamma\epsilon_{x,y} \approx 0.55\mu\text{m}$, $\beta \approx 36$ m,					
beam size $\sigma_{x,y} = 35$ μm , divergence ≈ 1 μrad , $\theta = \sigma_{\text{photon}}/L_{\text{gain}} \approx 4.5$ μrad , energy spread $\sigma_e = 0.18 \times 10^{-3}$					
Undulator parameters: $k = 1.87$, $\lambda_u = 2.571$ cm, section length 5 m					
Calculated FEL properties: saturated length $L_s \sim 55$ m, gain length $L_{\text{gain}} = 3.4$ m,					
radiation beam size $\sigma_{\text{photon}} \approx \sqrt{\sigma_{x,y} \sqrt{\lambda_1 L_g / 4\pi}} \approx 15$ μm					
Harmonic number	$n = 1$	$n = 2$	$n = 3$	$n = 4$	$n = 5$
Bessel coefficient f_n	0.82	0.09	0.31	0.09	0.16
Pierce parameter $\tilde{\rho}_n$	0.0004	0.00009	0.0002	0.00009	0.00014
Harmonic wavelength λ_n , nm	0.144	0.072	0.048	0.036	0.029
Saturated power $P_{E,n}$, W	1.0×10^{10}	2.0×10^6	1.0×10^8	1.0×10^6	6.0×10^6

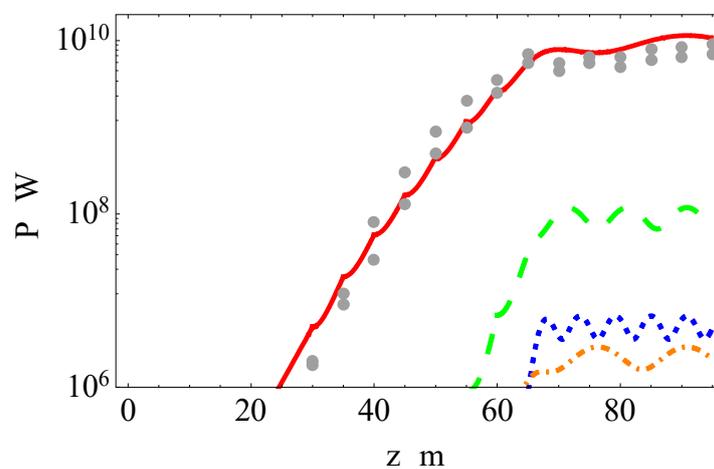


Figure 13. The harmonic power evolution along the undulators at PAL-XFEL for hard X-rays, $\lambda_1 = 0.144$ nm. The experimental average values are shown by dots, following the data in [39]. The harmonics are color coded: $n = 1$ —red solid, $n = 2$ —orange dot-dashed, $n = 3$ —green dashed, $n = 5$ —blue dotted.

No data are available for high harmonic generation in PAL-XFEL experiments. For soft X-rays, the energy spread was higher than in LCLS experiments for similar radiation wavelengths: $\sigma_e^{\text{soft}} = 0.0005 \cong \rho_1/2 \cong \rho_3 \cong 0.0005 > \rho_5 \cong 0.0003$. For the emittance we get pure value $\epsilon \cong 0.94 \times 10^{-10}$ m to be compared with $\lambda_1/4\pi = 1.2 \times 10^{-10}$ m. However, for the fifth harmonic we get $\lambda_5/4\pi = 2.5 \times 10^{-11}$ m and $\epsilon \cong \lambda_5/\pi = 1 \times 10^{-10}$ m. The third harmonic could appear with the power rate $\sim 0.7\%$ of the fundamental, the second harmonic would have the power rate $\sim 0.05\%$, as shown in Figures 12 and 13. Our estimation for the third harmonic at PAL-XFEL in soft X-rays, $P_3/P_1 \sim 0.7\%$, is roughly a half of that for a similar LCLS experiment, where $P_3/P_1 \sim 1.3\%$ for $\lambda_3 = 0.5$ nm [20] with similar radiation parameters. Thus, we can expect some weaker third harmonic at PAL-XFEL due to the smaller value of the undulator parameter k as compared with LCLS; moreover, the detrimental effect of the energy spread is higher for PAL-XFEL, $\sigma_e = 0.0002 \div 0.0005$, as compared with that in LCLS, where $\sigma_e = 0.0001$. The second hard X-ray harmonic at PAL-XFEL is weak; high High harmonics were not registered in the PAL-XFEL experiments.

