


Article

The Pauli Problem for Gaussian Quantum States: Geometric Interpretation

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Abstract: We solve the Pauli tomography problem for Gaussian signals using the notion of Schur complement. We relate our results and method to a notion from convex geometry, polar duality. In our context polar duality can be seen as a sort of geometric Fourier transform and allows a geometric interpretation of the uncertainty principle and allows to apprehend the Pauli problem in a rather simple way.

Keywords: covariance matrix; polar duality; uncertainty principle; reconstruction problem

1. The Pauli Problem and Quantum Tomography

The problem goes back to Pauli's question [1]:

The mathematical problem as to whether, for given probability densities $W(p)$ and $W(x)$, wave function $\psi(\dots)$ is always uniquely determined, has still not been investigated in its generality.

The answer to Pauli's question is negative [2]; there is a general nonuniqueness of the solution (for a detailed discussion of the Pauli problem and its applications, see [3]). The problem can actually be formulated as from statistical quantum mechanics as follows: can we estimate the density matrix of the said state using repeated measurements on identical quantum systems? After having obtained measurements on these identical systems, can we make a statistical inference about their probability distributions (e.g., [4])? Such a procedure is an instance of quantum state tomography, and is practically implemented using a set of measurements of a so-called quorum of observables. It can be performed using various mathematical techniques, for instance the Radon–Wigner transform that we discussed in [5]; the latter has important applications in medical imaging [6]. For details and explicit constructions, see [7–14], and [15] by Man'ko and Man'ko.

Remark 1. *Everything in this paper extends mutatis mutandis to time-frequency analysis, replacing the notion of wave function by that of a signal. In this case, one takes $\hbar = 1/2\pi$ and replaces phase-space variables (x, p) with time-frequency variables (x, ω) .*

2. A Simple Example

Let us discuss the Pauli problem on the simplest possible example, that of a Gaussian wave function in one spatial dimension. Assuming for simplicity, it is centered at the origin and is given by formula

$$\psi(x) = \left(\frac{1}{2\pi\sigma_{xx}}\right)^{1/4} e^{-\frac{x^2}{4\sigma_{xx}}} e^{\frac{i\sigma_{xp}}{2\hbar\sigma_{xx}}x^2} \quad (1)$$

where σ_{xx} is the variance in the position variable, and σ_{xp} the covariance in the position and momentum variables. Fourier transform

$$\hat{\psi}(p) = \frac{1}{\sqrt{2\pi\hbar}} \int_{-\infty}^{\infty} e^{-\frac{i}{\hbar}px} \psi(x) dx$$



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of the ψ is explicitly given by

$$\widehat{\psi}(p) = \left(\frac{1}{2\pi\sigma_{pp}}\right)^{1/4} e^{-\frac{p^2}{4\sigma_{pp}}} e^{-\frac{i\sigma_{xp}}{2\hbar\sigma_{pp}}p^2} \tag{2}$$

hence, the knowledge of σ_{xx} and of σ_{pp} , that is, of moduli $|\psi(x)|^2$ and $|\widehat{\psi}(p)|^2$, determines covariance σ_{xp} up to a sign because state ψ saturates the Robertson–Schrödinger inequality; so, we have

$$\sigma_{xx}\sigma_{pp} - \sigma_{xp}^2 = \frac{1}{4}\hbar^2 \tag{3}$$

This identity can be solved in σ_{xp} yielding $\sigma_{xp} = \pm(\sigma_{xx}\sigma_{pp} - \frac{1}{4}\hbar^2)^{1/2}$. The state and its Fourier transform are given by formulas

$$\psi_{\pm}(x) = \left(\frac{1}{2\pi\sigma_{xx}}\right)^{1/4} e^{-\frac{x^2}{4\sigma_{xx}}} e^{\pm\frac{i\sigma_{xp}}{2\hbar\sigma_{xx}}x^2} \tag{4}$$

and

$$\widehat{\psi}_{\pm}(p) = \left(\frac{1}{2\pi\sigma_{pp}}\right)^{1/4} e^{-\frac{p^2}{4\sigma_{pp}}} e^{\mp\frac{i\sigma_{xp}}{2\hbar\sigma_{pp}}p^2} . \tag{5}$$

