

Article **Modulations of Stochastic Modeling in the Structural and Energy Aspects of the Kundu–Mukherjee–Naskar System**

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Abstract: By using stochastic modeling, the investigation of the energy and wave characteristics of novel structures that develop in the sea and ocean currents becomes one of the most important advancements in the generation of sustainable and renewable energy. Theoretical examinations of random nonlinear Kundu–Mukherjee–Naskar (RNKMN) structures have become recommended in a random mode. The two-dimensional RNKMN equation permits exact and solved solutions that give rise to solitonic structures with adaptable properties. The obtained stochastic waves, under the influence of random water currents, represent a dynamically controlled system. It has been demonstrated that the stochastic parameter modulates wave forcing and produces energy wave collapse accompanied by medium turbulence. The fundamental wave characteristics establish an exact pattern for describing sea and ocean waves.

Keywords: stochastic modeling; stochastic waves; mathematical solver; spatio-temporal modeling; stochastic solutions; wave collapsing; sea and ocean wave applications

MSC: 35C05; 35R60; 60H30; 35Q55

1. Introduction

Different types of nonlinear partial differential equations (NPDEs) can be used to describe a variety of complex nonlinear physical processes [\[1](#page-9-0)[–3\]](#page-9-1). In fact, NPDEs have been the most extensively researched objects in various fields of applied science, such as molecular biology, optical fiber communications, chemical engineering, superfluids, solid-state physics, plasma physics, and many others $[4–6]$ $[4–6]$. The topic of optical solitons is critical for the exploration of soliton propagation via nonlinear fiber. One of the most well-known governing equations for the dynamics of optical solitons is the nonlinear Schrödinger equation (NLSE), which plays a crucial role in the dynamics of nonlinear effects in fiber optic communications. The NLSEs, which disclose solitary-type solutions, have become the primary representative method for defining wave behaviors in numerous vital applications [\[7–](#page-9-4)[10\]](#page-9-5).

Stochastic nonlinear partial differential equations (SNPDEs) are foundational models of physical systems with unpredictable inputs, interactions, or environments. Indeed, SNPDEs perform an important role in various fields of natural sciences [\[11](#page-9-6)[–14\]](#page-9-7). Therefore, finding solutions to SNPDEs is a prominent research topic. Brownian motion (Wiener process) is a classic stochastic process that is both a martingale and a Markov process [\[15\]](#page-9-8). In dispersive environments, Brownian motion is frequently employed as a stochastic process. It is used in biology, kinematics, new physics, engineering, and the study of flower pollen in water [\[16–](#page-9-9)[18\]](#page-10-0). Indeed, there is a vital link between stochastic processes and partial differential equations. There are numerous processes that depend on particles moving

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stochastically in random potentials. As previously discovered, noise can be incorporated directly into the equation when examining wave propagation using stochastic nonlinear dispersive equations.

In this work, we demonstrate how the newly extended NLSE can be understood in terms of Brownian motion. Specifically, we investigate the newly extended stochastic NLSE (SNLSE) with Kerr law nonlinearity, which is also known as the stochastic nonlinear Kundu–Mukherjee–Naskar (SKMN) equation [\[19](#page-10-1)[–22\]](#page-10-2):

$$
i\phi_t + \alpha \phi_{xy} + i\beta \phi (\phi \phi_x^* - \phi^* \phi_x) - i\delta \phi(x, t) \Xi_t = 0, \quad i = \sqrt{-1},
$$
 (1)

where $\varphi(x, y, t)$ represents the nonlinear wave envelope and $*$ refers to the complex conjugate. The first and second terms represent, respectively, the temporal evolution of the wave and the disturbance of the dispersion. The parameter β is distinct from conventional Kerr law nonlinearity. The noise Ξ_t is a Brownian time derivative of $\Xi(t)$, and δ denotes the noise amplitude [\[23\]](#page-10-3). Hole waves, oceanic rogue waves, bending of light beams, and erbium atoms are all described by Equation [\(1\)](#page-1-0) [\[21](#page-10-4)[,22\]](#page-10-2).