For hard X-rays the energy spread σ_e and Pierce parameters ρ_n are as follows: $\sigma_e^{\text{hard}} = 0.00018 \cong \rho_1/2 < \rho_3 = 0.00021 > \rho_5 = 0.00015, \rho_2 = 0.00006$; the relations between σ_e and ρ_n are close to those for soft X-ray radiation. Moreover, for hard X-rays the comparison of the pure emittance $\epsilon \cong 3.5 \times 10^{-11}$ m

with $\lambda_3/4\pi = 3.8 \times 10^{-12}$ m is not favorable for the third harmonic radiation: $\varepsilon_{x,y} \cong 10 \times \lambda_3/(4\pi)$. Unsurprisingly, high hard X-rays harmonics were not detected. However, if the energy spread and emittances are improved, then we can expect at the PAL-XFEL high harmonic generation as suggested in Figures 12 and 13. The off-axis deviation of the beam in PAL-XFEL amounted to ~ 10 μm on one undulator length [39]. This causes the off-axis angle ~ 2 μrad , comparable with the divergence, ~ 2 μrad for soft X-rays and ~ 1 μrad for hard X-rays. In the soft X-ray experiment the deviation of the beam in few undulator segments reached 20 μm on one undulator length [39]; this induces the angle 4 μrad . However, the electron-photon interaction on one gain length must be considered with the angle $\bar{\theta} \cong 14$ μrad , far exceeding the beam deviation. The latter angle causes the 2nd FEL harmonic, whose power is estimated $\sim 10^{-4}$ of the fundamental (see Figure 12). For the hard X-ray experiment we get much smaller value $\bar{\theta} \cong 4$ μrad , and the beam must be kept on the axis more precisely.

The proposed analytical approach allows theoretical study of the spectral line split and width in the PAL-XFEL experiments. Following the developed in Section 2 theory, we compute the split of the soft X-ray spectral line and obtain the main contribution from the subharmonics with $p = -4, \dots, +4$; higher subharmonics are negligible. Thus the fundamental tone at $\lambda_1 = 1.5$ nm is split in ~ 9 subharmonics. The total width of the line is $\Delta\lambda \sim 2.3$ pm, the relative value is $\Delta\lambda/\lambda \sim 1.5 \times 10^{-3}$. It is small, but it is higher than the respective value in the soft X-ray LCLS experiment, where $\Delta\lambda/\lambda \sim 5 \times 10^{-4}$. For hard X-rays in PAL-XFEL we have to account for more subharmonics: $p = -7, \dots, +7$. Nevertheless, the line remains rather narrow even with account for this split; the subharmonics are close to each other because of the electrons are ultrarelativistic. The absolute width of the hard X-ray line is $\Delta\lambda \sim 0.14$ pm and the relative width is $\Delta\lambda/\lambda \approx 1.0 \times 10^{-3}$. Compared with the respective values in the LCLS experiment, $\Delta\lambda/\lambda \approx 3 \times 10^{-5}$, the spectrum lines for PAL-XFEL radiation appear wider by ~ 1 – 2 orders of magnitude.