Both functions ψ_+ and $\psi_- = \psi_+^*$ and their Fourier transforms $\widehat{\psi}_+$ and $\widehat{\psi}_-$ satisfy conditions $|\psi_+(x)|^2 = |\psi_-(x)|^2$ and $|\widehat{\psi}_+(p)|^2 = |\widehat{\psi}_-(p)|^2$ showing that the Pauli problem does not have a unique solution. In Corbett’s [16] terminology ψ_+ and ψ_- are “Pauli partners”. Let us now have a look at these things from the perspective of the Wigner transform

$$W\psi(x, p) = \frac{1}{2\pi\hbar} \int_{-\infty}^{\infty} e^{-\frac{i}{\hbar}py} \psi(x + \frac{1}{2}y) \psi^*(x - \frac{1}{2}y) dy$$

of Gaussian ψ . A straightforward calculation involving Gaussian integrals [17] yields, setting $z = (x, p)$, normal distribution

$$W\psi_{\pm}(z) = \frac{1}{2\pi\sqrt{\det \Sigma_{\pm}}} e^{-\frac{1}{2}\Sigma_{\pm}^{-1}z \cdot z} \tag{6}$$

where covariance matrix

$$\Sigma_{\pm} = \begin{pmatrix} \sigma_{xx} & \pm\sigma_{xp} \\ \pm\sigma_{px} & \sigma_{pp} \end{pmatrix}$$

has determinant $\det \Sigma_{\pm} = \frac{1}{4}\hbar^2$ in view of equality (3); hence,

$$W\psi_{\pm}(z) = \frac{1}{\pi\hbar} e^{-\frac{1}{2}\Sigma_{\pm}^{-1}z \cdot z} . \tag{7}$$

Associated covariance matrices are thus

$$\Omega_{\pm} = \{z : \frac{1}{2}\Sigma_{\pm}^{-1}z \cdot z \leq 1 . \}$$

3. Multivariate Case: Asking the Right Questions

We generalize the discussion to the multivariate case where the real variables x and p are replaced with real vectors $x = (x_1, \dots, x_n)$, $p = (p_1, \dots, p_n)$.

The Wigner function cannot be directly measured, but its marginal distributions can (they are classical probability densities). In analogy with Formula (6) we determine a (centered) Gaussian, ψ such that

$$W\psi(z) = \left(\frac{1}{2\pi}\right)^n \frac{1}{\sqrt{\det \Sigma}} e^{-\frac{1}{2}\Sigma^{-1}z \cdot z} \tag{8}$$

where $z = (x, p)$, and the covariance matrix is

$$\Sigma = \begin{pmatrix} \Sigma_{XX} & \Sigma_{XP} \\ \Sigma_{PX} & \Sigma_{PP} \end{pmatrix}, \quad \Sigma_{PX} = \Sigma_{XP}^T. \tag{9}$$

Here, the the n -dimensional Wigner transform $W\psi$ is defined by

$$W\psi(x, p) = \left(\frac{1}{2\pi\hbar}\right)^n \int e^{-\frac{i}{\hbar}p \cdot y} \psi\left(x + \frac{1}{2}y\right) \psi^*\left(x - \frac{1}{2}y\right) d^n y.$$

The most straightforward way to determine this state is to use the properties of the Wigner transform itself. Let us start with the marginal properties [17]:

$$\int W\psi(x, p) d^n p = |\psi(x)|^2 \tag{10}$$

$$\int W\psi(x, p) d^n x = |\widehat{\psi}(p)|^2 \tag{11}$$

where the n -dimensional Fourier transform $\widehat{\psi}$ is given by

$$\widehat{\psi}(p) = \left(\frac{1}{2\pi\hbar}\right)^{n/2} \int e^{-\frac{i}{\hbar}p x} \psi(x) d^n x.$$

