

Article **Seasonal Variability of Near-Inertial Internal Waves in the Deep Central Part of the Black Sea**

Elizaveta Khimchenko 1,*, Alexander Ostrovskii ¹ [,](https://orcid.org/0000-0002-5906-0361) Alexey Klyuvitkin [1](https://orcid.org/0000-0002-8739-6321) and Leonid Piterbarg ²

- ¹ Shirshov Institute of Oceanology, Russian Academy of Sciences, 36 Nahimovskiy Prospect, 117997 Moscow, Russia; osasha@ocean.ru (A.O.); klyuvitkin@ocean.ru (A.K.)
- ² Department of Mathematics, University of Southern California, Kaprielian Hall, Room 108, 3620 Vermont Avenue, Los Angeles, CA 90089-2532, USA; piter@usc.edu
- ***** Correspondence: ekhym@ocean.ru

Abstract: This observational study is concerned with the seasonal variability of near-inertial internal waves (NIWs) in the central part of the Black Sea. Rotary spectral analysis of the nearly year-long time series of the sea current velocity measurements at 100 m and 1700 m revealed the prevailing anticyclonic component of the motions near the local inertial frequency *f*. Both the rotary spectra and the visual exploration of the time series showed that the peaks of NIWs were blue-shifted to higher frequencies. The monthly average blue-shift was stronger up to 1.038*f* in the summer. It was found that the minimum intensification of the NIWs occurred in summertime and the maximum intensification was characteristic of the autumn-winter period when the NIW packets included up to 16 waves with pronounced clockwise rotation of the velocity vectors.

Keywords: internal waves; near-inertial motions; currents; Black Sea; central basin

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1. Introduction

Ocean internal gravity waves propagate through the stratified interior of the ocean. As waves propagate, they interact with each other, producing an internal gravity wave continuum consisting of energy in many frequencies. The propagation of NIW energy can lead to mixing due to wave breaking in the thermocline [\[1](#page-18-0)[,2\]](#page-18-1). It is well known that the average annual contribution of inertial internal waves (IWs) to the global energy flux in the upper mixed layer in the ocean is comparable to that of internal tides [\[3\]](#page-18-2). Among possible generation mechanisms of inertial internal waves are wind forcing [\[4](#page-18-3)[,5\]](#page-18-4) and nonlinear interactions of IWs (which likely leads to the generation of inertial waves of higher modes), the formation of lee waves by geostrophic flow over the seabed topography, and generation due to the interaction of mesoscale and submesoscale motions [\[1\]](#page-18-0). An analytical solution for inertial internal waves in the boundary layers was given in [\[6\]](#page-18-5). In particular, the boundary layers were shown to be dependent on the horizontal extension of the wind forcing: the smaller the scale is, the sharper the boundary layer and the stronger the localization of inertial oscillations at the Ekman layer boundary. For inertial oscillations that propagate over regions of a horizontal size of baroclinic Rossby radius and larger, the boundary layers are significantly thicker and develop more slowly, and the fluctuations can reach the bottom.

Using an extensive moored current-meter database it was shown that the winter enhancement by a factor of 4–5 in the Northern Hemisphere for latitudes 25◦–45◦ at all depths less than 4500 m is in good agreement with the seasonal cycle of wind-forced mixed-layer energy flux [\[7\]](#page-18-6). The IWs were typically observed in short packets, so it was often difficult to define the wave frequency with a high degree of accuracy [\[1\]](#page-18-0). Below we study the seasonal change of IW during almost year-long observations using the current meter mooring data in the Black Sea.

The semi-enclosed Black Sea is a bowl-shaped and deep (>2000 m) basin, where inertial motions dominate the dynamics at time scale < 1 day throughout the water column [\[8\]](#page-18-7). In the Black Sea, the effect of near-inertial internal waves on the generation of turbulence and vertical mixing in the pycnocline is of a great importance [\[9](#page-18-8)[,10\]](#page-18-9). The Black Sea tides are so small [\[11\]](#page-18-10) that the tidal waves do not interfere with IWs. The significance of IWs in hydrophysical studies of the Black Sea is primarily determined by the role as an energy source for internal waves in a higher frequency range, which is important for mixing.

Instrumental observations of inertial motions in the Black Sea have been conducted since the 1960s [\[12–](#page-18-11)[14\]](#page-19-0). Measurements of the IWs in the Black Sea were carried out using moorings equipped with temperature sensors and acoustic Doppler current profilers [\[15](#page-19-1)[–17\]](#page-19-2). Observations were often conducted in the sea coastal zone during the summer-autumn season, although some data were obtained under winter stratification conditions. Measurements were rare at the continental slope due to difficulties with setting up the oceanographic instruments in the strong boundary current named the Rim Current (for the Black Sea circulation, see observational and modeling studies [\[18–](#page-19-3)[21\]](#page-19-4)). Recently, the IWs of the second mode were recorded [\[15\]](#page-19-1). These and other studies [\[22,](#page-19-5)[23\]](#page-19-6) were also concerned with the manifestation of abnormal inertial motions in the Black Sea, such as cyclonic rotation of hodographs of inertial currents, similar to those observed on the ocean shelf in the Northern Hemisphere [\[24\]](#page-19-7). Over the Black Sea, the IWs were also observed by remote sensing methods using radars, spectrometers, and radiometers onboard of Envisat, ASAR, ERS-2 SAR, and Landsat-5, 7, and 8 satellites [\[25,](#page-19-8)[26\]](#page-19-9).

