



Article

Invisible Gold in Pyrite from Epithermal, Banded-Iron-Formation-Hosted, and Sedimentary Gold Deposits: Evidence of Hydrothermal Influence

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Abstract: “Invisible gold” in pyrite is defined as an Au solid solution of the pyrite lattice, sub-microscopic Au nanoparticles (NPs) in the pyrite, or other chemisorption complexes of Au. Because the relationship between the Au and As concentrations in pyrite could indicate the genesis of the deposit, the purpose of this study is to assess the micro-analytical characteristics of the Au–As relationship in pyrite from epithermal and hydrothermally affected sedimentary Au deposits by secondary ion mass spectrometry. The Au and As concentrations in pyrite vary from 0.04 to 30 ppm and from 1 to 1000 ppm, respectively, in the high-sulfidation Nansatsu-type epithermal deposits; these concentrations are both lower than those of the low-sulfidation epithermal Hishikari deposit. The Au concentrations in pyrrhotite and pyrite reach 6 and 0.3 ppm, respectively, in the Kalahari Goldridge banded-iron-formation-hosted gold deposit, and Au in pyrrhotite may sometimes exist as NPs, whereas As concentrations in pyrrhotite and pyrite are both low and lie in a narrow range from 6 to 22 ppm. Whether Au is present as NPs is important in ore dressing. The Au and As concentrations in pyrite from the Witwatersrand gold field range from 0.02 to 1.1 ppm and from 8 to 4000 ppm, respectively. The shape of the pyrite grains might prove to be an indicator of the hydrothermal influence on deposits of sedimentary origin, which implies the genesis of the deposits.

Keywords: invisible gold; arsenic; pyrite; epithermal; banded iron formation; sedimentary; hydrothermal; secondary ion mass spectrometry (SIMS)

1. Introduction

It has been observed that gold recoveries from ore in many gold mining operations are generally lower than those determined by assay [1]. This is because Au has been found to occur as electrum or native Au and as “invisible gold” that is too small to see by optical microscope. The “invisible gold” comprises sub-microscopic Au nanoparticles (NPs; <100 nm, [2–4]) in minerals, such as sulfides, Au in a solid solution of sulfide crystals, Au in surficial nonautonomous phases [5], and chemisorbed Au polysulfide complexes [6]. Such Au types are referred to as “invisible gold”. It is of great concern whether the “invisible gold” occurs as NPs or as structurally bound in the lattices of the sulfides because the former could be extracted from finely pulverized concentrate, while the latter results in the ore being refractory. Microbeam analytical techniques have been applied to evaluate the presence of Au in sulfides in Au deposits [7–12]. Morishita and Hammond [13] observed Au peaks corresponding

to NPs in pyrrhotite from a banded-iron-formation (BIF)-hosted Au deposit in South Africa using secondary ion mass spectrometry (SIMS) depth profiling techniques.

Several studies have shown that there is generally a positive correlation between Au and As concentrations in hand specimens obtained from geochemical Au explorations (e.g., [14]). Pyrite is ubiquitous in most hydrothermal deposits, and several papers have reported that arsenian pyrite is a common host for invisible gold (e.g., [8]). The Au content in pyrite generally increases with an increasing As content [15–17].

Reich et al. [18] have shown that the solubility of Au in arsenian pyrite from Carlin-type and epithermal Au deposits increases with the As content in the pyrite. They proposed an empirical solubility limit of Au, which defines the boundary between Au being present as NPs (Au^0) or solid solutions (Au^+) in pyrite. Morishita et al. [19] found a positive correlation between Au and As in the Hishikari pyrite in a microscopic area of 3 μm . Additionally, Morishita et al. [20] demonstrated that there is a linear correlation between Au and As concentrations in pyrite from the three ore zones of the high-grade Hishikari deposit.

The purpose of this study is to present the micro-analytical characteristics of the Au–As relationship in pyrite and pyrrhotite (if applicable) from epithermal, BIF-hosted, and sedimentary Au deposits by SIMS for considering the genesis of these ore deposits.

2. Geological Setting and Mineralogy

2.1. High- vs. Low-Sulfidation Epithermal Deposits

In epithermal systems, deposits with high-sulfidation-state minerals, including enargite, tennantite, and covellite, are commonly referred to as “high-sulfidation” or “acid sulfate” deposits [21,22]. In contrast, “low-sulfidation” deposits (or “adularia–sericite” deposits) of epithermal Au mineralization related to near-neutral pH and relatively reduced fluids are associated with low-sulfidation-state minerals [22].

2.2. Epithermal Deposits in Kyushu, Japan

Kyushu is located at the intersection of the Southwestern Japan Arc and the Ryukyu arc on the eastern margin of the Eurasian plate facing the Philippine Sea Plate. Many epithermal Au deposits of Pliocene to Pleistocene age occur in Late Cenozoic subaerial andesitic to dacitic volcanic rocks in Kyushu, Japan (Figure 1). Kagoshima-ken (a prefecture in Kyushu) is situated in the southernmost part of Kyushu and constitutes two historically fixed districts, the Hokusatsu (Northern Kagoshima) and Nansatsu (Southern Kagoshima) districts (Figure 1). Three high-sulfidation epithermal deposits, the Kasuga, Iwato, and Akeshi deposits, are in the Nansatsu district of Kagoshima-ken and are thus commonly called the Nansatsu-type Au deposits. The low-sulfidation epithermal Hishikari deposit in the Hokusatsu district is about 90 km north of the Nansatsu-type Au deposits (Figure 1).

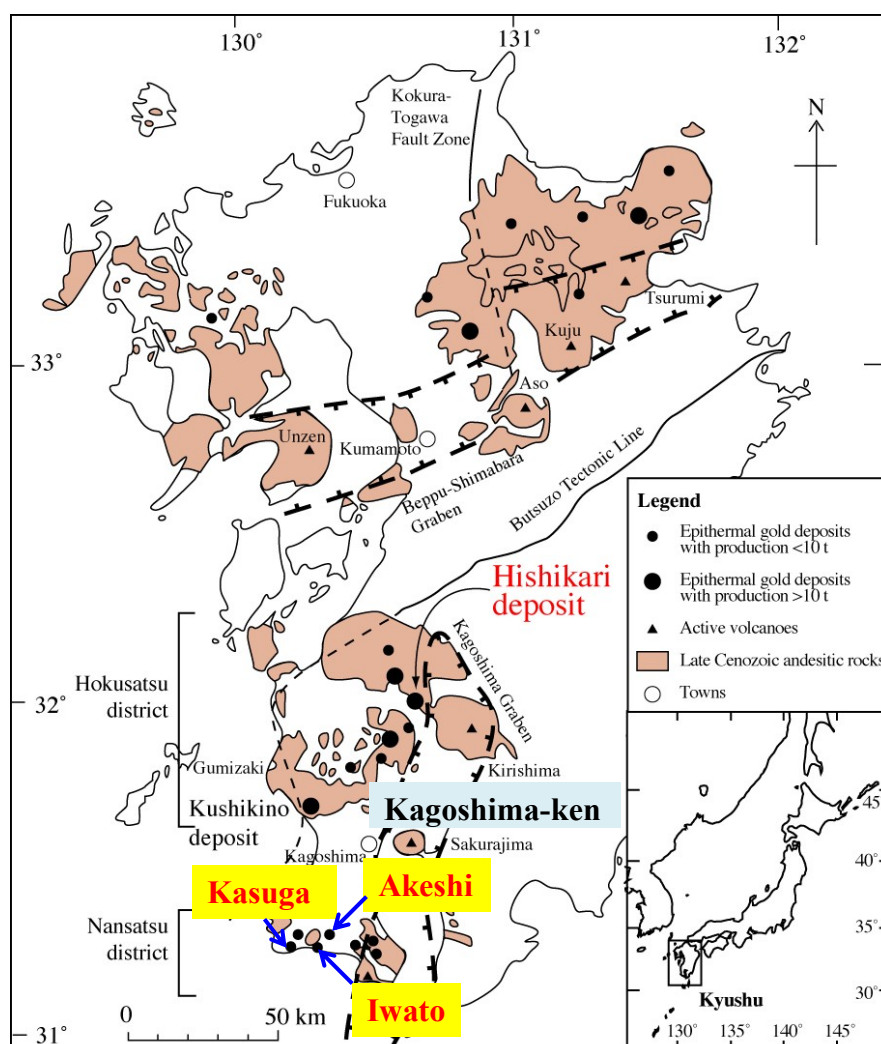


Figure 1. Epithermal gold deposits in Kyushu in Southwestern Japan [23,24]. The location of the Kasuga, Iwato, and Akeshi deposits are shown in addition to the Hishikari deposit.

2.3. High-Sulfidation Nansatsu-Type Au Deposits in the Nansatsu District of Kagoshima-ken

Calc-alkaline volcanism in Kagoshima-ken has moved eastward at an average speed of 0.6 cm/year from 8 Ma to date [25,26]. The eastward shift of the volcanic front may be due to a progressive steepening of the west-dipping subduction slab [26]. As hydrothermal activity has followed the volcanism, the location of Au mineralization has also moved eastward. Among the three Nansatsu-type Au deposits, the westernmost Kasuga deposit formed circa 5 Ma, and the intermediate Iwato and easternmost Akeshi deposits (younger than 4 Ma) have formed in turn in the Nansatsu district of Kagoshima-ken (Figure 1) [26]. The present-day active volcanic front occurs approximately 30 km east of the Nansatsu deposits. The Nansatsu district is underlain by sedimentary rocks of the Cretaceous Shimanto Supergroup (primarily of the Kawanabe Group). Kaolinite and its higher-temperature polymorph, dickite, are common in the Nansatsu deposits [27].

2.3.1. Kasuga Deposit

The Kasuga main ore body consists of a silicified zone extending 500 m E–W and 100 m N–S, with a thickness of approximately 80 m. The silica body is enveloped by an argillic zone that consists of an inner dickite (alunite–pyrophyllite) zone and an outer illite–chlorite zone. The average Au content of the silicified rock ranges from 2 to 3 g/t [28]. The principal ore minerals in the deposit are enargite and

native gold. Pyrite is ubiquitous throughout the ore body. Several quartz veinlet systems are oriented radially with respect to a basement depression in the vicinity of Kasuga [27]. The mineral paragenesis of the Kasuga deposit is given in Table 1.

Table 1. General description of the studied Au deposits.