This paper investigates certain aspects of the influence of noise, in the Itô sense, on the SKMN equation. This is an enormously broad and fascinating field, with active study in a variety of directions. This topic's ability to mix methods from both classical and stochastic analyses is one of its most intriguing aspects [\[24\]](#page-10-5). We employ a robust solver [\[25\]](#page-10-6) to reveal some new stochastic solutions for the SKMN equation. These solutions demonstrate various important physics related to hole waves, oceanic rogue waves, bending of light beams and erbium atoms, and many others.

The organization of this article is as follows. Section [2](#page-1-1) presents the SKMN in the Itô sense and its corresponding potential. Section [3](#page-2-0) presents the closed form of solutions to the SKMN equation using a powerful solver. Section [4](#page-3-0) provides a physical interpretation of the solution to the SKMN equation. The results are then summarized in Section [5.](#page-8-0)

2. The New Stochastic Solutions

The traveling wave solution [\[21\]](#page-10-4) in Equation [\(1\)](#page-1-0) is used as follows:

$$
\phi(x,y,t) = e^{i(-b_1x - b_2y + \Omega t + \theta) + \delta \Xi(t) - \delta^2 t} Q(\eta), \quad \eta = k_1 x + k_2 y - \omega t. \tag{2}
$$

Here, b_1 and b_2 represent the soliton frequencies in the *x*- and *y*-directions, θ is the phase constant of the soliton, and δ represents the wave number. k_1 and k_2 denote the inverse width of the soliton along the *x*- and *y*-directions, and ω denotes the soliton velocity. Equation [\(1\)](#page-1-0) is reduced to two equations in the form of

$$
-\alpha b_1 b_2 Q(\eta) - 2\beta b_1 Q^3(\eta) e^{2\delta E(t) - 2\delta^2 t} + \alpha k_1 k_2 Q''(\eta) - \Omega Q(\eta) = 0
$$
\n(3)

for real and imaginary parts:

$$
\alpha b_2 k_1 Q'(\eta) + \alpha b_1 k_2 Q''(\eta) + \delta^2 Q(\eta) = 0.
$$
\n(4)

In order to formulate the system of equations (real and imaginary), one can take the expectation on both sides of Equation [\(3\)](#page-1-2), yielding

$$
-\alpha b_1 b_2 Q(\eta) - 2\beta b_1 Q(\eta)^3 e^{-2\delta^2 t} E(e^{2\delta \Xi(t)}) + \alpha k_1 k_2 Q''(\eta) - \Omega Q(\eta) = 0.
$$
 (5)

Indeed, $E(e^{2\delta \Xi(t)}) = e^{2\delta^2 t}$, and then Equation [\(5\)](#page-1-3) becomes

$$
\alpha k_1 k_2 Q''(\eta) - 2\beta b_1 Q^3(\eta) - (\alpha b_1 b_2 + \Omega) Q(\eta) = 0.
$$
 (6)

By solving Equation [\(4\)](#page-1-4), the dispersion-constrained equation for Equation [\(6\)](#page-1-5) reads:

$$
-b_1\left(\alpha b_2 + 2\beta e^{\frac{2\delta^2 \eta}{\alpha b_2 k_1 + \alpha b_1 k_2 + \omega}}\right) + \frac{\alpha \delta^4 k_1 k_2}{(\alpha b_2 k_1 + \alpha b_1 k_2 + \omega)^2} - \Omega = 0. \tag{7}
$$

To find exact solutions to Equation [\(6\)](#page-1-5), we integrate it, and after some algebraic steps, it is transformed into an energy equation with the potential

$$
V = -\frac{1}{2\alpha k_1 k_2} \alpha b_1 b_2 Q^2(\eta) - \frac{1}{2\alpha k_1 k_2} \beta b_1 Q^4(\eta) - \frac{1}{2\alpha k_1 k_2} \Omega Q^2(\eta). \tag{8}
$$