The inertial motions can be identified in the frequency spectra by a peak close to the local inertial frequency *f* [\[27,](#page-19-10)[28\]](#page-19-11). Due to the shift relative to *f*, IWs are called nearinertial waves (NIWs) [\[29\]](#page-19-12). In the Black Sea, the frequency of near-inertial motions' peak in the energy spectrum was found to be 4–5% higher than *f* [\[8\]](#page-18-7), which is a so-called blueshift of frequency [\[30\]](#page-19-13). In the Black Sea the magnitude of this shift was consistent with theoretical and experimental values obtained in other regions of the World Ocean [\[31,](#page-19-14)[32\]](#page-19-15). In addition, the blue-shift increased with depth, which did not contradict the theory. For the Black Sea, the share of NIWs energy in the total kinetic energy increases from 10% at a depth of 50 m to 30% at 350 m [\[16\]](#page-19-16). This estimation was obtained for the averaged profile of the vertical stratification observed in the autumn, when the buoyancy frequency maxima were located in the seasonal pycnocline at a depth of 35 m, in the cold intermediate layer (CIL) at 50 m, and in the main pycnocline at 105 m [\[33](#page-19-17)[,34\]](#page-19-18). It should be noted that the seasonal thermocline isolates the CIL and the underlying water masses from atmospheric impact in the warm period of the year. Above the core of the CIL, there is a stable stratification both in temperature and salinity; deeper, the stratification is unstable in temperature and stable in salinity.

Although there have been collected various data about inertial motions in the Black Sea, most of the observations were limited to the coastal areas and the summer season. For the central part of the sea, the observations were of short duration and for the summer–autumn period only. The data analyzed in this study is based on frequent nearly year-long measurements in the CIL layer (the upper part of the main pycnocline) and in the deep layer (approximately 100 m above the sea bottom). Our study region is suitable for NIW observations for several reasons: the basin is microtidal, the study point is far from the continental shelf and coast, where the bottom is almost flat, and there is no sea breeze activity. Specific examples of the manifestations of NIWs are considered to compare the variability in different seasons. Additionally, an analysis of near-inertial fluctuations is presented for the annual series of sea temperature at 265 m depth in the northeastern part of the deep basin of the Black Sea (see Appendix [B\)](#page-17-0).

2. Materials and Methods

The anchored buoy station (ABS) was deployed in the northwestern part of the central basin of the Black Sea in 2016–2017. The mooring (ABS No. 1 in Figure [1\)](#page-2-0) was deployed at ~1800 m depth at $43°47'$ N 32°04' E. The top buoy of the mooring was located at a depth of 50–60 m in order to avoid the effect of waves and strong near-surface currents. The mooring was equipped with the acoustic Doppler current meters at fixed depths as
6. We also can interval in the interval instrumental instrumental instrumental instrumental instrumental instr follows: a Sontek Argonaut at 100 m and a Nortek Aquadopp at 1700 m. We analyzed Sontek data from 1 December 2016 to 5 October 2017 and Nortek data from 1 December 2016 to 30 November 2017. The temperature was recorded at 100 m. The measurement interval was 15 min and the internal instrumental averaging was 5 min. Noticeably, the measurement interval of 15 min is 0.0145 part of the local inertial period at the ABS measurement interval of 15 mm is 0.0145 part of the local mertial period at the ABS
No. 1 mooring site. Additionally, to get the time series of the sea temperature, the northeastern station (ABS No. 2) was deployed at the 2160 m depth at $43°26'$ N $36°22'$ E from 18 December 2017 to 4 December 2018 (results are discussed in the Appendix [B\)](#page-17-0). It was equipped with RBR Duet instrument fixed at a depth of 265 m. The sampling interval for the temperature was set to 30 s.

according to ETOPO1 data doi:10.7289/V5C8276M. 1—is for ABS No. 1, 2—is for ABS No. 2. **Figure 1.** The anchored buoy stations (ABS) locations in 2016–2018. Bathymetry of the Black Sea

The de-trended time series of the current velocity were processed using rotary specoceanography since the 1970s [35,36] in particular, in applications to tidal [37] and in[er](#page-19-19)tial currents [38]. The method is based on a representation of the horizontal velocity vector as a complex number: $\mathbf{a}(t) = \mathbf{a}(t) + i\mathbf{a}(t)$ (3) (3) tral analysis, or the method of rotational components, which has become widespread in complex number:

$$
\mathbf{u}(t) = u(t) + i v(t),\tag{1}
$$

 $\mathbb{Z}(\mathbf{y}) = \mathbb{Z}(\mathbf{y})$. The method is based on a representation of the horizontal velocity \mathbf{y} where t is time, u and v are zonal and increasonal components, and $t = \sqrt{t}$. The velocity vector can be decomposed for each frequency into two oppositely rotating motions, each with its amplitude (A^+, A^-) and phase (θ^+, θ^-) . Counterclockwise rotations correspond
to movements with positive frequencies, and electivise rotations correspond to positive ones. The components of the two-dimensional velocity vector at the inertial frequency
make periodic elliptic erbits in the bedeemaph plane. Frequency enectro are determined for both the negative and positive frequencies. The total energy spectrum *S*(ω), where ω where t is time, u and v are zonal and meridional components, and $i =$ √ −1. The velocity to movements with positive frequencies, and clockwise rotations correspond to negative make periodic elliptic orbits in the hodograph plane. Energy spectra are determined for

is the frequency, can be represented as the sum of the clockwise and counterclockwise components of the spectra ranges in the following form:

$$
S(\omega) = S^-(\omega) + S^+(\omega) \tag{2}
$$

As follows from the classic theory (for example, [\[28,](#page-19-11)[35\]](#page-19-19)) for the Northern Hemisphere, inertial oscillations appear in the clockwise component. One of the characteristics of a random vector process is the rotary coefficient $r(\omega)$, proposed in [\[35\]](#page-19-19):

$$
r(\omega) = \frac{S^-(\omega) - S^+(\omega)}{S^-(\omega) + S^+(\omega)}.
$$
\n(3)

For clockwise movements, the rotary coefficient is positive $r(\omega) \geq 0$ and for inertial oscillations $r(\omega) = 1$; in this case, the rotation of the velocity vector is observed as a circular orbit. When $r(\omega) \neq 1$, the orbit approaches an ellipse.