These formulas hold as soon as both ψ and $\widehat{\psi}$ are in $L^1(\mathbb{R}^n) \cap L^2(\mathbb{R}^n)$ [17]. These quantities allow for determining matrices

$$\Sigma_{XX} = (\sigma_{x_j x_k})_{1 \leq j, k \leq n} \quad \text{and} \quad \Sigma_{PP} = (\sigma_{p_j p_k})_{1 \leq j, k \leq n}$$

by usual formulas

$$\sigma_{x_j x_k} = \int x_j x_k |\psi(x)|^2 d^n x, \quad \sigma_{p_j p_k} = \int p_j p_k |\widehat{\psi}(p)|^2 d^n p$$

and an elementary calculation of Gaussian integrals yields the values

$$|\psi(x)| = \left(\frac{1}{2\pi}\right)^{n/4} (\det \Sigma_{XX})^{-1/4} e^{-\frac{1}{4} \Sigma_{XX}^{-1} x \cdot x} \tag{12}$$

$$|\widehat{\psi}(p)| = \left(\frac{1}{2\pi}\right)^{n/4} (\det \Sigma_{PP})^{-1/4} e^{-\frac{1}{4} \Sigma_{PP}^{-1} p \cdot p}. \tag{13}$$

Here, we are exactly in the situation discussed by Pauli: $|\psi(x)|$ and $|\widehat{\psi}(p)|$ are what we can measure, so we can determine covariance blocks Σ_{XX} and Σ_{PP} , but not covariance Σ_{XP} : knowledge of the latter (and hence of $\Sigma_{PX} = \Sigma_{XP}^T$) is necessary to entirely determine state ψ . In the previous section, the problem was solved: in case $n = 1$, blocks Σ_{XX} , Σ_{PP} , and Σ_{XP} were scalars σ_{xx} , σ_{pp} , and σ_{xp} , and these are related by the uncertainty principle in the form of $\sigma_{xx}\sigma_{pp} - \sigma_{xp}^2 = \frac{1}{4}\hbar^2$ yielding two possible values $\sigma_{xp} = \pm(\sigma_{xx}\sigma_{pp} - \frac{1}{4}\hbar^2)^{1/2}$, and hence the two states (5). In the multidimensional, case we also have a simple (but not immediately obvious) formula connecting the blocks of the covariance matrix. The way out of this problem consists in using a general formula [17–19], which was initially proved by Bastiaans [20] in connection with first-order optics. Let X and Y be real $n \times n$ matrices, such that $X = X^T > 0$ and $Y = Y^T$, and set

$$\psi_{X,Y}(x) = \left(\frac{1}{\pi\hbar}\right)^{n/4} (\det X)^{1/4} e^{-\frac{1}{2\hbar}(X+iY)x \cdot x}. \tag{14}$$

This function is normalized to unity: $\|\psi_{X,Y}\|_{L^2} = 1$, and its Wigner transform is given by

$$W\psi_{X,Y}(z) = \left(\frac{1}{\pi\hbar}\right)^n e^{-\frac{1}{\hbar}Gz \cdot z} \tag{15}$$

where G is the symmetric matrix

$$G = \begin{pmatrix} X + YX^{-1}Y & YX^{-1} \\ X^{-1}Y & X^{-1} \end{pmatrix}. \tag{16}$$

A fundamental fact, which is related to the uncertainty principle, is that G is a symplectic matrix, i.e., it belongs to symplectic group $\text{Sp}(n)$. Equivalently, since $G = G^T$,

$$G^T J G = G J G = J$$

where

$$J = \begin{pmatrix} 0_{n \times n} & I_{n \times n} \\ -I_{n \times n} & 0_{n \times n} \end{pmatrix}$$

the standard symplectic matrix. We have $G = S^T S$, where

$$S = \begin{pmatrix} X^{1/2} & 0 \\ X^{-1/2}Y & X^{-1/2} \end{pmatrix} \tag{17}$$

is clearly symplectic. Assuming that function ψ for which we are looking is a Gaussian, comparing Formulas (8) and (15) leads to identification

$$\Sigma = \frac{\hbar}{2} G^{-1}.$$

Since $G J G = J$ the inverse G^{-1} is $-J G J$, explicit formula

$$G^{-1} = \begin{pmatrix} X^{-1} & -X^{-1}Y \\ -YX^{-1} & X + YX^{-1}Y \end{pmatrix}$$

so that there remains to solve matrix equation

$$\begin{pmatrix} \Sigma_{XX} & \Sigma_{XP} \\ \Sigma_{PX} & \Sigma_{PP} \end{pmatrix} = \frac{\hbar}{2} \begin{pmatrix} X^{-1} & -X^{-1}Y \\ -YX^{-1} & X + YX^{-1}Y \end{pmatrix}. \tag{18}$$