By solving Equation [\(8\)](#page-2-1), an exact solution to Equation [\(5\)](#page-1-3) is expressed as:

$$
Q(\eta) = \frac{2(\alpha b_1 b_2 + \Omega) e^{\frac{\eta \sqrt{\alpha b_1 b_2 + \Omega}}{\sqrt{\alpha} \sqrt{k_1} \sqrt{k_2}}}}{\sqrt{-\beta b_1 (\alpha b_1 b_2 + \Omega)} \left(e^{\frac{2\eta \sqrt{\alpha b_1 b_2 + \Omega}}{\sqrt{\alpha} \sqrt{k_1} \sqrt{k_2}}} + 1\right)},
$$

$$
\phi(x, y, t) = \frac{2(\alpha b_1 b_2 + \Omega) e^{\frac{\eta \sqrt{\alpha b_1 b_2 + \Omega}}{\sqrt{\alpha} \sqrt{k_1} \sqrt{k_2}} e^{i(-b_1 x - b_2 y + \Omega t + \theta) + \delta \Xi(t) - \delta^2 t}}{\sqrt{-\beta b_1 (\alpha b_1 b_2 + \Omega)} \left(e^{\frac{2\eta \sqrt{\alpha b_1 b_2 + \Omega}}{\sqrt{\alpha} \sqrt{k_1} \sqrt{k_2}}} + 1\right)}.
$$
(9)

3. Closed-Form Solutions

We provide some new stochastic solutions to Equation [\(1\)](#page-1-0). In view of the unified solver [\[25\]](#page-10-6), the stochastic solutions to Equation [\(1\)](#page-1-0) are as follows:

Family I:

$$
Q_{1,2}(x,y,t) = \pm \sqrt{\frac{-(\alpha b_1 b_2 + \Omega)}{\beta b_1}} \operatorname{sech}\left(\pm \sqrt{\frac{\alpha b_1 b_2 + \Omega}{\alpha k_1 k_2}} (k_1 x + k_2 y - \omega t)\right).
$$
 (10)

Hence, the solutions to Equation [\(1\)](#page-1-0) are

$$
\phi_{1,2}(x,y,t) = \pm \sqrt{\frac{-(\alpha b_1 b_2 + \Omega)}{\beta b_1}} e^{i(-b_1 x - b_2 y + \Omega t + \theta) + \delta \Xi(t) - \delta^2 t} \operatorname{sech}\left(\pm \sqrt{\frac{\alpha b_1 b_2 + \Omega}{\alpha k_1 k_2}} (k_1 x + k_2 y - \omega t)\right).
$$
(11)

Family II:

$$
Q_{3,4}(x,y,t) = \pm \sqrt{\frac{-35\left(\alpha b_1 b_2 + \Omega\right)}{36\beta b_1}} \operatorname{sech}^2\left(\pm \sqrt{\frac{5\left(\alpha b_1 b_2 + \Omega\right)}{12\alpha k_1 k_2}} \left(k_1 x + k_2 y - \omega t\right)\right). \tag{12}
$$

Hence, the solutions to Equation [\(1\)](#page-1-0) are

$$
\phi_{3,4}(x,y,t) = \pm \sqrt{\frac{-35 \left(\alpha b_1 b_2 + \Omega\right)}{36 \beta b_1}} e^{i(-b_1 x - b_2 y + \Omega t + \theta) + \delta \Xi(t) - \delta^2 t} \operatorname{sech}^2\left(\pm \sqrt{\frac{5 \left(\alpha b_1 b_2 + \Omega\right)}{12 \alpha k_1 k_2}} \left(k_1 x + k_2 y - \omega t\right)\right). \tag{13}
$$