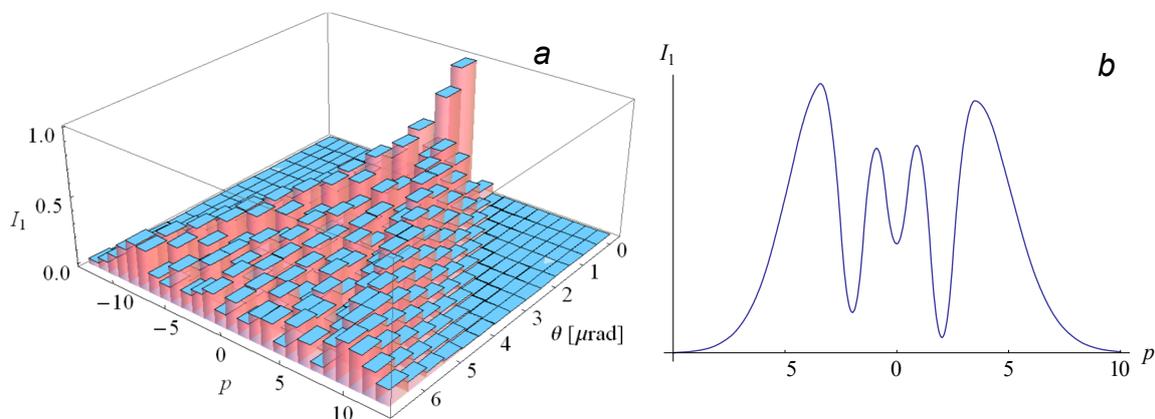


Figure 14. Split of the fundamental spectrum line $\lambda = 0.144$ nm for PAL-XFEL experiment.

The spectral width of the radiation depends on the position of the electrons in the beam. The split of the spectrum line for the hard X-ray radiation from the electrons at the outer extremity of the beam in PAL-XFEL experiment is demonstrated in Figure 14a. Observe that we must account for ~ 20 subharmonics for the radiation from the edges of the beam, while the theoretical line shape shown in Figure 14b.

We have modeled in a similar way other FEL experiments at other installations; in all the cases the results matched well with the measured data, and the analysis given above worked.

4. Conclusions

We have presented analytical formulation of the harmonic generation in FELs with multiperiodic magnetic fields accounting for the harmonic and constant field components and off-axis effects in undulators. We have obtained exact analytical expressions for the Bessel coefficients in the general

case of the multiperiodic elliptic undulator; they account for constant magnetic constituents, finite beam size and off-axis angles, and describe harmonic generation in wide electron beams and in high precision undulators, where fine alignment of narrow beams is required.

The Bessel coefficients for the general elliptic undulator and in its limiting cases of the elliptic and planar undulators with harmonics are provided. The effect of the constant non-periodic magnetic field on the UR line shape is formulated in terms of the generalized Airy function. The corrections for the Bessel coefficients, accounting for the constant magnetic components, off-axis radiation and beam position are given in the analytical integral form of generalized Bessel and Airy functions. The analysis shows that the relevant effects matter for the field-induced and off-axis angles $\gamma\theta > 0.05$. Exact analytical formulae for quantitative calculations of the UR are given in Section 2. Due to the betatron oscillations the UR lines are split in subharmonics; the split is very fine and for relativistic beams the subharmonics are very close to each other: $\delta\lambda/\lambda \sim 1/\gamma \ll 1$. Despite that, it causes noticeable broadening of the spectrum lines. The contribution of the betatron oscillations to the even harmonic generation is one–two orders of magnitude less than that caused by the off-axis and photon–electron interaction angles in real beams.

We have demonstrated theoretical spectrum lines of UR harmonics, their shapes and intensities with the help of the developed theoretical tools. The effect of the non-periodic field and off-axis effects in finite sized beams are clearly distinguished and elucidated in Figures 1–5.