It immediately follows that we have $X = \frac{\hbar}{2} \Sigma_{XX}^{-1}$ and $Y = -\frac{2}{\hbar} \Sigma_{XP} \Sigma_{XX}^{-1}$, so the unknown Gaussian for which we were looking is

$$\psi(x) = \left(\frac{1}{2\pi}\right)^{n/4} (\det \Sigma_{XX})^{-1/4} \exp \left[-\left(\frac{1}{4} \Sigma_{XX}^{-1} + \frac{i}{2\hbar} \Sigma_{XP} \Sigma_{XX}^{-1}\right) x \cdot x \right], \tag{19}$$

which is the n -dimensional variant of (1), replacing σ_{xx} with Σ_{XX} and σ_{xp} with Σ_{XP} . This does not solve completely our problem, however, because we do not know matrix Σ_{XP} . The crucial step is to notice that, as a bonus, we obtained from (18) the matrix form of the saturated Robertson–Schrödinger equality, namely,

$$\Sigma_{PP} \Sigma_{XX} - \Sigma_{XP}^2 = \frac{1}{4} \hbar^2 I_{n \times n}. \tag{20}$$

From this formula we can deduce Σ_{XP}^2 , and one finds two Pauli partners

$$\psi_{\pm}(x) = \left(\frac{1}{2\pi}\right)^{n/4} (\det \Sigma_{XX})^{-1/4} \exp \left[-\left(\frac{1}{4} \Sigma_{XX}^{-1} \pm \frac{i}{2\hbar} \Sigma_{XP} \Sigma_{XX}^{-1}\right) x \cdot x \right] \tag{21}$$

once a value of Σ_{XP} is determined (even if $\Sigma_{XP}^2 = 0$, we can have $\Sigma_{XP} \neq 0$). Here, we solved a so-called “phase retrieval problem” (see Klibanov et al. [21] for a good review of the topic): in view of Formula (12), we know that

$$\psi(x) = e^{i\Phi(x)} \left(\frac{1}{2\pi}\right)^{n/4} (\det \Sigma_{XX})^{-1/4} e^{-\frac{1}{4} \Sigma_{XX}^{-1} x \cdot x} \tag{22}$$

where Φ is an unknown real function of the position variable. We identified this phase here as being function

$$\Phi(x) = -\left(\frac{1}{2\hbar}\Sigma_{XP}\Sigma_{XX}^{-1}\right)x \cdot x.$$

4. Geometric Interlude

We introduce the notion of \hbar -polarity and duality; we see in the next section that this notion from convex geometry is quite unexpectedly related to the Pauli problem, of which it gives a limpid geometric interpretation. For a very detailed study of polarity, see Charalambos and Aliprantis [22]. In both sources, alternative competing definitions are also described; the one we use here is the most common and the best fitted to our needs.

Let X be a nonempty subset of n -dimensional configuration space \mathbb{R}_x^n ; this may be, for instance, a set of position measurements performed on some physical system with n degrees of freedom. One defines the polar set of X as the set X^o of all points $p = (p_1, \dots, p_n)$ in the momentum space \mathbb{R}_p^n , such that

$$px = p_1x_1 + \dots + p_nx_n \leq 1$$

for all points $x = (x_1, \dots, x_n)$ in X . Similarly, if P is a subset of \mathbb{R}_p^n , one defines its polar P^o as the set of all x in \mathbb{R}_x^n , such that $px \leq 1$ for all p in P . We use a rescaled variant of the notion of polarity here, which we call \hbar polarity. By definition, the \hbar -polar X^\hbar of X is the set of all p , such that

$$px = p_1x_1 + \dots + p_nx_n \leq \hbar$$

for all points x in X . We have $X^\hbar = \hbar X^o$ and $P^\hbar = \hbar P^o$ likewise.