For the rotary spectral analysis we used the software package jLab v.1.7.1 [\[39\]](#page-19-23) [http://www.jmlilly.net/software,](http://www.jmlilly.net/software) accessed on 1 March 2022. The availability of nearly year-long measurements of the currents with a 15-min sampling rate makes it possible to carry out a frequency spectral analysis that is statistically rigorous and reveals the seasonal variability of the observed NIWs. At the first stage of data processing, the energy spectra of the clockwise (anticyclonic) and counterclockwise (cyclonic) components of the current velocity were estimated for the entire observation period. Then, the rotary energy spectra were calculated for each month, and the annual variability of the NIWs was analyzed.

To assess the thermohaline stratification conditions near the ABS No. 1, the vertical profiles of the density and buoyancy frequency (Brunt–Väisäla) were calculated from the temperature and salinity data of the Argo floats (Nos. 6901832, 6901833, and 6901834) [\[40\]](#page-19-24) for 2017. Notice that the Argo floats are the primary research tools for contemporary oceanography of the Black Sea [\[41](#page-19-25)[–46\]](#page-20-0).

To assess the atmospheric forcing condition, the correlation analysis between the surface wind vector **U** and the current vector **u** was performed. The wind data were taken from the ERA-5 meteorological reanalysis data, which combine model data and observations on a grid with 0.25° latitude and longitude spacing. The hourly data of the near-surface wind at the height of 10 m from the sea surface at a grid point close to the ABS No. 1 site were used. To find the correlations between the time series $U(t)$ and $u(t)$, the complex correlation method proposed by [\[47\]](#page-20-1) was used. The complex correlation coefficient was defined as the normalized inner product of two vector time series:

$$
\rho_0 = \rho \exp(i\varphi) = \frac{\langle \mathbf{U}(t)\mathbf{u}^*(t) \rangle}{\langle \mathbf{U}(t)\mathbf{U}^*(t) \rangle^{1/2} \langle \mathbf{u}(t)\mathbf{u}^*(t) \rangle^{1/2}},\tag{4}
$$

where the asterisk (*) is the complex conjugate operator, $\langle \ \rangle$ denotes the ensemble-average operator, *ρ* is the the magnitude of the correlation, and *ϕ* gives the mean counterclockwise rotation angle of the current velocity vector relative to the wind velocity vector.

To estimate the correlation coefficient *ρ^τ* with a time lag (*τ*), we have:

$$
\rho_{\tau} = \frac{\langle \mathbf{U}(t)\mathbf{u}^*(t+\tau) \rangle}{\langle \mathbf{U}(t)\mathbf{U}^*(t)\rangle^{1/2} \langle \mathbf{u}(t+\tau)\mathbf{u}^*(t+\tau) \rangle^{1/2}}.
$$
(5)

Before calculating the correlations, the time series $U(t)$ and $u(t)$ were de-trended. Since the reanalysis wind data included diurnal breeze signal (which is somewhat atypical for a point in the central part of the sea), the wind data were filtered. The fourth-order Butterworth bandpass filter was applied to extract near diurnal velocities. The cutoff frequency band was [0.916, 1.083] *fdiurnal*, where the diurnal wind frequency *fdiurnal* = 0.0416 cph. For the correlation calculation, time windows of 96–100 h and lags $\tau \leq 48$ h were selected.

The estimates of the NIW period based on the results of spectral analysis of relatively short (1 month) time series may possess large uncertainty. This could be due to the phase
and delivered NIW packets and his proposition at the time interval which wish to seem modulation of NIWs generated by successive atmospheric storms, which might occur between the creation of two adjacent waves of the creative waves of the velocity occasionally. Thereby, the total number of NIW crests in the time series of the velocity components for the identified NIW packets for each month was calculated manually. To do this, we visually identified the NIW packets and estimated the time interval between the crests of two adjacent waves in the packet. The average period of waves was determined for each month as the arithmetic means of the sum of the average periods of identified packets. Taking into account the measurement interval of 15 min, the error of the individual estimate of the distance between two inertial waves was 1.45% of the local inertial period. The NIW intermittency was calculated as the ratio of the duration of the identified NIWs to **3. Results** the total observation time for each month.

3. Results

3.1. Vertical Stratification and Near-Inertial Currents The vertical stratification was obtained from the Argo float data available in the Argo fl

The vertical stratification was obtained from the Argo float data available in the observation region for 2017 (Figure [2\)](#page-4-0). In the warm season, the vertical stratification was characterized by seasonal and main pycnoclines. The upper quasi-homogeneous layer Inducting by seasonal and main pycnoclines. The apper quasi nonogeneous layer
began to form in April due to the warming of surface waters and could be observed until october. The seasonal pycnocline was sharpest in June at depths of 30–40 m. In autumn, the position of its lower boundary gradually shifted to a depth of 50 m. Two maxima characterized the buoyancy frequency profile: subsurface $N = 23$ cph and deep maximum *N* > 10 cph in the main pycnocline that was located usually below 80 m.