Family III:

$$
Q_{5,6}(x,y,t) = \pm \sqrt{\frac{-(\alpha b_1 b_2 + \Omega)}{2\beta b_1}} \tanh\left(\pm \sqrt{\frac{-(\alpha b_1 b_2 + \Omega)}{2\alpha k_1 k_2}} (k_1 x + k_2 y - \omega t)\right).
$$
 (14)

Hence, the solutions to Equation [\(1\)](#page-1-0) are

$$
\phi_{5,6}(x,y,t) = \pm \sqrt{\frac{-(\alpha b_1 b_2 + \Omega)}{2\beta b_1}} e^{i(-b_1 x - b_2 y + \Omega t + \theta) + \delta \Xi(t) - \delta^2 t} \tanh\left(\pm \sqrt{\frac{-(\alpha b_1 b_2 + \Omega)}{2\alpha k_1 k_2}} (k_1 x + k_2 y - \omega t)\right).
$$
(15)

4. Physical Interpretation

In this work, we study the system of the Kundu–Mukherjee–Naskar model, which is a powerful nonlinear model that exhibits the characteristics of the dynamics of hole and oceanic waves. Namely, we consider this model via the Brownian motion process. This process is defined as a stochastic process, which is continuous in time. For more properties of the Brownian motion process $\{W(t)\}_{t\geq0}$, see [\[15\]](#page-9-8).

In order to obtain some unique stochastic solutions to this model with multiplicative noise in the Itô sense, we use the unified solver approach. This theoretical solver process has been applied to identify some new and brief random solutions to the RNKMN model with multiplicative random parameters, resulting in Solutions (9) , (11) , (13) , and (15) , which exhibited a diversity of solitonic and dissipative structures. The obtained results account for a number of intriguing wave phenomena with applications in water wave physics and engineering. In fact, the provided solutions included blow-up and shock structure solutions to the nonlinear system, in addition to bright, rational explosive, breather, shocklike, explosively dissipated, and dark solitons.

The model with the noise term and random function $E(t)$ reduces to Equation [\(6\)](#page-1-5). The expectation of [\(3\)](#page-1-2) with $E(e^{2\delta \Xi(t)}) = e^{2\delta^2 t}$ transforms the equation into a probability density function that is solved to yield several important solutions. Using Matlab Release 18 [\[26\]](#page-10-7), we plot 2D, 3D, and contour graphs to explain the wave structures for some chosen solutions to Equation [\(6\)](#page-1-5) with appropriate parametric values, i.e., $b_1 = b_2 = k_1 = k_2 = 0.5$, $\alpha = 2, \beta = -1.7.$

Solution [\(9\)](#page-2-2) expresses a set of random solitonic depictions, as shown in Figure [1.](#page-4-0) Figure [1a](#page-4-0),b present the effect of the intensive randomness coefficient on the structure, amplitude, bandwidth, and energy. Figure [1b](#page-4-0) shows that with increasing time *t*, the influence of randomness increases along with the rate of wave collapse. It was noted that near-complete collapse occurs at $t = 4$. By increasing δ , we found that both wave amplitude and width decreased and the wave began to collapse, as shown in Figure [1c](#page-4-0).

In the same context, it was found that the dark solution [\(15\)](#page-3-1), which exhibits a dissipative wave, is affected by time *t* and the random coefficient δ , as shown in Figure [2.](#page-5-0) By increasing *t*, the rate of collapse of the dissipative wave increased, as shown in Figure [2b](#page-5-0). Also, the parameter δ caused the wave to collapse and transform into a superwave with a small amplitude, as shown in Figure [2c](#page-5-0). In the absence of random effects, the studied solutions are of great importance due to their different solitary characteristics, some of which we aim to highlight. For example, Solution [\(9\)](#page-2-2) yielded breather waves and both localized and super solitons, as shown in Figure [3.](#page-6-0) Solution [\(13\)](#page-2-4) formed two types of applicable waves: the first was a shock-like soliton and the second was a winged soliton, as shown in Figure [4.](#page-7-0) Figure [4a](#page-7-0),b depicts the overlap between the solitonic and dissipative waves, which is dependent on the system parameters.