From now on, we assume for simplicity that X and P are convex bodies, i.e., they are convex, compact, and with a nonempty interior; we also assume that they are symmetric (i.e., $X = -X$), which implies, by convexity, that they contain 0 in their interior. Simple examples of such sets are balls and ellipsoids centered at the origin. Polar duals have the following remarkable properties:

- *Biduality:* $(X^\hbar)^\hbar = X$
- *Antimonotonicity:* $X \subset Y \implies Y^\hbar \subset X^\hbar$
- *Scaling property:* $L \in GL(n, \mathbb{R}) \implies (LX)^\hbar = (L^T)^{-1}X^\hbar$.

Let $\mathcal{B}_X^n(R)$ (resp. $\mathcal{B}_p^n(R)$) be the ball $\{x : |x| \leq R\}$ in \mathbb{R}_x^n (resp. $\{p : |p| \leq R\}$ in \mathbb{R}_p^n). We have

$$\mathcal{B}_X^n(\sqrt{\hbar})^\hbar = \mathcal{B}_p^n(\sqrt{\hbar}) \tag{23}$$

and one can show that $\mathcal{B}_X^n(\sqrt{\hbar})$ is the only self \hbar -dual set in \mathbb{R}_x^n . Let us extend this to the case of ellipsoids. An ellipsoid in \mathbb{R}_x^n centered at the origin (which is just an ordinary plane ellipse when $n = 1$) can always be viewed as the image of ball $\mathcal{B}_X^n(\sqrt{\hbar})$ by some invertible linear transformation L , in which case, it is given by inequality

$$L^{-1}x \cdot L^{-1}x = (LL^T)^{-1}x \cdot x \leq \hbar.$$

Conversely, if A is a positive definite symmetric matrix, inequality $Ax \cdot x \leq \hbar$ always defines an ellipsoid, since it is equivalent to the above inequality, taking for L inverse square root $A^{-1/2}$ of A . It immediately follows from the scaling property that the \hbar -polar of the ellipsoid is obtained by inverting the matrix of the ellipsoid:

$$X : Ax^2 \leq \hbar \iff X^\hbar : A^{-1}p \cdot p \leq \hbar \tag{24}$$

(that we have an equivalence follows from biduality property $(X^\hbar)^\hbar = X$).

5. The Pauli Problem and Polar Duality

Let us return to the Wigner transform of Gaussian states; using Formula (15), we can explicitly calculate $W\psi_{\pm}$, and one finds

$$W\psi_{\pm}(z) = (\pi\hbar)^{-n} e^{-\frac{1}{2}\Sigma_{\pm}^{-1}z \cdot z}$$

where covariance matrices Σ_{\pm} are given by

$$\Sigma_{\pm} = \begin{pmatrix} \Sigma_{XX} & \pm\Sigma_{XP} \\ \pm\Sigma_{PX} & \Sigma_{PP} \end{pmatrix},$$

with $\Sigma_{PX} = \Sigma_{XP}^T$. Two ellipsoids Ω_{\pm} centered at the origin correspond to Σ_{\pm} . Let us determine orthogonal projections $\Omega_{X,\pm}$ and $\Omega_{P,\pm}$ of Ω_{\pm} on the position and momentum spaces \mathbb{R}_x^n and \mathbb{R}_p^n .