Figure 2. Vertical profiles of the temperature *T*, potential density σ_{θ} , and the Brunt-Väisäla frequency quency near to ABS No. 1 for 11 January, 13 April, 13 June, and 11 October 2017, based on data *N* near to ABS No. 1 for 11 January, 13 April, 13 June, and 11 October 2017, based on data from Argo floats Nos. 6901832, 6901833, and 6901834. Profiles are given to a depth of 500 m to show the changes in the upper layer of the sea in more detail.

was $\langle |u| \rangle = 0.087 \text{ ms}^{-1}$, the standard deviation was $\sigma_u = 0.04 \text{ ms}^{-1}$, and the current was oriented mainly to the southwest (207°) . Seasonal changes in the current velocity were observed: the highest values were noted in winter, mainly in December (maximum of 0.28 ms⁻¹), and the minimum occurred in summer. At this depth, the temperature varied from 8.35 °C (in December) to 8.5 °C (in March and May) at a depth of 100 m (Figure 3). The According to ABS No. 1 observations at a depth of 100 m, the average current velocity

100 m depth at ABS No. 1 from 1 December 2016 to 5 October 2017. The graphs are based on data 100 m depth at ABS No. 1 μ at ABS No. 1 μ December 2017. The graphs are based on data b **Figure 3.** Time series of current speed (**top**) and direction (**middle**) and temperature (**bottom**) at the averaged over 1 h.

During winter months, mainly in January–February, strong inertial motions were distinguished by the rotation of the vector **u** with near-inertial frequency (Figure [4\)](#page-6-0). At 100 m depth, the inertial orbital motion was of 0.5–1.5 km diameter. At 1700 m depth, the hodograph orbits were not closed. Instead, the hodographs were arc-shaped with a typical

horizontal scale of 0.5–0.7 km. Note that the baroclinic Rossby radius for the deep-water part of the Black Sea could be 15–25 km [\[48\]](#page-20-2).

00:00 20 January 2017 to 03:00 24 January 2017. The data were averaged over 1 h. The horizontal and vertical axes show the distances in longitude and latitude from the starting points of the data ensembles. The distances in longitude and latitude and latitude from the starting points of the data ensembles. **Figure 4.** Progressive vector diagram of the currents at 100 m and 1700 m depths of ABS No. 1 from *3.2. Rotary Spectra of the Currents*

3.2. Rotary Spectra of the Currents

the power spectra, a peak close to the local meritar nequency $y = 0.056$ Cpm, and
clockwise rotation indicate the frequent presence of NIW at 100 m in the central basin of the Black Sea (Figure 5a). At frequencies [cl](#page-6-1)ose to the local inertial one, the value of the rotary coefficient approached 1, which indicates that clockwise rotation dominated at the inertial frequency, and the velocity vector made an ellipse-like turn (Figure [5b](#page-6-1)). In the power spectra, a peak close to the local inertial frequency $(f = 0.058 \text{ cph})$ and

Figure 5. (a) Rotary spectra of the clockwise and counterclockwise current vector rotation motions at at 100 m at ABS No. 1. In the left plot, the vertical dotted line indicates the local inertial frequency 100 m at ABS No. 1. In the left plot, the vertical dotted line indicates the local inertial frequency *f* at at the mooring site. (**b**) Rotary coefficient in the range of near-inertial frequencies. the mooring site. (**b**) Rotary coefficient in the range of near-inertial frequencies.

The clockwise power spectra of the current velocity at 100 m for each month are shown in Figure [6.](#page-7-0) A distinct peak was found at a frequency of 0.058–0.061 cph (in the range In rigule 6. A distinct peak was found at a flequency of 0.056–0.061 cpn (in the range
of periods 17.24–16.39 h) every month. At 100 m depth, the energy peak was highest in February and lowest in July (inset in Figure 6). The ratio of the peaks in February and in July was about 8.5.

The dashed line indicates the local inertial frequency in the study area. The small image in the upper right corner is an enlargement of the figure to better visualize the peaks close to the inertial periods. **Figure 6.** The rotary spectra of clockwise rotation motion at 100 m at ABS No. 1 in different months.

periods. In the deep layer at 1700 m, there was also a significant peak near the inertial frequency (Figure [7\)](#page-8-0), and the anticyclonic velocity component prevailed over the cyclonic. The NIWs were most energetic in January and weakest in June. The ratio between the local spectral peak maximum and the local spectral peak minimum was as high as 14.0. A comparison of the spectra at 100 m and 1700 m shows that the energy level in the upper layer was an order of magnitude higher than that in the deep layer. In addition, the smaller maximum at the superinertial frequency 2*f* was observed in winter (December–February), spring (March–May), and September.

PSD [(cm/s)²cph]

 $10[°]$ 10^{-3} 10^{-2} 10^{-1} $10⁰$ ω [cph]

The dashed line indicates the local inertial frequency in the study area. The small image in the upper right corner is an enlargement of the figure to better visualize the peaks close to the inertial periods. the upper right corner is an enlargement of the figure to better visualize the peaks close to the peaks close **Figure 7.** The rotary spectra of clockwise rotation motion at 1700 m at ABS No. 1 in different months.