The obtained rigorous theoretical results are employed for FEL radiation studies with the help of the phenomenological FEL model. We have analyzed the PAL-XFEL experiment at POHANG laboratory [60], where soft and hard X-rays were produced. Our analytical results are in good agreement with the reported values (see Figures 12 and 13). We have modeled possible FEL harmonic behaviors accounting for the beam sizes, divergences, electron-photon interaction angles, energy spread and diffraction; the second harmonic would be very weak, in particular, for hard X-rays. The spectral line in hard X-ray experiment is split in ~ 15 subharmonics; despite that, we get quite narrow line, $\Delta\lambda/\lambda \sim 1.0 \times 10^{-3}$, $\Delta\lambda \sim 0.14$ pm due to $\gamma \sim 1500 \gg 1$. However, this is >10 times wider than in the LCLS experiment at the same wavelength, where $\Delta\lambda/\lambda \approx 3 \times 10^{-5}$. The radiation of high harmonics at PAL-XFEL is limited by a rather high energy spread: for hard X-rays $\sigma_e^{hard} = 0.00018$, the Pierce parameters for the $n = 1, 2, 3, 5$ harmonics are $\rho_1 \cong 0.0004$, $\rho_3 \cong 0.0002$, $\rho_5 \cong 0.00015$, $\rho_2 \cong 0.00006$; for soft X-rays the energy spread is $\sigma_e^{soft} = 0.0005$, and the Pierce parameters are $\rho_1 \cong 0.0010$, $\rho_3 \cong 0.0006$, $\rho_5 \cong 0.0004$, $\rho_2 \cong 0.00007$. We have demonstrated possible theoretical harmonic radiation at PAL-XFEL; however, due to relatively high energy spread, we can hardly expect radiation of the harmonics higher than the third.

The analysis of the SACLA facility reveals the spectrum and power evolution for the undulator line BL3 with two different designs and $k = 1.8$ and $k = 2.1$ with respective electron energies 7 GeV and 7.8 GeV. The modeling for the 7 GeV agrees with the experiment; the modeling for 7.8 GeV gives the prediction of the harmonic evolution in the absence of the measured data. The saturated harmonic powers in all cases agree with those measured. The theoretical spectrum line split is demonstrated for the soft and hard X-rays for the lines BL1 and BL3; the radiation line at 0.124 nm is split is $>\pm 5$ subharmonic, the spectrum line $\lambda = 12.4$ nm at BL1 is split in three subharmonics. The spectrum lines at SACLA are narrower than at PAL-XFEL.

The results, obtained with the help of the developed theoretical formalism for FEL power and spectrum evaluation agree with the experiments in X-ray and other bands. The analytical formulae are relatively simple and the relevant calculations do not require special knowledge and programmer skills; they can be done on any PC. The predictions are accurate and agree with the measurements. This allows the theoretical study of current and planned FEL experiments and estimation of performance, spectrum and harmonic generation in operating and constructed FELs.

Funding: This research received no external funding.

Acknowledgments: The authors are grateful to Professor A. Borisov and Leading Researcher A. Lobanov for useful discussions.

Conflicts of Interest: The authors declare no conflict of interest.

Appendix A. Phenomenological Model of Harmonic Power Evolution in High-Gain FELs

The Pierce parameter ρ_n reads accounting for the diffraction as follows [3–6,23,45–53]:

$$\rho_n = \frac{J^{1/3}(\lambda_u k_{eff} |f_n|)^{2/3}}{2\gamma(4\pi i)^{1/3}} \chi, \quad \chi = 1 / \left(1 + \frac{\lambda_u \lambda_n}{16\pi \rho_n \Sigma} \right)^{1/3}, \quad (A1)$$