5.1. Case $n = 1$

We begin with case $n = 1$, and projections are line segments. Here, $\Sigma_{XX} = \sigma_{xx}$, $\Sigma_{PP} = \sigma_{pp}$, and $\Sigma_{XP} = \Sigma_{PX} = \sigma_{xp}$ and covariance ellipses Ω_{\pm} are defined by

$$\frac{\sigma_{pp}}{2D}x^2 \mp \frac{\sigma_{xp}}{D}px + \frac{\sigma_{xx}}{2D}p^2 \leq 1 \tag{25}$$

where $D = \sigma_{xx}\sigma_{pp} - \sigma_{xp}^2 = \frac{1}{4}\hbar^2$ (cf. Formula (3)). Orthogonal projections $\Omega_{X,\pm}$ and $\Omega_{P,\pm}$ of Ω_{\pm} on the x and p axes are the same:

$$\Omega_X = [-\sqrt{2\sigma_{xx}}, \sqrt{2\sigma_{xx}}] , \quad \Omega_P = [-\sqrt{2\sigma_{pp}}, \sqrt{2\sigma_{pp}}] . \tag{26}$$

Let Ω_X^{\hbar} be the polar dual of Ω_X : it is the set of all numbers p , such that $px \leq \hbar$ for $-\sqrt{2\sigma_{xx}} \leq x \leq \sqrt{2\sigma_{xx}}$ and is thus the interval

$$\Omega_X^{\hbar} = [-\hbar/\sqrt{2\sigma_{xx}}, \hbar/\sqrt{2\sigma_{xx}}] .$$

Since $\sigma_{xx}\sigma_{pp} \geq \frac{1}{2}\hbar$, we have inclusion

$$\Omega_X^{\hbar} \subset \Omega_P \tag{27}$$

and this inclusion reduces to equality $\Omega_X^{\hbar} = \Omega_P$ if and only if the Heisenberg inequality is saturated, i.e., $\sigma_{xx}\sigma_{pp} = \frac{1}{4}\hbar^2$, which is equivalent to $\sigma_{xp} = 0$.

5.2. General Case

We have similar properties in arbitrary dimension n . To study this case, we first must find the orthogonal projections of covariance ellipsoid Ω on the position and momentum spaces. Ellipsoid Ω is given by equation $Mz \cdot z \leq \hbar$ where $M = \frac{\hbar}{2}\Sigma^{-1}$ is symmetric and positive definite ($M > 0$). Writing M in block form

$$M = \begin{pmatrix} M_{XX} & M_{XP} \\ M_{PX} & M_{PP} \end{pmatrix}$$

where $M_{XX} = M_{XX}^T$, $M_{PP} = M_{PP}^T$, and $M_{XP} = M_{PX}^T$ are $n \times n$ matrices; since $M > 0$, we also have $M_{XX} > 0$ and $M_{PP} > 0$. Then, the projections of Ω on \mathbb{R}_x^n and \mathbb{R}_p^n are ellipsoids given by, respectively [23],

$$\Omega_X : (M/M_{PP})x \cdot x \leq \hbar \quad , \quad \Omega_P : (M/M_{XX})p \cdot p \leq \hbar \tag{28}$$

where symmetric matrices

$$M/M_{PP} = M_{XX} - M_{XP}M_{PP}^{-1}M_{PX} \tag{29}$$

$$M/M_{XX} = M_{PP} - M_{PX}M_{XX}^{-1}M_{XP} \tag{30}$$

are Schur complements in M of M_{PP} and M_{XX} ; we have $M/M_{PP} > 0$ and $M/M_{XX} > 0$ so that Ω_X and Ω_P are nondegenerate (see Zhang’s treatise [24] for a detailed study of the Schur complement). To prove that inclusion $\Omega_X^{\hbar} \subset \Omega_P$ holds, we must show that cf. implication (24)) that

$$(M/M_{PP})(M/M_{XX}) \leq I_{n \times n} , \tag{31}$$

that is, that the eigenvalues of $(M/M_{PP})(M/M_{XX})$ must be smaller than 1. To prove this, we use the following essential remark: we showed above that matrix $M = \frac{\hbar}{2}\Sigma^{-1}$ is symplectic; therefore, its entries obey some constraints. Considering that M is also symmetric, these constraints are

$$M_{XX}M_{PP} - M_{XP}^2 = I_{n \times n} \tag{32}$$

$$M_{XX}M_{PX} = M_{XP}M_{XX} \tag{33}$$

$$M_{PX}M_{PP} = M_{PP}M_{XP} . \tag{34}$$

Using Identities (33) and (34), it follows that Schur complements (29) and (30) can be rewritten as