3.3. Seasonal Variation of Near-Inertial Waves

3.3.1. Winter Season

At the beginning of the winter season, the intensification of NIWs alternated with $\frac{1}{2}$ tions weakened, and then modulation occurred, possibly caused by a sharp change in the near-surface wind direction. A time shift of 36–40 h between the wind action and the response was noted (Figure 8b)[. F](#page-9-0)rom 8 to 13 December, the NIWs decreased, due to a stabilizing effect of the southwestern wind blowing for more than 100 h (Figure [8a](#page-9-0)). The
suggest area of uses as duesd although the hade susphere this possible the inestial fraction and earent speed was reduced, alaloagh the housgappis roducent with the filerular requestly persisted. The next packet of near-inertial motions developed again after 13 December and continued until 20 December (Figure [8a](#page-9-0)). During these episodes, 10 crests of inertial waves passed. This wave packet was preceded by increased speed of the southeastern wind up to 8 m s^{−1} on 12–13 December. For this observation ensemble, the maximum correlation between wind and current vectors was $\rho_\tau = 0.53$ at the time lag $\tau =$ 45–48 h. On 21 December, the current direction changed. Then the new packet of regular and intense near-inertial the current direction changed. Then the new packet of regular and intense near-inertial becomments of an increased period arrived. In the first wave packet in the second part of December, the period was about 15.81 h. The wave period extended to 17.73 h in the second wave packet and then turned to 16.94 h in the third packet. $\frac{1}{1}$ periods of calm at a 100 m depth. Strong intermittence or enhanced temporal variability of wave packets was observed (Figure [8a](#page-9-0)). From 1 to 4 December, near-inertial oscillacurrent speed was reduced, although the hodographs rotation with the inertial frequency oscillations of an increased period arrived. In the first wave packet in the second part of

Figure 8. (a) Top—wind speed $U(t)$ vector plot from the ERA-5 meteorological reanalysis data. The second plot from the top is the current velocity $\mathbf{u}(t)$ at 100 m at ABS No. 1 in December 2016, and the third plot from the top shows the u - and v -components of the current velocity. Bottom—wind speed after band-pass filtering for 22–26 h. (**b**) Top—lagged correlations ρ_{τ} between the wind and speed after band-pass filtering for 22–26 h. (**b**) Top—lagged correlations ρ_{τ} between the wind and $\frac{1}{2}$ speculate vand plass find the $\frac{1}{2}$ december 2016, both $\frac{1}{2}$ december 2016, both $\frac{1}{2}$ december 2016, both $\frac{1}{2}$ decreases for $\frac{1}{2}$ current vectors for 1–4, 5–8, and 13–16 December 2016, bottom—the angle φ of rotation of the current velocity vector relative to the wind velocity vector.

An intensification of NIWs was observed on 6-10 January 2017 (Appendix [A,](#page-16-0) A1). This amplification was preceded by a sharp increase in the north wind speed. An-Figure [A1\)](#page-16-1). This amplification was preceded by a sharp increase in the north wind speed. other intense NIW packet was observed on 11–16 January, when five waves were rec-Another intense NIW packet was observed on 11–16 January, when five waves were recorded. Starting from 17 January, a well-defined sequence of 13 waves with a period of 16.79 h arrived. The wave amplitude gradually decreased by 25 January. The correlamaxima between the wind and current velocity vectors for 12–16 January and 20–24 tion maxima between the wind and current velocity vectors for 12–16 January and 20–24 January were $\rho_{\tau} = 0.43$ and $\rho_{\tau} = 0.44$, with time lags of 30–37 and 30–43 h, respectively. The rotation angle between the current velocity and wind vector φ varied from $-165°-170°$ to $160^{\circ} - 180^{\circ}$.

As noted above, persistent long-term forcing by intense wind could suppress in-ertial oscillations, such as those observed in February 2017 (Figure [9\)](#page-10-0). There were no inertial oscillations during the period of the northeastern wind on 8-14 February. The intensification of inertial fluctuations happened again after 14 February and could be observed for almost 14 days. These oscillations could originate due to changes in the strength and direction of the wind. The packet of 15 inertial internal waves was recorded during this period, reaching the maximum amplitude on 18-19 February. During the passage of the NIW packets, positive correlations were noted between the wind and current

velocity vector at $\tau = 31$ –36 h. Notably, NIWs were developed during the prevailing southeasterly. The intensification of NIWs peaked after 5–6 periods, and then a slow decay followed (Figure 9). The inertial fluctuations dissipated after the wind direction changed to followed (Figure 9). [Th](#page-10-0)e inertial fluctuations dissipated after the wind direction changed the northeastern.

Figure 9. (a) Top—wind speed $U(t)$ vector plot from the ERA-5 meteorological reanalysis data. The second plot from the top is the current velocity $\mathbf{u}(t)$ at 100 m at ABS No. 1 in February 2017, and the third $\frac{1}{2}$ the third plot the third plot from the α -components of the current velocity. Botthird plot from the top shows the *u*- and *v*-components of the current velocity. Bottom—wind speed after band-pass filtering for 22–26 h. (**b**) Top—lagged correlations ρ_{τ} between the wind and current vectors for 16–20 and 21–25 February 2017, bottom—the angle φ of rotation of the current velocity vector relative to the wind velocity vector.

3.3.2. Spring–Summer Period of 2017, the NIW energy was getting lower than that in α

In the spring–summer period of 2017, the NIW energy was getting lower than that in the winter season (Figures 6 and 7). NIWs were observed more clearly in the current speed data, and the velocity hodographs underwent significant changes. In March, there were certain periods of intensification of inertial oscillations with a maximum after 27 March (Figure [A2\)](#page-16-2). The wind was usually weak 3 m s^{−1} or less increasing to 8 m s^{−1} on some days. The most intense NIWs were observed just after such an intensification of the wind forcing. The most method is the west cose were just and such The maximum correlation between the surface wind and current vectors was relatively low $\rho_{\tau} = 0.44$ with a time lag $\tau = 31-45$ h.