On the other hand, Solution [\(15\)](#page-3-1) is considered one of the most useful physical applications in the study of dissipative and explosive waves. Figure [5a](#page-8-1),b show the formation of the dissipated oscillatory waves with time. By varying *x*, the oscillatory form turned into a rational explosive structure with high energy. Finally, Figure [6a](#page-8-2),b depict a dissipative blow-up wave.

In fact, the results of random waves exhibiting the properties of stochastic pulses obtained in this work represent great progress in explaining physical phenomena with decaying and damping characteristics that can cause a sharp decrease in wave amplitude and dissipation coefficients. They are consistent with many of the observations made in various fields, such as electrostatic waves in fiber optics, space plasma, and deep ocean water waves propagating in random environments [\[20](#page-10-8)[,27](#page-10-9)[–29\]](#page-10-10). If the stochastic properties of the random pulse are neglected, our results are consistent with previous work on solitonic and shock solutions [\[19–](#page-10-1)[21\]](#page-10-4). We also provided a precise description of the explosive, explosive-shock, and dissipative solutions of shock-like solitons, in addition to the winged soliton solutions. This highlights the accuracy and importance of the method used in this work [\[18](#page-10-0)[,29](#page-10-10)[–31\]](#page-10-11).

In summary, the merits of the stochastic structures of the RNKMN model with the noise term inspired the dynamical energy properties of the obtained solitary and dissipative waves.

(**a**) Trajectory of $\phi(x, y, t)$ for the deterministic case.

(**b**) Trajectory of $\phi(x, y, t)$ for the stochastic case.

(c) Plot of $\phi(x, y, t)$ with x, δ .

Figure 1. A set of random solitonic depictions of $\phi(x, y, t)$.

(**a**) Trajectory of $\phi_5(x, y, t)$ for the deterministic case.

(**b**) Trajectory of $\phi_5(x, y, t)$ for the stochastic case.

(c) Plot of $\phi_5(x, y, t)$ with x, δ .

Figure 2. A set of random solitonic depictions of $\phi_5(x, y, t)$.

(a) Plot of $\phi(x, y, t)$ with x, t .

(**b**) Plot of $|\phi(x, y, t)|$ with *x*, *t*.

(c) Plot of $\phi(x, y, t)$ with x, y .

Figure 3. A set of solitonic depictions of $\phi(x, y, t)$.

(**d**) Plot of $|\phi_3(x, y, t)|$ with x, t . **Figure 4.** A set of solitonic depictions of $\phi_3(x, y, t)$.

(a) 3D plot of $\text{Re}\phi_5(x, y, t)$ with x, t .

(**b**) Contour plot of $\text{Re}\phi_5(x, y, t)$ with x, t .

Figure 5. A set of solitonic structural of $\phi_5(x, y, t)$.

(a) 3D plot for blow-up solution of $\text{Re}\phi_5(x, y, t)$.

(**b**) Contour plot for blow-up solution of $\text{Re}\phi_5(x, y, t)$.

Figure 6. Plot of $\text{Re}\phi_5(x, y, t)$ with x, y .

5. Conclusions

We implemented a unified solver to solve the RNKMN equation. Significant wave characteristics for exact solitary, oscillatory shock, breather, and super soliton waves were

examined in the RNKMN equation. The modulations of stochastic parameters on the obtained solution amplitude and energy were investigated. It was noted that the stochastic effects can demonstrate some amendments in collapsing dissipative and explosive water waves. This theoretical investigation could be used in sea and ocean wave applications. Our goal is to expand this approach to include decaying and damped characteristic dispersion and dissipation nonlinear equations in future work. It is also necessary to investigate the explosive, explosive-shock, and dissipative solutions of our model under fractal random environmental conditions.

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