$$\begin{aligned} M/M_{PP} &= M_{XX} - M_{PP}^{-1}M_{PX}^2 \\ &= M_{PP}^{-1}(M_{PP}M_{XX} - M_{PX}^2) \\ &= M_{PP}^{-1} \end{aligned}$$

the last equality by using the transpose of Identity (32). Similarly,

$$M/M_{XX} = M_{PP} - M_{XX}^{-1}M_{XP}^2 = M_{XX}^{-1}$$

So, summarizing, Schur complements are given by

$$M/M_{PP} = M_{PP}^{-1} , M/M_{XX} = M_{XX}^{-1} . \tag{35}$$

It follows that

$$(M/M_{PP})(M/M_{XX}) = M_{PP}^{-1}M_{XX}^{-1} = (M_{XX}M_{PP})^{-1} .$$

We show that $(M/M_{PP})(M/M_{XX}) \leq I_{n \times n}$; equivalently, $M_{XX}M_{PP} \geq I_{n \times n}$. Now, since $M = \frac{\hbar}{2}\Sigma^{-1}$ is symplectic, so is matrix

$$M^{-1} = \frac{2}{\hbar}\Sigma = \begin{pmatrix} \frac{2}{\hbar}\Sigma_{XX} & \frac{2}{\hbar}\Sigma_{XP} \\ \frac{2}{\hbar}\Sigma_{PX} & \frac{2}{\hbar}\Sigma_{PP} \end{pmatrix}$$

hence, reinverting,

$$M = \begin{pmatrix} \frac{2}{\hbar}\Sigma_{PP} & -\frac{2}{\hbar}\Sigma_{PX} \\ -\frac{2}{\hbar}\Sigma_{XP} & \frac{2}{\hbar}\Sigma_{XX} \end{pmatrix} \tag{36}$$

so that $M_{XX}M_{PP} = \frac{4}{\hbar^2}\Sigma_{PP}\Sigma_{XX}$. In view of the generalized RSUP (20), we have

$$\Sigma_{PP}\Sigma_{XX} - \Sigma_{XP}^2 = \frac{1}{4}\hbar^2 I_{n \times n} \tag{37}$$

hence

$$M_{XX}M_{PP} = I_{n \times n} + \frac{4}{\hbar^2}\Sigma_{XP}^2 \tag{38}$$

and we are finished, provided that we can prove that $\Sigma_{XP}^2 \geq 0$ (which is obvious if $n = 1$), or, which amounts to the same $M_{XP}^2 \geq 0$. For this, since $M_{XX}M_{PX} = M_{XP}M_{XX}$ (Formula (33)), we have

$$M_{XP} = M_{XX}M_{PX}M_{XX}^{-1}; \quad (39)$$

hence, M_{XP} and M_{PX} have the same eigenvalues; since $M_{PX} = M_{XP}^T$, these eigenvalues must be real, and those of M_{XP}^2 must be ≥ 0 .

For completeness, we still need to discuss what happens when $\Omega_X^h = \Omega_P$. In view of Formulas (28) and Equivalence (24), this means that (31) reduces to equality

$$(M/M_{PP})(M/M_{XX}) = I_{n \times n}$$

that is, by (35), $M_{XX}M_{PP} = I_{n \times n}$. Taking (38) into account, we must thus have $M_{XP}^2 = 0$, which does not imply that $M_{XP} = 0$. We are in the presence of states (21) in this case, saturating the Heisenberg inequalities.

6. Discussion and Outlook

Our discussion of polar duality suggests that a quantum system localized in the position representation in a set X cannot be localized in the momentum representation in a set smaller than that of its polar dual X^h . The notion of polar duality thus appears informally as a generalization of the uncertainty principle of quantum mechanics, as expressed in terms of variances and covariances (see [23]). The idea of such generalizations is not new, and can already be found in the work of Uffink and Hilgevoord [25,26]; see Butterfield's discussion in [27]. It would certainly be interesting to explore the connection between convex geometry and quantum mechanics, but very little work has been conducted so far.

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