where n is the harmonic number, $J = I_0/\Sigma$ [A/m²] is the current density, $\Sigma = 2\pi\sigma_x\sigma_y$ is the beam section, $\sigma_{x,y} = \sqrt{\varepsilon_{x,y}\beta_{x,y}}$ are the sizes of the beam, $\varepsilon_{x,y} = \sigma_{x,y}\theta_{x,y}$ are the emittances, $\beta_{x,y} = \varepsilon_{x,y}/\theta_{x,y}^2$ are the betatron average values, $\theta_{x,y}$ are the divergences, $i \cong 1.7045 \times 10^4$ is the constant of Alfvén current [A], $k_{eff} = k\sqrt{\omega}$ (see (3) for ω) is the effective undulator parameter, which reduces for the common planar undulator to $k = \frac{eH_0\lambda_u}{2\pi mc^2} \approx 0.934H_0\lambda_u$ [T · cm], H_0 is the magnetic field amplitude on the undulator axis, f_n is the Bessel factor for the n -th UR harmonic. The Bessel factors f_n in the general case of the two-dimensional field with harmonics (2) are given by (22)–(24) and (31) accounting for the finite beam size effects and constant magnetic components, which cause even harmonics. We assume the fundamental harmonic is not suppressed and it dominates. The saturated n -th harmonic power can be calculated accounting for the loss factors following [53]: $P_{n,F} = \sqrt{2}P_e\eta_n\eta_1\chi^2\rho_1f_n^2/(n^{5/2}f_1^2)$, where $P_e = I_0E$ is the beam power, I_0 is the beam current [A], E is the electron energy [eV], ρ_1 is the Pierce parameter. The gain length for the n -th harmonic is $L_{n,g} \cong \Phi_n\lambda_u/(4\pi\sqrt{3}n^{1/3}\chi\rho_n)$, where λ_u is the undulator period, Φ_n η_n are the loss factors. For the fundamental tone we denote $L_{1,g} \equiv L_g$; the fundamental tone saturation length is $L_s \cong 1.07L_{1,g} \ln(9P_{1,F}/P_{1,0})$.

The correction factors phenomenologically describe major losses as follows:

$$\tilde{\Phi}_n = \Phi_n|_{\mu_{e,n} \rightarrow \tilde{\mu}_{e,n}}, \quad \Phi_n \cong (\zeta^n + 0.165\mu_{e,n}^2) \exp(0.034\mu_{e,n}^2), \quad (A2)$$

$$\tilde{\mu}_{e,n}(\sigma_e, n) \cong 2n^{2/3}\sigma_e/(\chi\rho_n), \quad \mu_{e,n}(\sigma_e, n) \cong 2\sigma_e/(n^{1/3}\chi\rho_n), \quad (A3)$$

$$\tilde{\eta}_n = \eta_n|_{\Phi_n \rightarrow \tilde{\Phi}_n}, \quad \eta_n \cong (e^{-\Phi_n(\Phi_n-0.9)} + 1.57(\Phi_n - 0.9)/\Phi_n^3)/1.062. \quad (A4)$$

The coefficient ζ in (A2) is calculated by the cumbersome formula, involving the betatron parameters (see [24,53]); in matched beam, $\zeta \approx 1 - 1.05$, and this correction is small. In X-ray FELs, it is even smaller: $\zeta \leq 1.02$ and often $\zeta = 1$. The beam diffraction lowers the Pierce parameter value ρ_n ; the beam energy spread σ_e and the emittances $\varepsilon_{x,y}$ prolong the gain $L_{n,g}$ and reduce the saturated powers $P_{n,F}$ for harmonics. For stable FEL amplification, weak conditions $\sigma_e \circ \rho_n/2$, $\varepsilon_{x,y} \circ \lambda_n/4\pi$ should be fulfilled (see, for example, [3–6,23,45–50]); however, failure to satisfy them exactly, especially in X-ray band, does not mean these harmonics will not be radiated at all.

For the harmonic power evolution in the initially unbunched electron beam, we use formula [53]:

$$P_{L,n}(z) \cong \frac{P_{0,n}A(n,z)e^{0.223z/L_s}}{1 + (A(n,z) - 1)\frac{P_{0,n}}{P_{n,F}}}, \quad A(n,z) \cong \frac{1}{3} + \frac{\cosh \frac{z}{L_{n,g}}}{4.5} + \frac{\cos \frac{\sqrt{3}z}{2L_{n,g}} \cosh \frac{z}{2L_{n,g}}}{0.444}, \quad (A5)$$