In the summer conditions of June 2017 (Figure [10\)](#page-11-0), the NIW energy at 100 m depth was at the annual minimum. Four short packets of NIWs appeared during this month. Each event consisted of 3–7 oscillations. At the beginning of June, the strong NIWs $(|u| \approx 0.2 \text{ m s}^{-1})$ were of a period 16.33 h. Later the current speed in the NIWs diminished to 0.05–0.10 m s^{−1}. In June, the values of ρ_{τ} were insignificant; only on 16–20 June, ρ_{τ} = 0.44 at τ = 38–48 h (Figure [10b](#page-11-0)).

cated the main direction of the water mass transport. The rotation of the hodographs was

Figure 10. (a) Top—wind speed $U(t)$ vector plot from the ERA-5 meteorological reanalysis data. The $\frac{1}{2}$ second plot from the top is the current velocity $u(t)$ at 100 m at ABS No. 1 in June 2017, and $\frac{1}{2}$ second plot from the top is the current velocity $\mathbf{u}(t)$ at 100 m at ABS No. 1 in June 2017, and the third plot from the top shows the *u*- and *v*-components of the current velocity. Bottom—wind speed after band-pass filtering for 22–26 h. (**b**) Top—lagged correlations ρ_{τ} between the wind and current vectors for 16–20 June 2017, bottom—the angle *φ* of rotation of the current velocity vector relative to the wind velocity vector.

In the warm season as compared with winter conditions, the orbits of near-inertial motions significantly changed (Figure [11\)](#page-11-1). The hodographs did not close and had the shape of an arc elongated to the southwest in March and northwest in June, which indicated the main direction of the water mass transport. The rotation of the hodographs was insignificant at the 1700 m depth.

Figure 11. Progressive vector diagrams of currents at 100 m depth for selected periods in March **Figure 11.** Progressive vector diagrams of currents at 100 m depth for selected periods in March (**left**) (**left**) and June (**right**) of 2017. and June (**right**) of 2017.

3.3.3. Autumn Season

As shown in above Figures [6](#page-7-0) and [7,](#page-8-0) the energy level at the inertial frequency in September was close to the January values. The inertial oscillations of the velocity components underwent temporal variations in September (Figure [12\)](#page-12-0). Long packets of 3–13 inertial cycles were observed. In the longer packets, the intensification happened at the 4th–6th oscillation. The maximum amplitudes of the current velocity fluctuations reached 0.2 m s⁻¹. Then the NIWs slowly dissipated. The results of the correlation analysis between the wind and current velocity vectors show significant positive correlations maximum $\rho_{\tau} = 0.52$ for the NIW packets during 1–4 and 16–20 September (Figure [12b](#page-12-0)) at the time lags *τ* of 25 h and 28 h, respectively. In September, meteorological conditions were characterized by a frequent change in the wind direction with a frequency close to inertial (Figure [12a](#page-12-0)).

Figure 12. (a) Top—wind speed $U(t)$ vector plot from the ERA-5 meteorological reanalysis data. The second plot from the top is the current velocity $\mathbf{u}(t)$ at 100 m at ABS No. 1 in September 2017, and the third plot from the top shows the *u*- and *v*-components of the current velocity. Bottom—wind speed after band-pass filtering for 22–26 h. (**b**) Top—lagged correlations ρ_{τ} between the wind and current vectors for 1–4 and 16–20 September 2017, bottom—the angle φ of rotation of the current velocity vector relative to the wind velocity vector.

3.4. The Seasonal Variability of the Blue-Shift in the Near-Inertial Frequency

The spectral analysis revealed the blue-shift frequency of NIWs (see Section [3.2](#page-6-2) above). This was confirmed by visual identification of the NIW periods, which showed that the average periods were mainly shorter than the inertial ones (Table [1\)](#page-13-0). The shortest periods varied from 14.8 h in January to 15.3 h in August, while the largest varied more broadly from 19.25 h to 20.8 h in April. The monthly average periods were almost the same as the local inertial period in March and April. In May and June, the observed NIW's periods became shorter ~16.6 h. In July and September, these were nearly 16.8 h and later in December, they shortened again to 16.6 h. In terms of frequency, the monthly average shift was in the range from 1.002*f* to 1.038*f*, i.e., 0.2–3.8%. The high temporal intermittency was a remarkable feature of NIWs. The modulation time scale increased in winter (~11.3 days in December, 10.5 days in February) and shortened in the spring–summer period (1.4 days in May, 1.3 days in August). The maximum number of NIWs was observed in December and September (39 and 34 fluctuations, respectively), and the minimum, in June and July (21 and 25 fluctuations, respectively). The intermittency was at maximum of 87% in December and at minimum of 48% in June.

Table 1. Near-inertial waves at 100 m depth for individual months, 2016–2017.

The thermohaline stratification was modified in spring and autumn. These transformations occurred with delay from the astronomical seasons. For example, the stratification was not yet of a winter type in December. It was established in January and persisted until the end of April. During this period, the modulation time scales of NIWs were the longest. The seasonal pycnocline began to develop in May, and the strongest shifts of near-inertial frequency and shortest modulation time scales were observed. Also from the spectral analysis, it was found that during transitional seasons, the power spectra peaks were also found at frequency ~2*f*. The overtones of inertial oscillation could be associated with the observed shortness of the NIW packets.