Type	Sub-Type/Region	Name	Mineral Assemblage
Epithermal	high-sulfidation Nansatsu-type	Kasuga	quartz, enargite, pyrite, native gold, alunite, kaolinite, dickite, covellite, native sulfur, goethite [27,28]
Epithermal	high-sulfidation Nansatsu-type	Iwato	quartz, pyrite, native gold, alunite, kaolinite, dickite, goethite [28]
Epithermal	high-sulfidation Nansatsu-type	Akeshi	quartz, luzonite, pyrite, native gold, covellite, alunite, native sulfur, kaolinite, dickite, smectite, scorodite, goethite [28]
Epithermal	low-sulfidation vein-type	Hishikari	quartz, adularia, smectite, pyrite, marcasite, electrum, naumannite-aguilarite, pyrargyrite, chalcopyrite [29]
BIF-hosted hydrothermal	Kraaipan greenstone belt	Kalahari Goldridge	quartz, siderite, ankerite, calcite, pyrite, pyrrhotite, stilpnomelane, chlorite, magnetite, epidote, plagioclase [30]
Sedimentary & hydrothermal	auriferous palaeoplacer/Witwatersrand basin	South Deep	quartz, pyrite, native gold
Sedimentary & hydrothermal	auriferous palaeoplacer/Witwatersrand basin	KDC West	quartz, pyrite, native gold

BIF: banded iron formation. KDC: Kloof-Driefontein Complex.

2.3.2. Iwato Deposit

The Iwato silicified zone extends more than 1200 m E–W and approximately 200 m N–S, with a thickness of approximately 100 m. Three groups of ore bodies occur within the silicified zone; from east to west, these are the Maruyama ore bodies, the Arabira ore bodies, and the Iwashita ore body. Although the silicified zone continues further to the southwest, minable Au has not been found there. Most of the silicified rocks there contain more than 0.2 g/t Au. The high-grade zone (above 4 g/t) often has a mushroom shape [28]. The mineral paragenesis of the Iwato deposit is given in Table 1.

2.3.3. Akeshi Deposit

A silicified zone (300 m × 200 m, with a thickness of approximately 150 m) contains high-grade zones consisting of gray or milky white quartz. The No. 1 ore body (100 m NE–SW with a width of 40 m) plunges SW at an angle of 60° from the surface (75 m above sea level) to 75 m below sea level. The Akeshi mineralization is the youngest of the three Nansatsu deposits. The Kasuga and Iwato ore bodies formed before deposition of the Nansatsu Middle Volcanics, which host the Akeshi deposit [27]. Primary Au is closely associated with luzonite and pyrite [31]. The mineral paragenesis of the Akeshi deposit is given in Table 1.

2.4. Low-Sulfidation Au Deposits in the Hokusatsu District of Kagoshima-ken

Au mineralization in the Hokusatsu district (Figure 1) is typical of low-sulfidation quartz veins with adularia and calcite. The epithermal Au deposits, including the Hishikari deposit and the Kushikino deposit [32], which is the second largest in Kyushu, have formed from the Pliocene to Pleistocene in the Hokusatsu district.

Hishikari Deposit

The Hishikari deposit, located in the southern part of Kyushu, is the largest Au deposit in Japan (Figure 1). The Hishikari deposit is a low-sulfidation epithermal deposit and is quite famous for its high-grade ore. The Hishikari deposit consists of the Sanjin, Honko, and Yamada ore zones. The Sanjin and Honko ore zones occur in the Cretaceous basement sedimentary rocks of the Lower Shimanto Group. In contrast, veins in the Yamada ore zone occur in Pleistocene andesitic rocks that overlie the basement sedimentary rocks. The mineralization of the Hishikari deposit began at 1.15 Ma and ended by 0.60 Ma [24]. The Hishikari veins consist mainly of quartz, adularia, and smectite, and the most abundant metallic mineral is pyrite. The mineral paragenesis of the Hishikari deposit is given in Table 1. The occurrence of As-bearing minerals, such as arsenopyrite, is not known.

2.5. BIF-Hosted Deposit

Kalahari Goldridge Deposit

The Kalahari Goldridge deposit is a typical lode-gold deposit hosted by BIF in the Archaean Kraaipan Greenstone Belt in the Northwestern Province of South Africa (Figure 2). The genesis of the deposit has been documented [30,33,34]. The Kalahari Goldridge deposit is hosted in a strata-bound iron-formation (3.41 Ga) [35]. It comprises the D zone, A zone, Watertank, and Windmill ore bodies. The entire ore body stretches along a strike length of approximately 6.5 km N–S. The ore zone is characterized by meta-sedimentary rocks.

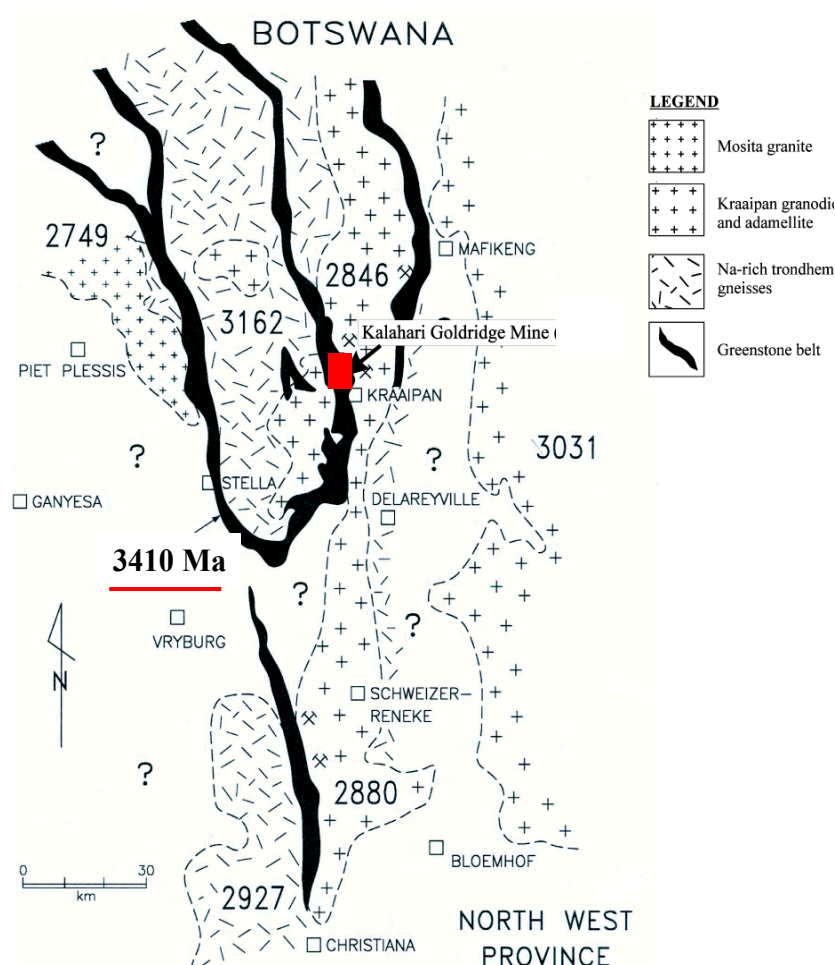


Figure 2. Simplified geologic map of the Kraaipan terrane [35], showing the locality of the Kalahari mine. Numbers indicate approximate ages (Ma) of BIFs and related rocks.

Principal mineralogy at Kalahari Goldridge includes silicates (chlorite, stilpnomelane, minnesotaite, and muscovite), carbonates (siderite, ankerite–dolomite series, and calcite), and opaque minerals (magnetite, pyrite, pyrrhotite, and chalcopyrite) [30,33,34]. The Kalahari Goldridge deposit hosted by BIF was formed under relatively reduced circumstances, and pyrite and pyrrhotite were found in the deposit. The mineral paragenesis of the Kalahari Goldridge deposit is given in Table 1.

Mineralogical studies at the Kalahari Goldridge deposit have revealed that despite the relatively high assay values found in bulk ores, particularly in pyrrhotite-rich areas, samples showed minor or no visible Au grains during microscopic observations [33,34]. These observations suggest that Au likely occurs at submicroscopic levels or is structurally bound in the lattices of the sulfides in these samples. The presence of submicroscopic Au in sulfides has been a major contributor to refractory Au ores, which presents great difficulty in using conventional extractive metallurgy to extract the Au from these ores. At the Kalahari Goldridge mine, Au recovery stands at 80%, indicating the possible occurrence of submicroscopic Au in the ores at the deposit.

2.6. Sedimentary Au Deposits (Witwatersrand Au Fields)

The Witwatersrand sedimentary basin is situated in the central part of the Kaapvaal Craton of South Africa. This basin consists of a 6000 m-thick layer of clastic sedimentary rocks, the Archaean Witwatersrand Supergroup, and yields auriferous and uraniferous quartz-pebble conglomerates (reefs) [36]. The Witwatersrand Supergroup, which is divided into the lower West Rand and upper Central Rand Groups, has contributed more than one-third of the world's total historic production. The bottom of the Witwatersrand Supergroup has been dated to 2970 Ma [37], whereas the top has been dated to 2714 Ma [38]. The reefs are considered to represent vast fluvial fans. The Au is mostly of detrital origin, deposited syngenetically with the conglomerates [39]. Pyrite is ubiquitous in the Witwatersrand Au-bearing conglomerates [37].

The initial source of Au was the surrounding Archaean granite–greenstone terrain of the Kaapvaal Craton [36]. The Au occurred in association with pyrite, carbon, and uranium, and most of it has been subsequently modified during later hydrothermal activity [36]. Although debate is ongoing as to whether the genesis of the Witwatersrand Au deposits can be attributed to a paleo-placer source or a hydrothermal source, the modified placer process (a combination of the two sources) is thought to be more plausible [40].

2.6.1. South Deep Deposit

The South Deep Au mine in the West Rand Goldfield lies at the northwestern edge of the Witwatersrand Basin at latitude 26°25' S, longitude 27°40' E (Figure 3). The mining target reefs are both the Ventersdorp Contact Reef (VCR) and the Upper Elsburg formation conglomerates [36]. The VCR overlies footwall lithologies of the Turffontein Subgroup in the western part of the mining area. The Upper Elsburg reefs consist of multiple stacked reefs. The Upper Elsburg reef is well-understood thanks to information integrated from surface drilling, three-dimensional (3D) surface seismic surveys, and underground long inclined drillings [36]. Au production amounted to 4885 kg in 2018 [41].

2.6.2. KDC West Deposit

The Kloof–Driefontein Complex (KDC) is situated in the Far West Rand Goldfields at latitude 26°24' S and longitude 27°30' E (Figure 3). The Kloof Au Mine (KDC East) and Driefontein Au Mine (KDC West) are situated in the mining area. Au mineralization occurs within quartz-pebble conglomerate reefs, which are developed above unconformable surfaces near the basin margin. The mining target reefs are the Carbon Leader Reef (CL), Ventersdorp Contact Reef (VCR), and Middelvlei Reef (MR). The CL represents most of the current mineral reserves in the KDC West. It varies in thickness from a mere carbon streak to a multiple-band conglomerate of about 200 cm thick [39]. The deepest operating level is 3420 m below the surface (50 Level at 5 Shaft) in the KDC West. The current

mineral reserves will possibly remain operating until 2039 [42]. In 2018, the underground production from the KDC West amounted to 8952 kg [43].