where we ad-hoc introduce $\bar{P}_{n,f} = P_{n,f}(1 + 0.3\cos(n(z - L_s)/1.3L_g))/1.3$ to describe the saturated FEL power oscillations. The match with FEL experiments appears good, as seen in Figures 11–13. In cascaded FELs, the previous cascade feeds the next cascade with the prebunched beam. Even in the case the radiation on the n -th harmonic is suppressed for some reason, the initial power of the n -th UR harmonic can be provided by the bunching, $P_{n,0} \cong d_n b_n^2 P_{n,F}$, which is induced by the dominant harmonic and $d_{n=1,2,3,4,5} \approx \{1, 3, 8, 40, 120\}$. The fundamental tone induces the bunching [48] $b_n(z) \cong h_n(P_1(z)/P_e\rho_1)^{n/2}$, where $h_{1,2,3,4,5} \cong \{1, 1.5, 2.4, 4.3, 7.7\}$. If the fundamental

tone is suppressed, the dominant harmonic induces its sub-harmonics (see [15,16]). The independent harmonic power evolution in the cascade is described by the following formula [53]:

$$P_{L,n}(z) \cong \frac{P_{0,n}F(n,z)}{1 + F(n,z)\frac{P_{0,n}}{\bar{P}_{n,F}}}, \quad F(n,z) \cong 2 \left| \cosh \frac{z}{L_{n,g}} - \cos \frac{z}{2L_{n,g}} \cosh \frac{z}{2L_{n,g}} \right|. \quad (\text{A6})$$

We have reevaluated the contribution of the initial shot noise P_{noise} in the self-amplified spontaneous emission (SASE) FEL with respect to all earlier works; fitting with available measurements from many FELs on average yields:

$$N_n(z) \cong \frac{P_{\text{noise}}}{9n} \frac{S_n(z)}{1 + 30P_{\text{noise}}S_n(z)/nP_{n,f}}, \quad (\text{A7})$$

$$S_n(z) \cong 2 \left| \cosh \frac{z}{L_{n,g}} - e^{-\frac{z}{2L_{n,g}}} \cos\left(\frac{\pi}{3} - \frac{\sqrt{3}z}{2L_{n,g}}\right) - e^{\frac{z}{2L_{n,g}}} \cos\left(\frac{\pi}{3} + \frac{\sqrt{3}z}{2L_{n,g}}\right) \right|.$$

The dominant FEL harmonic (usually the fundamental) generates subharmonics in nonlinear regime: the harmonic powers then grow as the n -th power of the dominant harmonic $\propto \exp(nz/L_g)$ [23, 45–50]. The electron-photon interaction at high harmonic wavelengths is more sensitive to losses than that at the fundamental wavelength. Improving the phenomenological description in [12–16] and other earlier works, we now describe gradual harmonic saturation by two terms:

$$Q_n(z) \cong \frac{\tilde{P}_{n,F}}{\left(e^{-nz/L_g}/d_n b_n^2\right) + \left(1 - e^{-nz/L_g}\right)} + \frac{\bar{P}_{n,F}}{\left(e^{-nz/L_g}/b_n^2\right) + \left(1 - e^{-nz/L_g}\right)}, \quad (\text{A8})$$

where the bunching $b_n \cong (P_{0,1}/9P_e\rho_1)^{n/2}$ is induced by the fundamental harmonic with the initial power $P_{0,1}$, the n -th harmonic in nonlinear generation begins to saturate at the power level $\tilde{P}_F = P_F|_{\eta_n \rightarrow \tilde{\eta}_n}$ and saturates with oscillations around the power $\bar{P}_{n,f}$. In an elliptic undulator two polarizations are radiated and the effective Pierce parameter is modified accordingly (see, for example, [16]). The above analytical model of the FEL harmonic power evolution describes independent and induced harmonic contributions, multistage harmonic saturation, power oscillations, all major losses and different sensitivity of the photon-electron interaction at different harmonic wavelengths; it agrees with the available results of FEL experiments in a wide range of conditions and radiated wavelengths.

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