4. Discussion

The rotary spectral analysis of the velocity vectors revealed a peak at the inertial frequency with clockwise rotation. These features indicate the predominant influence of inertial oscillations in the study area. Similar results for the deep-water part of the Black Sea were obtained previously based on a short time series of observation data [\[14\]](#page-19-0). The power spectra maxima were basically shifted toward frequencies higher than the local inertial frequency. The analysis of the current measurement data in the Black Sea showed a frequency shift of 1.002*f* –1.038*f* (0.2–3.8%). The power spectra analyses indicated relatively broader peaks. An apparent reason for this was the phase modulation of NIWs. However, both the spectral estimates and ad-hoc intuitive analysis showed the tendency of increase

in the NIW's period during the most stable vertical stratification in the sea. The blue-shift of the inertial motions was often observed in the World Ocean [\[29\]](#page-19-12). The blue-shift was observed in many areas of the World Ocean, for example, in the South China Sea [\[32\]](#page-19-15), where the meandering of the Kuroshio Current caused this shift. A frequency shift of NIWs to the low-frequency band ("red-shift") was also noted, which could be associated with the Doppler shift in the flow [\[49\]](#page-20-3) or by negative background vorticity, such as mesoscale eddy [\[49,](#page-20-3)[50\]](#page-20-4).

Based on our data analysis, the peaks of the power spectra of NIWs were more pronounced in September–February. Noteworthy that according to the results of global observations at moored stations, the energy of NIWs in the Northern Hemisphere middle latitudes also varies with season, with a maximum in winter [\[7\]](#page-18-6).

There could be several reasons for the blue-shift in the NIW frequency in the Black Sea study site. The seasonal variation in the wind stress over the Black Sea could be characterized by an increase in cyclonic vorticity in winter, which peaked in February [\[51\]](#page-20-5), while in summer, the atmospheric forcing change most likely resulted in the anticyclonic vorticity of the wind stress. The seasonal variation of atmospheric forcing contributed to the intensification of the Rim Current (cyclonic vorticity) in winter and its weakening in summer. In turn, the Rim Current instability could serve as a source of generation of NIWs [\[14\]](#page-19-0). The horizontal vorticity of the background currents affected inertial oscillations [\[52\]](#page-20-6). If the background vorticity was cyclonic (anticyclonic), then the frequency of inertial oscillations increased (decreased). Previous observations showed a strong impact of mesoscale vorticity on the NIW propagation in the ocean [\[53,](#page-20-7)[54\]](#page-20-8). The observed variation in the blue-shift magnitude could be due to the change of the background flow and the development of mesoscale anticyclones [\[55\]](#page-20-9). In particular, the Rim Current weakens during the spring–summer period [\[55,](#page-20-9)[56\]](#page-20-10). The seasonal change of the Rim Current was shown by the Black Sea Forecasting System ocean model, which assimilated near real time, in-situ and satellite observational products [\[57\]](#page-20-11). The seasonal variation of the mesoscale eddies was discussed in [\[55](#page-20-9)[,58\]](#page-20-12), in particular, the maximum anticyclone eddies were shown to appear often during summer in the Black Sea.

Analysis of the velocity data at 100 m depth and the near-surface wind suggested that the development of inertial motions was prompted by an increase in the wind speed magnitude over the study area, regardless of its direction. Furthermore, at 100 m depth, the NIW packets occur with a time lag of 25–45 h with respect to the wind forcing. Significant correlations were usually obtained for the winter and autumn seasons when the winds were much stronger. In the spring–summer period, there were practically no stable correlations (except for a few events). This could be associated with the formation of a barrier layer involving a seasonal pycnocline above the main pycnocline [\[14](#page-19-0)[,33\]](#page-19-17). Perhaps, more intense inertial motions could develop at shallower depths in the seasonal pycnocline in the summer, similar to those reported for the Gulf of Lion in the Mediterranean Sea [\[59\]](#page-20-13). However, the wind reanalysis sometimes did not accurately represent actual wind over the deep part of the basin.

In analytical solutions, Ref. [\[60\]](#page-20-14) noticed that the amplitude of inertial motions increases during the passage of atmospheric storm and later slowly decreases. Sometimes the storm results in rapid decay of the inertial oscillations that existed before it. Recently, similar results were obtained using the general ocean circulation models [\[61\]](#page-20-15). The inertial motions were generated as a sea response to a sharp change in wind speed. Under the steady wind, the ocean current speed in the inertial motions gradually weakened [\[60](#page-20-14)[,61\]](#page-20-15). However, studies only dealt with variability at the 30 m depth and argued that inertial motions decay with the depth and are unlikely to appear at greater depths [\[61\]](#page-20-15). Our observations demonstrated strong near-inertial oscillations at a depth of 100 m. There was also a correlation between the direction of the wind and the generation of the NIWs. Based on the data analysis, Ref. [\[62\]](#page-20-16) argued that high-frequency fluctuations in the wind speed enhance inertial oscillations of the currents in the upper ocean layer. According to our data, such intensification occurred in September when the sea current velocity rotated in accord with

leads to resonant excitation of inertial oscillations of ocean currents, e.g., the generation of inertial motions, as discussed in [\[63\]](#page-20-17). In particular, the highest resonant response to the effect of wind occurs when the wind rotates along with inertial oscillations. In this case, the frequency of the wind action must be close to the local effective Coriolis frequency to generate the near-inertial motions.

Unfortunately, there is not enough data to quantitatively relate the moments of generation of inertial internal waves packets at a depth of 100 m (and even more so at 265 m) with the variability of the near-surface wind. It can be suggested that the observed waves generated owing to the wind intensification propagate throughout the entire water column. These are most likely vertical standing waves, similar to those found previously in the Black Sea [\[13\]](#page-19-26). The inertial motion generation could be a result of a combination of forcing factors, such as the wind strength, the mean current speed, and the ocean eddy activity at the observation site. As noted by [\[64\]](#page-20-18), the variety of mechanisms of generation and levels of excitation of IWs make any attempts to interpret the properties of the spectrum of oceanic IWs within the assumption of a single mechanism meaningless. In addition, it is not easy to substantiate the possible generation of the observed internal waves with available data on the deep-water depth only. It is necessary to have data on the sea current speed and temperature throughout the water column, which requires particular experimental study.