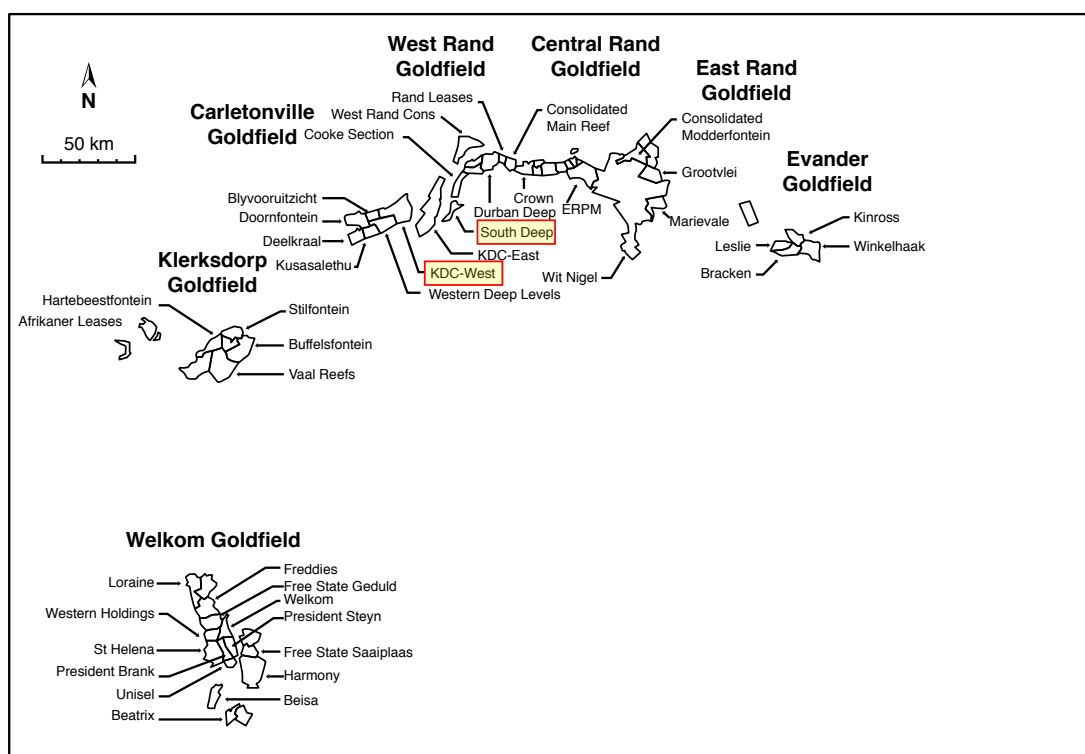


Figure 3. Location of the Witwatersrand gold mines [44], showing the localities of the South Deep and KDC (Kloof–Driefontein Complex) West mines.

3. Materials and Methods

3.1. Materials

Pyrite is the most abundant sulfide and is ubiquitous throughout the high-sulfidation epithermal deposits in the Nansatsu district (the Kasuga, Iwato, and Akeshi deposits) and the low-sulfidation epithermal Hishikari deposit (Table 1). Pyrite samples from the Kasuga deposit were collected from drill core 11MANU-1, which was drilled vertically from the 100 mL (100 m above sea level) bench of the mine [45]. The Iwato pyrite samples were from the 150 mL bench of the Maruyama open pit. The Akeshi pyrite samples containing covellite and native sulfur were from the −50 mL (underground) bench of the No. 1 ore body. Pyrite and pyrrhotite from the Kalahari Goldridge deposit, as well as pyrite from the South Deep and KDC West deposits in the Witwatersrand Basin, South Africa, were sampled for analysis.

The size and shape of the analyzed minerals are described in Table 2. Because the size of a mineral grain is measured on a thin section, it roughly represents the actual 3D substance. Table 2 gives the lengths of the major and minor axes of the grains. The roundness [46], which indicates the degree of abrasion, is another important factor in describing mineral grains in sedimentary-origin Au deposits. The mineral grains assessed in this study were divided into four categories based on their roundness: Crystalline, angular, subangular, and rounded, in order of increasing roundness. The crystalline grains retain their original crystal shapes.

Table 2. Results of secondary ion mass spectrometry (SIMS) analysis.

Deposit	Sample Name	Locality	Analysis Spot	Analysis No.	Mineral	Au (ppm)	As (ppm)	Analysis Position-Size-Roundness
Kasuga	S113	19.6 m depth	N1	Au5*246.adp	pyrite	0.06	44	core-260 μm \times 320 μm -subangular
Kasuga	S113	19.6 m	N11	Au5*256.adp	pyrite	0.04	2	core-the same grain as above
Kasuga	S113	19.6 m	N2	Au5*247.adp	pyrite	0.09	6	core-120 μm \times 130 μm -subangular
Kasuga	S113	19.6 m	N3	Au5*248.adp	pyrite	0.05	44	rim-the same grain as above
Kasuga	S113	19.6 m	N4	Au5*249.adp	pyrite	0.10	2	rim-the same grain as above
Kasuga	S113	20.5 m	N6	Au5*251.adp	pyrite	0.24	3	core-400 μm \times 400 μm -subangular
Kasuga	S113	20.5 m	N7	Au5*252.adp	pyrite	0.17	3	rim-50 μm \times 60 μm -subangular
Kasuga	S113	90.2 m	N8	Au5*253.adp	pyrite	0.17	64	core-200 μm \times 300 μm -subangular
Kasuga	S113	138.3 m	N10	Au5*255.adp	pyrite	0.14	1	core-100 μm \times 250 μm -subangular
Kasuga	S113	19.6 m	N12	Au5*257.adp	pyrite	0.09	2	core-50 μm \times 90 μm -subangular
Kasuga	S113	19.6 m	N13	Au5*258.adp	pyrite	0.09	17	core-50 μm \times 100 μm -subangular
Kasuga	S113	19.6 m	N14	Au5*259.adp	pyrite	0.06	2	core-120 μm \times 140 μm -angular
Kasuga	S113	19.6 m	N15	Au5*260.adp	pyrite	23.58	176	core-20 μm \times 40 μm -subangular
Kasuga	S113	19.6 m	N16	Au5*261.adp	pyrite	22.40	395	core-20 μm \times 30 μm -subangular
Iwato	S176	Maruyama	Py1a	Au5a*285.adp	pyrite	1.59	24	rim-150 μm \times 200 μm -mosaic
Iwato	S176	Maruyama	Py1b	Au5a*287.adp	pyrite	2.84	14	rim-the same grain as above
Iwato	S176	Maruyama	Py4a	Au5a*286.adp	pyrite	0.16	19	core-100 μm \times 200 μm -mosaic
Iwato	S176	Maruyama	Py3a	Au5a*288.adp	pyrite	0.25	43	core-30 μm \times 40 μm -subangular
Akeshi	S174	No.1 ore body	#2	Au5a*290.adp	pyrite	0.17	847	rim-50 μm \times 70 μm -crystalline
Akeshi	S174	No.1 ore body	#4	Au5a*292.adp	pyrite	1.60	1083	rim-20 μm \times 40 μm -subangular
Akeshi	S174	No.1 ore body	#5	Au5a*293.adp	pyrite	0.08	403	core-30 μm \times 50 μm -subangular
Akeshi	S174	No.1 ore body	#6	Au5a*294.adp	pyrite	0.04	420	rim-20 μm \times 50 μm -subangular
Akeshi	S174	No.1 ore body	#3	Au5a*295.adp	pyrite	33.53	703	rim-30 μm \times 60 μm -irregular
Akeshi	S174	No.1 ore body	#7	Au5a*296.adp	pyrite	0.07	56	core-50 μm \times 90 μm -subangular
Akeshi	S174	No.1 ore body	#8	Au5a*297.adp	pyrite	0.54	111	core-50 μm \times 60 μm -crystalline
Kalahari	S79	D zone 118/21/11A	#3	Au4*2.adp	pyrite	0.01	22	core-500 μm \times 1000 μm -subangular
Kalahari	S82	D zone 118/21/15A	K2_2	0703_8 #1	pyrite	0.32	17	core-80 μm \times 100 μm -subangular
Kalahari	S82	D zone 118/21/15A	#31 Po	Au5aa*31.adp	Pyrrhotite	0.53	8	core-250 μm \times 250 μm -mosaic
Kalahari	S82	D zone 118/21/15A	#32 Po	Au5aa*32.adp	Pyrrhotite	0.80	12	core-the same grain as above
Kalahari	S82	D zone 118/21/15A	#23 Po	Au5aa*33.adp	Pyrrhotite	0.86	6	core-the same grain as above
Kalahari	S82	D zone 118/21/15A	K1_3	0703_2 #1	Pyrrhotite	0.87	19	core-the same grain as above
Kalahari	S82	D zone 118/21/15A	K1_4	0703_3 #1	Pyrrhotite	0.25	17	core-the same grain as above
Kalahari	S82	D zone 118/21/15A	K1_5	0703_4 #1	Pyrrhotite	0.07	13	core-the same grain as above
Kalahari	S82	D zone 118/21/15A	K1_6	0703_5 #1	Pyrrhotite	0.01	8	core-50 μm \times 50 μm -mosaic
Kalahari	S82	D zone 118/21/15A	K2_1	0703_7 #1	Pyrrhotite	0.01	8	core-200 μm \times 250 μm -rounded
Kalahari	S82	D zone 118/21/15A	K2_3	0703_9 #1	Pyrrhotite	0.01	8	core-100 μm \times 100 μm -mosaic

Table 2. Cont.