5. Conclusions

The observations from 2016–2017 analyzed above were unique, since they were based on more than 10-month-long measurements with a high sampling rate in the CIL layer of the main pycnocline and at a depth of 1700 m, i.e., 100 m above the sea bottom. This study made it possible to supplement the existing knowledge about NIWs in the Black Sea and reveal their seasonal variability in its deep-water part.

The analysis showed that NIWs dominate over the internal-wave continuum of 0.01 cph $\lt \omega \lt 2$ cph in the Black Sea. The horizontal scales of the inertial rotation can be as large as 0.5–1.5 km and 0.5–0.7 km at 100 m and 1700 m depths, respectively. During the autumn–winter period, pronounced rotation of the current velocity vectors was observed. The hodographs appeared as ellipses with closed orbits. In the spring, inertial motions became infrequent and weak so that the hodographs of inertial currents strongly elongated in a slow current direction of subinertial time scales. An analysis of wind activity and the intensification of inertial motions suggested that each period of generation of internal waves was preceded by a change in wind speed and direction.

The maximum NIW energy occurred in the winter and the minimum in the summer. The blue-shift of the NIW's frequency was in the range 1.002*f* –1.038*f*. The shift tended to be more significant in the summer season. The NIWs occurred more frequently in the months of transient weather in March and September, also in December and January, when the wind forcing was the strongest. During the winter, intermittency of the NIWs was 67–87%, and long packets consisting of up to 11–16 near-inertial oscillations were observed at a depth of 100 m. In the winter, the NIW's packets had the longest modulation time scale, so the current speed grew during initial 3–6 fluctuations, i.e., by the mid of a wave packet. The duration of the longest NIW's packet was close to the time scale of \sim 10 days, which is expected to be consistent with the frontal-geostrophic adjustment [\[65\]](#page-20-19). By contrast, short packets of 1.3–9 oscillations were typical for the summer. The relatively short time scales may indicate that the ocean fronts are rare at the intermediate depths in the central part of the Black Sea in the summer.

The preliminary results we presented look promising and raise new questions. For future investigations to identify the space-time propagation of the NIWs, it is necessary to obtain data from several points and for the entire water column. This could be a challenge for the problems of internal wave propagation and generation, which remain open for the Black Sea. Further progress requires continuing such investigations using the slab model [\[4,](#page-18-3)[66\]](#page-20-20) to directly calculate the near-inertial energy flux input into the sea by the wind.

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Appendix A Appendix A Appendix A

Figure A1. Top—wind speed $\mathbf{U}(t)$ vector plot from the ERA-5 meteorological reanalysis data. The second plot from the top is the current velocity $\mathbf{u}(t)$ at 100 m at ABS No. 1 in January 2017, and the third plot from the top shows the u - and v -components of the current velocity.

 ϵ econd plot from the top is the current velocity $\mathbf{u}(t)$ at 100 m at ABS No. 1 in March 2017, and the third plot from the top shows the u - and v -components of the current velocity. **Figure A2.** Top—wind speed () vector plot from the ERA-5 meteorological reanalysis data. The second plot from the top is the current velocity $\mathbf{u}(t)$ at 100 m at ABS No. 1 in March 2017, and the the top is the current velocity $\mathbf{u}(t)$ at 100 m at ABS No. 1 in March 2017, and the **Figure A2.** Top—wind speed **U**(*t*) vector plot from the ERA-5 meteorological reanalysis data. The

Appendix B

Near-Inertial Fluctuations of the Deep Sea Temperature in the Northeastern Part of the Black Sea Basin

The RBR Duet instrument for observations of temperature and pressure was deployed at 265 m, i.e., much deeper than the main pycnocline in the northeastern Black Sea basin, from 18 December 2017 to 4 December 2018. At this depth, temperature *T* and pressure *P* fluctuations with a time scale close to the inertial period were also identified. Figure $A3$ shows an excerpt of the record for January 2018. The most intense fluctuations were observed on 12–18 January. Over the near-inertial period, the temperature *T* changed by only 0.02 ◦C, and the pressure sensor readings changed up to 1 dbar due to the wave passage. In such a way, oscillations with a near-inertial period were observed every month. The fluctuations were within 0.01–0.015 $°C$, and the pressure amplitude did not exceed 1.5 dbar. Notice that inertial oscillations were clear-cut mostly in autumn and winter and almost decayed in summer.

> The estimates of the power spectra of the time series $T(t)$ for each month show the peak at a near-inertial frequency of 0.059–0.06 cph (16.95–16.67 h) (Figure [A4\)](#page-18-12). At this mooring site, the local inertial frequency *f* was 0.057 cph (17.54 h). Alike the sea current oscillations in the central part of the Black Sea (see the main body of this paper above), the analysis of the temperature variability also revealed the seasonal intensification of near-inertial fluctuations in December–January and, surprisingly, in June. The ratio of the maximum and minimum spectral energy of near-inertial oscillations was \sim 9.5. In the spring and autumn seasons, the power of near-inertial fluctuations was approximately the same.

Figure A3. Temporal variability of sea temperature and pressure at 265 m in January, 2018, from the cheap of ABS No. 2 observations of ABS No. 2.

Figure A4. Power spectra of temperature fluctuations at 265 m depth at ABS No. 2 for different **Figure A4.** Power spectra of temperature fluctuations at 265 m depth at ABS No. 2 for different months. The dashed line indicates the local inertial frequency at the mooring site. The small image in the upper right corner is an enlargement of the figure to better visualize the peaks close to the in the upper right corner is an enlargement of the figure to better visualize the peaks close to the inertial periods.

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