Deposit	Sample Name	Locality	Analysis Spot	Analysis No.	Mineral	Au (ppm)	As (ppm)	Analysis Position-Size-Roundness
South Deep	S196	MIT in Upper Elsburgs	S11	0702_4 #1	pyrite	0.04	21	core-350 μm \times 450 μm -rounded
South Deep	S196	MIT in Upper Elsburgs	S12	0702_5 #1	pyrite	0.02	19	rim-the same grain as above
South Deep	S196	MIT in Upper Elsburgs	S13	0702_6 #1	pyrite	0.03	103	core-200 μm \times 400 μm -subangular
South Deep	S196	MIT in Upper Elsburgs	S14	0702_7 #1	pyrite	0.03	827	rim-the same grain as above
South Deep	S196	MIT in Upper Elsburgs	S15	0702_8 #1	pyrite	0.22	1988	core-200 μm \times 270 μm -angular
South Deep	S196	MIT in Upper Elsburgs	S16	0702_9 #1	pyrite	0.34	2297	rim-the same grain as above
South Deep	S196	MIT in Upper Elsburgs	S26	1026_6 #1	pyrite	0.14	2789	rim-the same grain as above
South Deep	S196	MIT in Upper Elsburgs	S27	1026_7 #1	pyrite	1.09	4042	core-the same grain as above
South Deep	S196	MIT in Upper Elsburgs	S21	1026_1 #1	pyrite	0.05	8	core-300 μm \times 600 μm -round
South Deep	S196	MIT in Upper Elsburgs	S22	1026_2 #1	pyrite	0.05	115	rim-160 μm \times 250 μm -subangular
South Deep	S196	MIT in Upper Elsburgs	S23	1026_3 #1	pyrite	0.04	49	core-the same grain as above
South Deep	S196	MIT in Upper Elsburgs	S24	1026_4 #1	pyrite	0.05	102	core-200 μm \times 300 μm -subangular
South Deep	S196	MIT in Upper Elsburgs	S25	1026_5 #1	pyrite	0.04	177	rim-the same grain as above
South Deep	S196	MIT in Upper Elsburgs	S28	1026_8 #1	pyrite	0.03	33	rim-600 μm \times 650 μm -rounded
South Deep	S196	MIT in Upper Elsburgs	S29	1026_9 #1	pyrite	0.02	19	core-the same grain as above
South Deep	S196	MIT in Upper Elsburgs	S30	1026_10 #1	pyrite	0.08	59	core-50 μm \times 80 μm -subangular
KDC West	S197	2B (Carbon Leader)	D3	Au5aa*1.adp	pyrite	0.12	1546	core-250 μm \times 350 μm -subangular
KDC West	S197	2B (Carbon Leader)	D1	Au5aa*3.adp	pyrite	0.22	1238	rim-250 μm \times 350 μm -subangular
KDC West	S197	2B (Carbon Leader)	D2	Au5aa*2.adp	pyrite	0.62	200	rim-300 μm \times 300 μm -rounded
KDC West	S197	2B (Carbon Leader)	D14	0702_3 #1	pyrite	0.63	153	rim-the same grain as above
KDC West	S197	2B (Carbon Leader)	D13	0702_2 #1	pyrite	0.42	100	core-the same grain as above
KDC West	S197	2B (Carbon Leader)	D12	0702_1 #1	pyrite	0.04	22	core-150 μm \times 200 μm -rounded

#1: determined by Atmosphere and Ocean Research Institute (AORI) NanoSIMS at the University of Tokyo.

3.2. Analytical Methods

The pyrite and pyrrhotite were quantitatively analyzed ($3\ \mu\text{m} \times 3\ \mu\text{m}$ field) using a Cameca IMS 1270 spectrometer at the Geological Survey of Japan (GSJ), National Institute of Advanced Industrial Science and Technology (AIST). A defocused Cs^+ primary ion beam was restricted to $20\ \mu\text{m}$ in diameter by a circular aperture to obtain a homogeneous primary beam of approximately $100\ \text{pA}$. Sputtered negative secondary ions with an accelerating voltage of $-10\ \text{kV}$ were extracted from the carbon-coated polished sample surface and were detected using a total impact energy of $20\ \text{kV}$ (primary accelerating voltage of $+10\ \text{kV}$ and secondary extraction voltage of $-10\ \text{kV}$). The secondary $^{197}\text{Au}^-$ and $^{75}\text{As}^-$ ions (both are monoisotopic), as well as the matrix $^{34}\text{S}^-$ used as an internal reference, were detected using an electron multiplier (EM). A rectangular field aperture of $300\ \mu\text{m} \times 300\ \mu\text{m}$ was introduced into the secondary ion optics, limiting the analyzed area on the sample surface to a central square measuring $3\ \mu\text{m} \times 3\ \mu\text{m}$ to avoid the crater edge effect. The centering of the primary beam was ascertained each time by looking at the beam position in the field aperture of $2\ \text{mm} \times 2\ \text{mm}$ (primary beam of $20\ \mu\text{m}$ in diameter). An electron flood gun was used on the sample surface for the purpose of charge compensation to achieve better stability of the secondary ions. The isobaric interference of $^{133}\text{Cs}^{32}\text{S}_2$ is clearly separated from ^{197}Au ($M/\Delta M = 1728$) by a mass-resolving power of 3300 without energy filtering [20]. Measurements at 3, 5, and 10 s for $^{34}\text{S}^-$, $^{75}\text{As}^-$, and $^{197}\text{Au}^-$ ions, respectively, were allocated for each measurement cycle. It takes 24 s for a one-cycle measurement, which includes the waiting time for the magnetic hysteresis stabilization. The detection limits, which were computed as three times the standard deviation of the background noise, for the Au and As measurements were 15 ppb and 0.6 ppm, respectively, for an ^{34}S count of 3×10^5 cps if the background noise is assumed to have a Poisson distribution. The analysis size using the IMS 1270 was $3\ \mu\text{m}$.

The sulfur isotope ratios of pyrite grains from the KDC West deposit were measured using the multi-collection system of the SIMS spectrometer with Faraday cups. The analyzed area on the sample surface is a circle with a diameter of $20\ \mu\text{m}$, the same area as the primary beam bombardment; therefore, depth profiling is not available. The primary beam intensity was increased to $1\ \text{nA}$ to obtain a high intensity of ^{32}S secondary ions (2×10^8 cps) for high-precision sulfur isotope measurements. The measured data were normalized using a standard pyrite sample ($\delta^{34}\text{S} = +5.1\text{‰}$) and are reported relative to Canyon Diablo Troilite (CDT). The accuracy is less than $\pm 0.2\text{‰}$ (2σ). The analysis size using the IMS 1270 for sulfur isotope ratios was $20\ \mu\text{m}$.

A NanoSIMS 50 at the Atmosphere and Ocean Research Institute (AORI) of the University of Tokyo was used for the analysis of several samples, as indicated in Table 2. After pre-sputtering with a $500\ \text{pA}$ Cs^+ rastered beam over a $30\ \mu\text{m} \times 30\ \mu\text{m}$ area to remove the carbon coat from the sample surface and stabilize the secondary ion intensity by Cs^+ implantation, a focused 30 to $50\ \text{pA}$ Cs^+ primary ion beam was scanned over the inner $20\ \mu\text{m} \times 20\ \mu\text{m}$ area for analysis. An electron flood gun was used for the purpose of charge compensation. Negative secondary ions were extracted by an accelerating voltage of $8\ \text{kV}$ and introduced into the mass spectrometer. The matrix $^{34}\text{S}^-$ and the secondary $^{75}\text{As}^-$ and $^{197}\text{Au}^-$ ions were measured simultaneously using multi-collectors with EMs. The beam intensity of $^{34}\text{S}^-$ was approximately 3.3×10^5 to 3.8×10^5 cps. The dead time ($\sim 44\ \text{ns}$) of the counting system was corrected. The entrance slit was set to a width of approximately $40\ \mu\text{m}$, and the exit slits for each EM were set to a width of $75\ \mu\text{m}$. A mass-resolving power of 8000 was attained with adequate flat-topped peaks. To reduce background noise and contamination from the area surrounding the analyzed part, beam blanking was adopted. Only the inner approximately $10\ \mu\text{m} \times 10\ \mu\text{m}$ area of the scanned area was measured (32×32 pixels out of 64×64 pixels), meaning that only 25% of the total secondary ions produced were counted. It took about 1000 s to carry out one analysis run of 50 cycles. The detection limits, computed as three times the standard deviation of the background noise, for the Au and As measurements were 7 ppb and 0.3 ppm, respectively, for a ^{34}S count of 3.3×10^5 cps if the background noise is assumed to have a Poisson distribution. The analysis size using the NanoSIMS 50 was $10\ \mu\text{m}$.

The pyrite grains in ore samples from the Witwatersrand Au fields were analyzed by electron probe microanalysis (EPMA) for Fe, Cu, Co, Ni, As, and S using a JEOL JXA-8900 probe at Shizuoka University operated at 20 kV and 12 nA, with natural pyrite (FeS₂) and Arsenopyrite (FeAsS) and pure metal samples (Co, Cu, Ni) as standards. The detection limit was 0.01 wt% for all the measured elements. The analysis size was 10 μm.

3.3. Standardization

Quantitative SIMS analysis requires the use of reference standards for each element and the analyzed host mineral because of matrix effects on secondary ion intensities. Standard samples for Au measurements in pyrite and pyrrhotite and for As measurements in pyrite were prepared. Calibration of the Au concentration in a mineral was conducted using external standard samples of pyrite and pyrrhotite that were implanted with ¹⁹⁷Au ions at a dosage of 1×10^{15} atoms/cm² with an implantation energy of 1.6 MeV using the 1.7 MV Tandetron accelerator located at the University of Western Ontario, Canada [20]. Calibration of the As concentration in pyrite was also performed using external standard samples that were implanted with As ions at a dosage of 3×10^{15} atoms/cm², with an implantation energy of 0.4 MeV. After analyzing a standard for depth profiling, the depth of the sputtered borehole was measured using a surface profiler to determine the relative sensitivity factor (RSF), which was used to calculate the concentration in the unknown sample [20]. The calibration of the As concentration in pyrrhotite was conducted using EPMA-determined pyrrhotite samples.

4. Results

The results of the Au and As measurements are given in Table 2. Because the detection limits for Au and As measurements were 15 ppb and 0.6 ppm, respectively, at the GSJ SIMS, the last figures of the Au and As concentrations have some errors at low-level intensities. The detection limits of 8 ppb and 0.3 ppm for Au and As measurements, respectively, at the AORI SIMS might be low enough to guarantee the whole figures in Table 2.

The Au and As concentrations in the pyrite from the Kasuga deposit vary from 0.04 to 24 ppm and from 1 to 400 ppm, respectively (Table 2). The Au and As concentrations in the pyrite from the Iwato deposit vary from 0.16 to 2.8 ppm and from 14 to 43 ppm, respectively (Table 2). The Au and As concentrations in pyrite from the Akeshi deposit vary from 0.04 to 34 ppm and from 56 to 1100 ppm, respectively (Table 2). The Iwato pyrite grains contain less As than those from the other deposits.

The Au and As concentrations in the pyrite and pyrrhotite from the Kalahari Goldridge deposit are given in Table 2. In addition to the above results, substantial Au analytical data without As content was obtained at the beginning of this study (Table A1). The Au concentrations in pyrrhotite and pyrite vary from 0.01 to 5.62 ppm and from 0.01 to 0.32 ppm, respectively (Table 2, Table A1), whereas the As concentrations in pyrrhotite and pyrite are both low, with a narrow range from 6 to 22 ppm (Table 2). Although the Au–As relationship cannot be ascertained from the Au data alone (Table A1), Au depth profiles were acquired during every Au measurement. The Au concentration of pyrite is very low (less than 0.3 ppm).

The analytical results of the pyrite from the South Deep and KDC West deposits in the Witwatersrand Basin are given in Table 2. The Au and As concentrations in the pyrite from the South Deep deposit vary from 0.02 to 1.1 ppm and from 8 to 4000 ppm, respectively (Table 2). The Au and As concentrations in the pyrite from the KDC West deposit vary from 0.04 to 0.6 ppm and from 22 to 1500 ppm, respectively (Table 2).

The EPMA analysis results for Fe, Cu, Co, Ni, As, and S in pyrite from the South Deep and KDC West deposits are given in Tables A2 and A3, respectively. The analysis positions were allocated near the SIMS analysis points. Most pyrites contain Co and Ni in addition to As, but the Cu concentration is generally negligible (Tables A2 and A3).

Two sulfur isotope ratios of the pyrite from the KDC West deposit were obtained. The analysis positions were near D3 (core 250 $\mu\text{m} \times 350 \mu\text{m}$ subangular) and near D2 (rim 300 $\mu\text{m} \times 300 \mu\text{m}$ rounded). The $\delta^{34}\text{S}$ values in these positions are +3.0‰ and +5.8‰, respectively.

5. Discussion

5.1. Epithermal Au Deposits

Figure 4 shows the Au and As concentrations in pyrite from the high-sulfidation Nansatsu-type Au deposits. Three samples from the Nansatsu-type Au deposits have relatively high Au concentrations of around 30 ppm (Figure 4, Table 2), which are more than the solubility of Au in synthesized pure pyrite (up to 3 ppm; [47]). Concerning natural arsenic pyrite, relatively large Au^+ ions are able to enter the crystal structure when the structure is distorted by As incorporation into the structure [3,48]. Among the three high Au samples, two small grains of less than 20 μm from the Kasuga deposit have high Au concentrations, and the appearance of the pyrite from the Akeshi deposit is irregular. On the other hand, the low-Au pyrite grains, which contain less Au than the solubility of Au in synthesized pure pyrite [47] regardless of the As content, are all bigger than 50 μm , with a crystalline appearance. The fluid in the stage (probably late stage) might contain little Au but a moderate amount of As, and the fluid might be slightly saturated with pyrite.

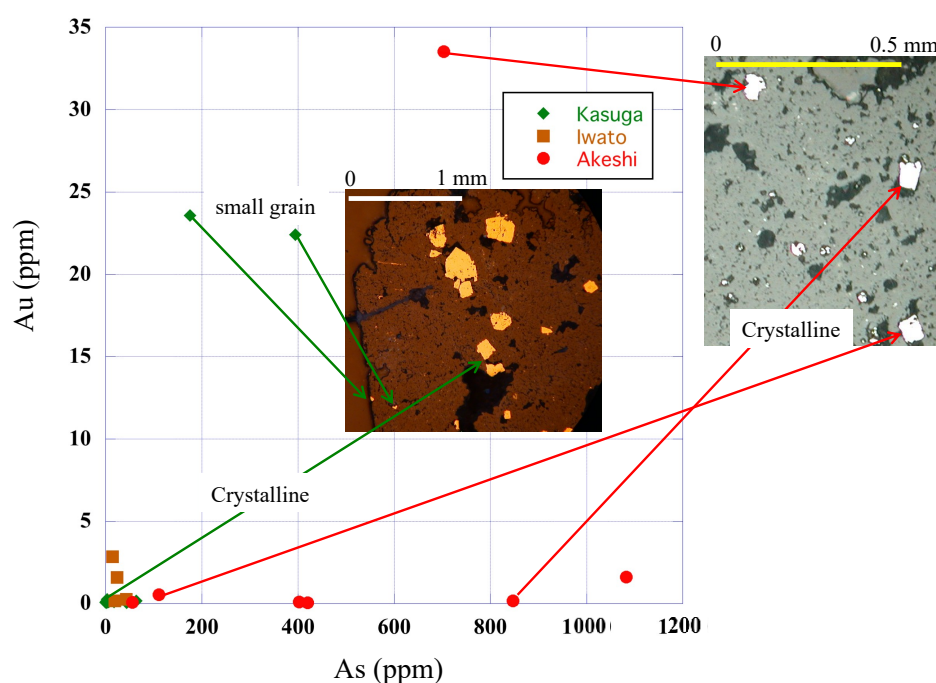


Figure 4. Au and As concentrations (ppm) in pyrite from the high-sulfidation Nansatsu-type (Kasuga, Iwato, and Akeshi) gold deposits. Reflected-light photos showing the shapes of pyrite grains are added to the graph. See Section 5.1 for details.

Regarding the As content, the pyrite grains from the Iwato deposit contain very little As. The other two deposits, the Kasuga and Akeshi deposits, contain relatively high concentrations of As in their pyrite grains. In accordance with the As content in pyrite, the As-containing minerals enargite and luzonite (both of the chemical compositions are Cu_3AsS_4) occur in the Kasuga and Akeshi deposits, respectively; however, no such minerals were found in the Iwato deposit (Table 1).

At the Hishikari deposit (a typical low-sulfidation Au deposit), the specific value of the Au–As ratio in pyrite proves to be characteristic of the specific Au-bearing vein and its specific depth [20]. The Au–As ratio (3.4×10^{-2}) is high in the Sanjin ore zone, which consists of high-grade veins in or above the Cretaceous accretionary sedimentary rocks. The Au–As ratio (1×10^{-2}) is moderate in the

Honko ore zone, which consists of several high-grade veins, and also occurs in or above the basement sedimentary rocks [20]. In contrast, the Au–As ratio (0.4×10^{-2}) is gentle in the Yamada ore zone, which consists of relatively low-grade veins, and occurs in Pleistocene andesitic pyroclastic rocks that unconformably overlie the basement sedimentary rocks [20]. According to [20], it may be possible to infer the nature of unknown veins by determining their Au–As ratios.

5.2. Kalahari Goldridge Deposit

The Au and As concentrations in the pyrite and pyrrhotite from the Kalahari Goldridge deposit are both low (Table 2), and no relationship was found between the two elements. However, the Au concentration in the pyrrhotite varies in a wide range from 0.01 to 5.6 ppm (Table A1). Therefore, the occurrence of Au in pyrrhotite was examined by SIMS depth analysis.

Figure 5 shows the SIMS depth profile of Au in a Kalahari pyrrhotite sample (Au1*14.adp in Table A1). The sample surface was bombarded by a Cs^+ beam, which formed a crater on the sample surface, and the Au^- ion count was measured along with the matrix $^{34}\text{S}^-$ ion count as a reference. The measurement of the depth of the resulting crater allows the conversion of the time scale to a depth scale (horizontal axis in Figure 5) because the primary beam current stability was fairly good (fluctuation of 1%–2%) during the measurement. The left-hand vertical axis shows the Au intensity in counts per second, and the right-hand vertical axis shows the raw data converted into the Au concentration in parts per million based on the relative sensitivity factor (RSF). Because the Au intensity is normalized by the intensity of the reference ^{34}S , the measured Au variation is not caused by instrumental instability or fluctuation but represents the actual variation in Au content with the depth. Figure 5 shows that the pyrrhotite contains Au in concentrations that vary with depth from 1 to 16 ppm, with an average concentration of 5.6 ppm (Table A1). The presence of several peaks in the Au concentration with widths of 50 nm, as shown in Figure 5, suggest that Au occurs as NPs rather than as a solid solution in the pyrrhotite lattice. The high spatial resolution of 3 nm along the depth direction in the SIMS analysis is characteristically useful in assessing the presence of Au NPs.

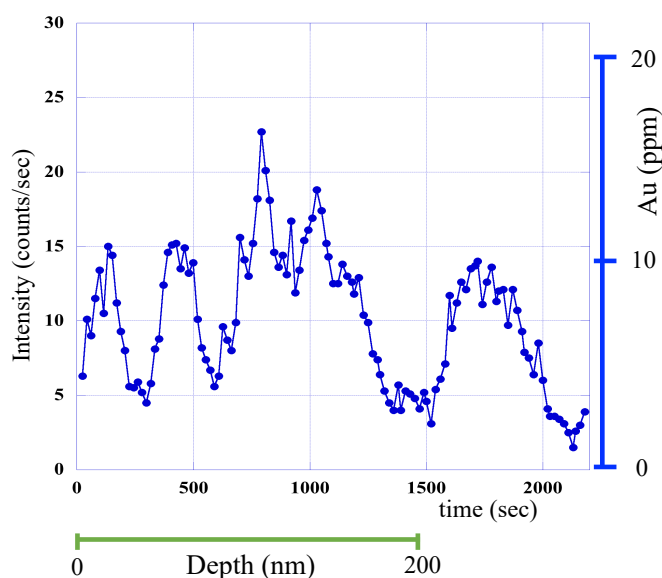


Figure 5. Secondary ion mass spectrometry (SIMS) depth profile of Au in pyrrhotite (Au1*14.adp in Table A1) from the Kalahari Goldridge deposit. The analysis started at the surface of the pyrrhotite grain (left end of the graph) and proceeded toward the interior (right end of the graph), resulting in the formation of a crater. The left vertical axis shows the raw data in counts per second. The right vertical axis shows the Au concentration (ppm) in pyrite calculated from the raw data using the $\text{RSF}_{\text{Au-pyrrhotite}}$ value. Note that the crater is very shallow and several peaks of Au with of approximately 50 nm in width (Au nanoparticles (NPs)) were observed. RSF: relative sensitivity factor.

Au NPs in pyrrhotite are not uncommon in the Kalahari Goldridge deposit, so similar depth profiles in pyrrhotite are shown in Figure A1 (34S_Au4*111.adp in Table A1) and Figure A2 (34S_Au4*105.adp in Table A1). In contrast, pyrite has little Au, and no Au NPs were detected in the Kalahari Goldridge deposit samples. The depth profile in Figure A3 (Au1*47.adp in Table A1) shows only low signals. The existence of Au NPs is important when the preconcentration of Au by sulfide flotation is considered.

5.3. Witwatersrand Au Field

Pyrite grains from the South Deep deposit show a positive correlation between Au and As (Figure 6). The coefficient of determination R^2 is 0.71. The angular grains have more As than the subangular or rounded grains (Figure 6, Table 2). Figure 7 shows an area of lower concentrations of the Au–As relationship in the subangular and rounded pyrite grains from the South Deep deposit. The rims of the pyrite grains contain more As than their cores do (Figure 7, Table 2). Both the rim and core of the rounded pyrite grains contain little Au or As.

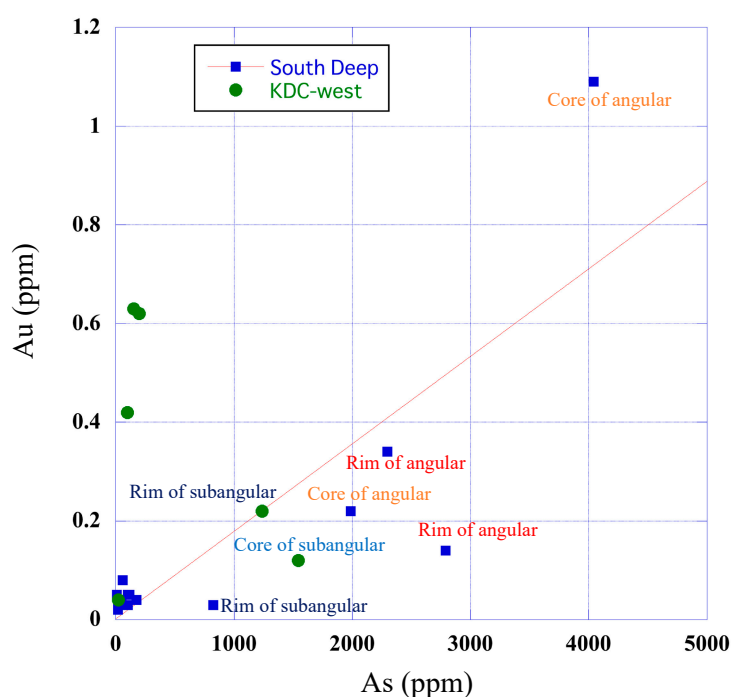


Figure 6. Relationship between Au and As concentrations (ppm) in pyrite from the South Deep and KDC West deposits. A rough positive correlation ($R^2 = 0.71$) between the Au and As concentrations was obtained for the South Deep pyrite, whereas the relationship for the KDC West deposit showed different behavior from that of the South Deep deposit. The analysis position (rim or core) and the shape of the pyrite are given near the symbol.

The roundness of the grains is an indicator of the degree of abrasion during transportation in the sedimentary basin. Conversely, angularity may be caused by precipitation from various types of hydrothermal fluids. In the latter case, the rim of a grain may be composed of precipitation from a fluid while the core is of sedimentary origin. These suggest a hydrothermal influence on pyrite formation in the South Deep deposit. The rounded pyrite is sedimentary in origin, and the angular pyrite is thought to have precipitated from hydrothermal fluids. The subangular pyrite may be partly influenced by a late-stage hydrothermal fluid. The As concentration in rounded pyrite from the KDC West deposit is also low, which is similar to the case of the South Deep rounded pyrite; however, in contrast to the South Deep rounded pyrite, its Au content varies (Figure 6).

The $\delta^{34}\text{S}$ values of the D3 sample (subangular pyrite) and the D2 sample (rounded pyrite) from the KDC West deposit are +3.0‰ and +5.8‰, respectively. In a previous paper, the broader range

of the $\delta^{34}\text{S}$ values for rounded pyrite (from -4.7‰ to $+6.7\text{‰}$) confirms its heterogeneous sulfur composition [49]. This heterogeneity supports the presumption that the rounded pyrite has a placer origin [40,49]. Additionally, the euhedral pyrite and pyrite overgrowths have a restricted $\delta^{34}\text{S}$ range (from -0.5‰ to $+2.5\text{‰}$) [49]. The $\delta^{34}\text{S}$ values of the rounded and subangular pyrite grains at the KDC West deposit are consistent with published data. The $\delta^{34}\text{S}$ value of subangular pyrite is slightly higher than the upper limit of the range because the subangular grain might be a mixture of pyrites with placer and hydrothermal origins. Concerning the subangular pyrite from the KDC West deposit, the Au–As relationship is similar to that of the angular pyrite from the South Deep deposit (Figure 6), which might also be influenced by hydrothermal fluids. The shape of the pyrite grains might prove to be an indicator of hydrothermal influence on sedimentary deposits.

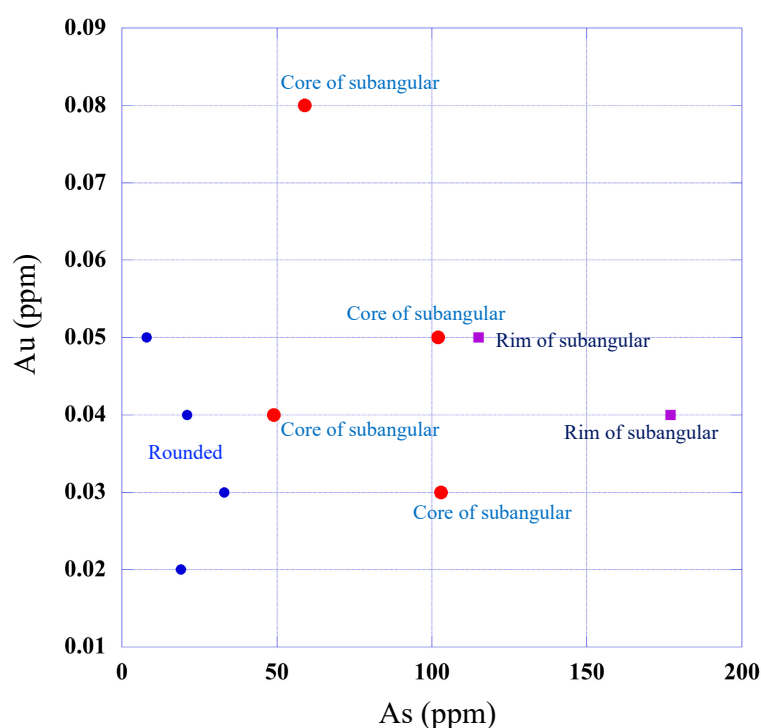


Figure 7. Relationship between Au and As concentrations (ppm) in pyrite from the South Deep deposit (an area of lower Au and As concentrations of Figure 6). The analysis position (rim or core) and the shape of the pyrite are given near the symbol. Violet and red symbols represent data collected at the rim and core of subangular grains, respectively. Rounded grains are shown in blue.

5.4. Au–As Relationship

Figure 8 shows all the SIMS data of the Au and As concentrations in pyrite and pyrrhotite (Table 2) in a log–log plot of their molar content. The solid line represents the empirical solubility limit of Au as a function of the As concentration [18]. Almost all the data lie on or below the line. The blue symbols represent high-sulfidation epithermal pyrite from the three Nansatsu-type deposits. The data are scattered over wide ranges of both Au and As contents. The analysis of one sample from the Kasuga deposit (Au5*260.adp, Table 2) is above the Au solubility limit; however, there are no Au NPs in the pyrite, because the depth profile shows constant Au intensity. Green symbols represent pyrite grains from Au deposits in the Witwatersrand Au field. The Au and As concentrations are scattered over relatively wide ranges, but a high Au content is not found in any sample. Orange circles represent pyrite (solid) and pyrrhotite (open) from the BIF-hosted Kalahari deposit. They have a narrow range of As concentrations but are scattered over a wide range of Au concentrations. Other than the rough positive correlation between Au and As concentrations in the South Deep deposit ($R^2 = 0.71$; Figure 6), clear Au–As correlations are not seen in Figure 8. The reason why there are few Au–As correlations

might partly be because of later-stage hydrothermal Au and As precipitations onto the existing pyrite surfaces, and additionally, because the majority of Au concentrations in Figure 8 are below 3 ppm (experimental Au solubility limit for pure pyrite) [47].

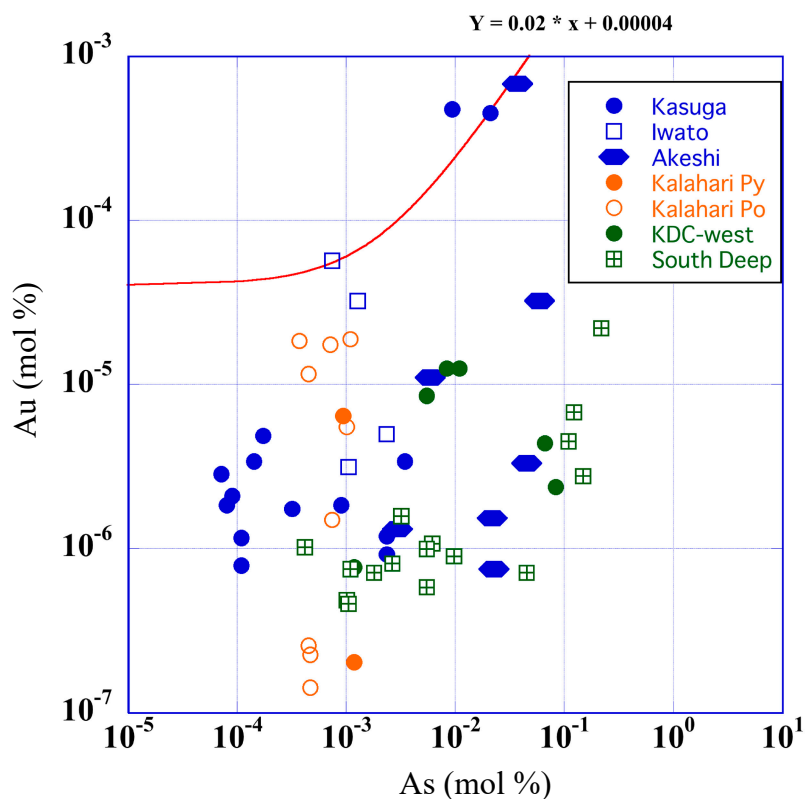


Figure 8. Log–log plot of the relationship between the Au and As concentrations (mol%), including all the SIMS data for the pyrite and pyrrhotite. Every deposit is allocated a symbol in the legend. The solid line represents the empirical solubility limit of Au in pyrite as a function of As concentration [18]. Note that almost all the data lie on or below the line. See Section 5.4 for details. The labels ‘Kalahari Py’ and ‘Kalahari Po’ represent pyrite and pyrrhotite grains, respectively, from the Kalahari deposit.

Figure 5 shows that the Au in pyrrhotite grains exists as Au NPs. Although the As concentration of the grains was not analyzed, the narrow range of As concentrations for the Kalahari deposit would restrict the value. The Au content of 5.62 ppm (Table A1) corresponds to 1.23×10^{-4} mol%, which is above the solubility limit of Au for pyrite in Figure 8 when the As concentration is in the range for the Kalahari pyrrhotite. In contrast, the As concentrations of pyrrhotite for A1 (34S_Au4*111.adp in Table A1) and A2 (34S_Au4*105.adp in Table A1) correspond to 2.8×10^{-5} and 1.9×10^{-5} mol%, respectively, which are below the solubility limit of Au for pyrite in Figure 8 when the As concentration is in the range for the Kalahari pyrrhotite. This suggests either the Au solubility limit of pyrrhotite is lower than that for pyrite or Au could exist in the form of Au NPs despite the sulfide being below the Au solubility limit [20,50]. In any case, the information on the existence of Au NPs is of great significance in the processing of Au-bearing ore by sulfide flotation because sufficient extraction of Au depends on the mode of Au occurrence in the sulfides.

Figure 9 compares the Au–As relationship with that from the high-grade low-sulfidation epithermal Hishikari data [20]. The Au solubility limit is also shown. The Hishikari pyrite grains contain more Au and As than the pyrite and pyrrhotite grains analyzed in this study. The As range of the Hishikari host rocks is scattered similarly to that of the Hishikari ore, and the Au range of the Hishikari host rocks is similar to that of the Witwatersrand Basin.

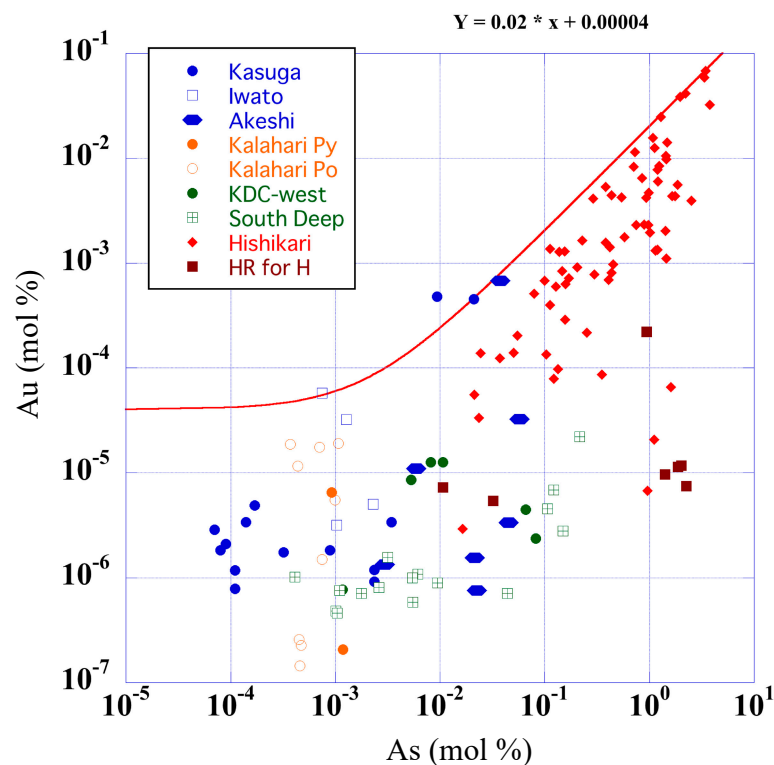


Figure 9. Log–log plot of the relationship between the Au and As concentrations (mol%), including all the SIMS data for the pyrite and pyrrhotite. The Au–As relationship in pyrite from the Hishikari deposit [20] is included for reference. The solid line represents the empirical solubility limit of Au in pyrite as a function of As concentration [18]. The labels ‘Hishikari’ and ‘HR for H’ represent pyrite grains in ore veins and host rocks, respectively, from the Hishikari deposit.

6. Conclusions

Pyrite is the principal sulfide of most Au deposits and is ubiquitous in such deposits. The Au–As relationships in pyrite and pyrrhotite from epithermal, BIF-hosted, and sedimentary Au deposits were determined by SIMS. The Au and As concentrations in pyrite varied from 0.04 to 30 ppm and from 1 to 1000 ppm, respectively, in the high-sulfidation Nansatsu-type epithermal deposits. Both concentrations are lower than those of the low-sulfidation epithermal Hishikari deposit, where the Au and As concentrations in pyrite range from 0.1 to 2800 ppm and from 0 to 70,000 ppm (7 wt%), respectively [20]. The Au concentrations in pyrrhotite and pyrite reach 6 and 0.3 ppm, respectively, in the Kalahari Goldridge BIF-hosted Au deposit, and the Au contained in the pyrrhotite might sometimes exist in the form of NPs, whereas the As concentrations in pyrrhotite and pyrite are low and lie in a narrow range from 6 to 22 ppm. The existence of Au NPs is important when preconcentration of Au by sulfide flotation is considered, while the existence of lattice Au cannot be confirmed. The Au and As concentrations in pyrite from the Witwatersrand Au field range from 0.02 to 1.1 ppm and from 8 to 4000 ppm, respectively. The Au and As concentrations of almost all the pyrites lie on or below the Au solubility limit [18], indicating that Au^+ is the dominant form of Au in the studied pyrites. The shape of the pyrite from a sedimentary deposit might prove to be an indicator of the hydrothermal influence on the deposit, which is supposed to relate with the genesis of the deposit. This study will further demonstrate that the analysis of Au and As concentrations in pyrite from other ore deposits could elucidate the genesis of the deposits.

Author Contributions: Conceptualization: Y.M., N.Q.H., and H.U. Methodology: Y.M. Formal analysis: Y.M. Investigation: Y.M., Y.S., and N.T. (SIMS); and K.M. and R.K. (EPMA). Resources: Y.M., N.Q.H., K.M., R.K., and H.U. Writing—original draft preparation: Y.M. Writing—review and editing: All authors. Project administration: Y.M., N.Q.H., and H.U. Funding acquisition: Y.M.

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Conflicts of Interest: The authors declare no conflict of interest.

Appendix A

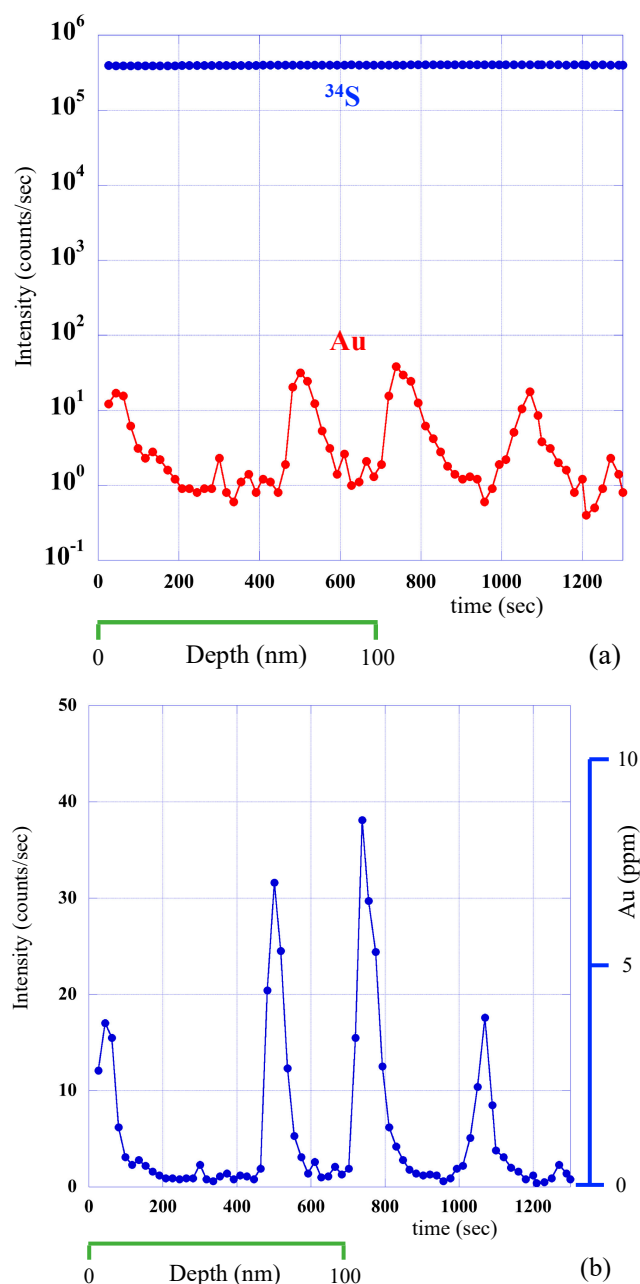


Figure A1. Depth profiles in pyrrhotite (34S_Au4*111.adp in Table A1) from the Kalahari Goldridge deposit. (a) Au and ^{34}S intensities (cps) on a logarithmic scale plotted against time. The corresponding depth is shown along the time scale. (b) Au intensity on a linear scale. The corresponding Au concentration is shown along the intensity scale. Several peaks corresponding to Au NPs (nanoparticles) are present in the profile.

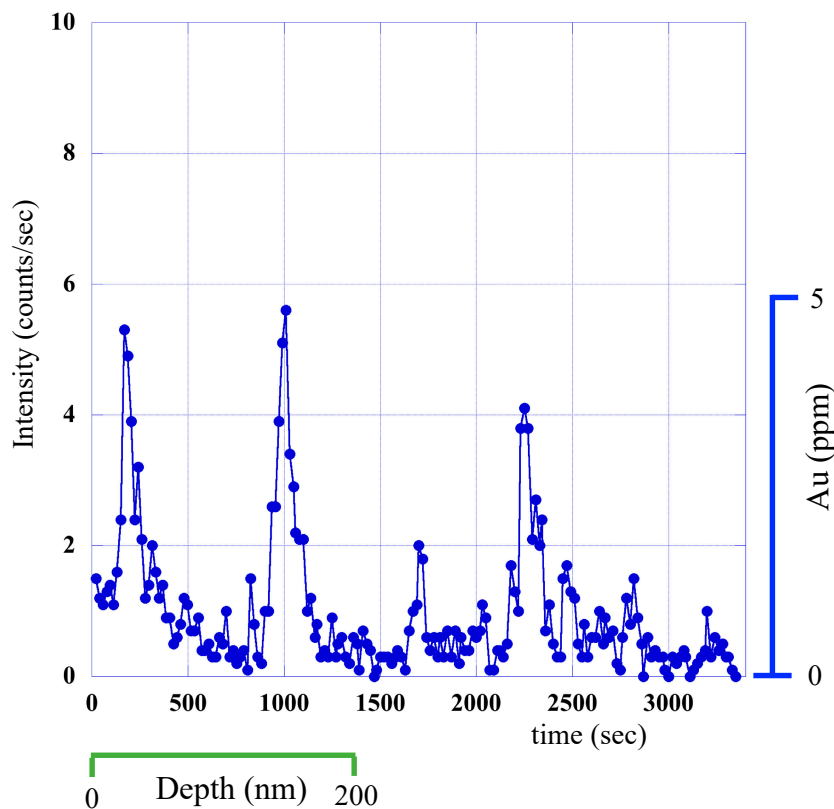


Figure A2. Depth profile of Au concentration in pyrrhotite (34S_Au4*105.adp in Table A1) from the Kalahari Goldridge deposit. Several peaks corresponding to Au NPs are present in the profile.

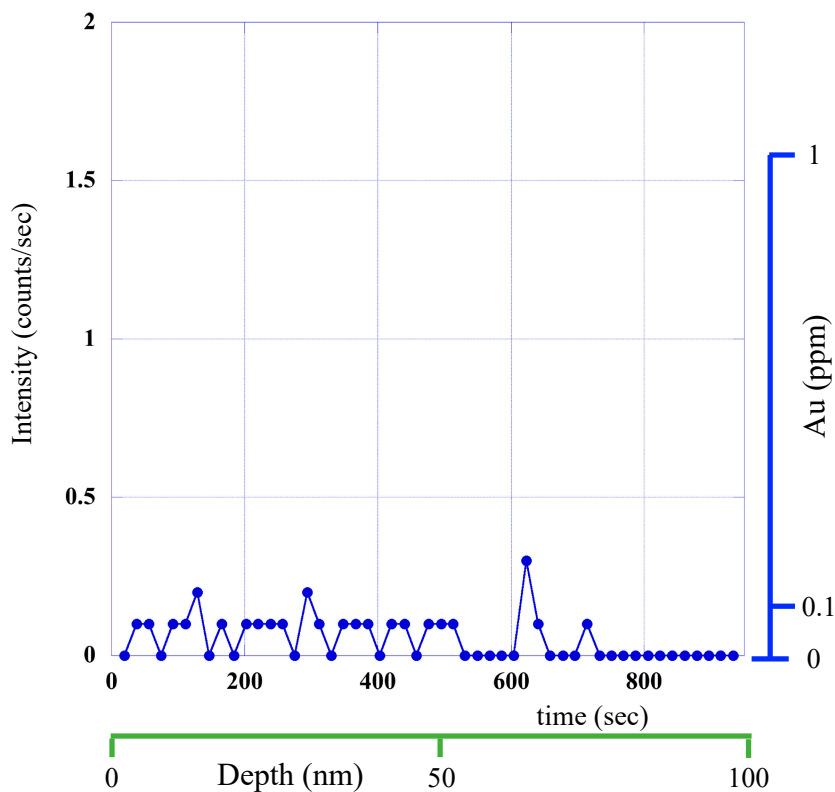


Figure A3. Depth profile of Au concentration vs. depth in pyrite (Au1*47.adp in Table A1) from the Kalahari Goldridge deposit. No peaks corresponding to Au NPs are present in the profile. The Au concentration is very low and includes EM (electron multiplier) background noise.

Appendix B

Table A1. Results of the SIMS Au analysis of pyrite and pyrrhotite grains from the Kalahari deposit.

Sample Name	Deposit	Loc	Analysis Spot	Analysis No.	Mineral	Au (ppm)
GDP636/72/B	D zone		#9	Au1*13.adp	Pyrrhotite	0.05
GDP636/72/B	D zone		#10	Au1*14.adp	Pyrrhotite	5.62
GDP636/72/B	D zone		#12	Au1*16.adp	Pyrrhotite	0.06
118/21/15A	D zone	north pit BIF	#6 py-1	Au1*23.adp	Pyrite-crystalline	0.26
118/21/15A	D zone	north pit BIF	#7 po-3	Au1*24.adp	Pyrrhotite	0.05
118/21/15A	D zone	north pit BIF	#11 po-5	Au1*28.adp	Pyrrhotite	0.69
118/21/15A	D zone	north pit BIF	#14 po-8	Au1*31.adp	Pyrrhotite	0.55
118/21/15A	D zone	north pit BIF	#16 po-10	Au1*34.adp	Pyrrhotite	1.75
DD18/18/1B	D zone	south pit BIF	#2 po-1	Au1*45.adp	Pyrrhotite	0.02
DD18/18/1B	D zone	south pit BIF	#3 py-1	Au1*47.adp	Pyrite	0.03
DD18/18/1B	D zone	south pit BIF	#5 po-2	Au1*48.adp	Pyrrhotite	0.04
DD18/18/1B	D zone	south pit BIF	#6 po-3	Au1*49.adp	Pyrrhotite	0.04
DD18/18/1B	D zone	south pit BIF	#8 po-4	Au1*51.adp	Pyrrhotite	0.09
118/21/11A	D zone	north pit BIF	#3	Au1*52.adp	Pyrite	0.04
118/21/11A	D zone	north pit BIF	#8	Au1*56.adp	Pyrrhotite	0.14
118/21/11A	D zone	north pit BIF	#9	Au1*57.adp	Pyrrhotite	0.28
GDP636/72/B	D zone		#3 py-1	Au1*58.adp	Pyrite	0.14
GDP636/72/B	D zone		#8 po-4	Au1*64.adp	Pyrrhotite	0.11
GDP636/72/B	D zone		#9 po-5	Au1*65.adp	Pyrrhotite	0.09
118/21/8B	D zone	north pit BIF	#2 py-1	Au1*68.adp	Pyrite	0.03
118/21/8B	D zone	north pit BIF	#5 po-2	Au1*70.adp	Pyrrhotite	0.07
118/21/8B	D zone	north pit BIF	#7 py-2	Au1*73.adp	Pyrite	0.01
118/21/8B	D zone	north pit BIF	#9 po-5	Au1*75.adp	Pyrrhotite	0.07
118/21/15A	D zone	north pit BIF	#20 po-20	34S_Au4*102.adp	Pyrrhotite	0.08
118/21/15A	D zone	north pit BIF	#21 po-21	34S_Au4*103.adp	Pyrrhotite	0.10
118/21/15A	D zone	north pit BIF	#22 po-22	34S_Au4*104.adp	Pyrrhotite	1.29
118/21/15A	D zone	north pit BIF	#22 po-22	34S_Au4*111.adp	Pyrrhotite	1.28
118/21/15A	D zone	north pit BIF	#23 po-23	34S_Au4*105.adp	Pyrrhotite	0.89
118/21/15A	D zone	north pit BIF	#23 po-23	34S_Au4*112.adp	Pyrrhotite	0.66
118/21/15A	D zone	north pit BIF	#24 po-24	34S_Au4*106.adp	Pyrrhotite	0.06
118/21/15A	D zone	north pit BIF	#25 po-25	34S_Au4*107.adp	Pyrrhotite	0.49
118/21/15A	D zone	north pit BIF	#26 po-26	34S_Au4*108.adp	Pyrrhotite	0.58
118/21/15A	D zone	north pit BIF	#27 po-27	34S_Au4*109.adp	Pyrrhotite	0.41

Table A2. Results of the EPMA analysis of pyrite from the South Deep deposit.

Analysis Spot	Analysis Position	Fe (wt%)	Cu (wt%)	Co (wt%)	Ni (wt%)	As (wt%)	S (wt%)	Total (wt%)
11	core S11	45.85	-	0.36	0.08	0.04	53.60	99.93
12	rim S12	46.41	-	0.10	0.11	0.01	54.32	100.94
14	core S13	46.73	0.02	0.04	0.01	-	53.77	100.57
15	rim S14	46.15	-	0.09	0.08	0.05	54.01	100.37
5	core S15, S27	46.05	0.04	0.05	-	0.18	53.85	100.17
6	rim S16, S26	46.67	-	0.05	0.02	0.32	53.82	100.88
7	rim S16, S26	46.21	0.01	0.07	0.04	0.25	54.03	100.60
8	rim S16, S26	46.57	0.03	0.07	0.03	0.13	52.47	99.30
9	core S24	46.37	-	0.05	0.01	0.02	53.75	100.19
10	rim S25	46.33	-	0.10	0.03	-	52.60	99.06
3	Rim S28	46.99	0.01	0.08	-	-	51.87	98.95
4	Rim S28	46.32	-	0.11	0.02	0.02	53.95	100.42
1	core S29	45.62	-	0.05	-	-	54.70	100.37
2	core S29	46.54	-	0.07	0.02	0.00	52.72	99.35
13	core S30	45.99	-	0.22	-	-	53.74	99.95

Table A2. Cont.

Analysis Spot	Analysis Position	Fe (mol%)	Cu (mol%)	Co (mol%)	Ni (mol%)	As (mol%)	S (mol%)	Total (mol%)
11	core S11	32.83	-	0.25	0.06	0.02	66.85	100.00
12	rim S12	32.86	-	0.07	0.07	0.01	67.00	100.00
14	core S13	33.27	0.01	0.03	0.01	-	66.69	100.00
15	rim S14	32.86	-	0.06	0.05	0.02	67.00	100.00
5	core S15, S27	32.87	0.03	0.04	-	0.09	66.97	100.00
6	rim S16, S26	33.16	-	0.04	0.01	0.17	66.62	100.00
7	rim S16, S26	32.86	0.01	0.05	0.03	0.13	66.93	100.00
8	rim S16, S26	33.70	0.02	0.05	0.02	0.07	66.14	100.00
9	core S24	33.11	-	0.03	0.00	0.01	66.85	100.00
10	rim S25	33.56	-	0.07	0.02	-	66.36	100.00
3	Rim S28	34.18	0.00	0.05	-	-	65.74	99.98
4	Rim S28	32.98	-	0.07	0.01	0.01	66.92	99.99
1	core S29	32.36	-	0.04	-	-	67.60	100.00
2	core S29	33.61	-	0.05	0.01	0.00	66.33	100.00
13	core S30	32.89	-	0.15	-	-	66.95	99.99

core S11: Core of the grain, where it is close to S11 in Table 2.

Table A3. Results of the EPMA analysis of pyrite from the KDC West deposit.

Analysis Spot	Analysis Position	Fe (wt%)	Cu (wt%)	Co (wt%)	Ni (wt%)	As (wt%)	S (wt%)	Total (wt%)
1	rim D1	46.03	-	0.11	-	0.06	52.32	98.51
9	core SgD1	45.25	-	0.06	0.03	0.11	54.17	99.62
5	rim SgD3	45.62	-	0.09	0.11	0.12	53.78	99.71
4	rim D2	45.24	0.06	0.40	0.07	0.04	53.94	99.75
12	rim D2	46.25	-	0.29	0.13	0.02	53.17	99.86
13	rim D2	45.84	0.01	0.55	0.09	0.04	52.06	98.58
14	rim D2	46.23	-	0.44	0.09	0.02	53.46	100.24

Analysis Spot	Analysis Position	Fe (mol%)	Cu (mol%)	Co (mol%)	Ni (mol%)	As (mol%)	S (mol%)	Total (mol%)
1	rim D1	33.52	-	0.08	-	0.03	66.37	100.00
9	core SgD1	32.37	-	0.04	0.02	0.06	67.51	99.99
5	rim SgD3	32.68	-	0.06	0.07	0.06	67.12	100.00
4	rim D2	32.38	0.04	0.27	0.05	0.02	67.24	100.00
12	rim D2	33.20	-	0.20	0.09	0.01	66.50	100.00
13	rim D2	33.42	0.01	0.38	0.06	0.02	66.11	100.00
14	rim D2	33.05	-	0.30	0.06	0.01	66.58	100.00

rim D1: rim of the grain, where it is close to D1 in Table 2. core SgD1: core of the same grain as D1 grain.

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