

Article

# Prescribed Performance Rotating Formation Control of Multi-Spacecraft Systems with Uncertainties

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**Abstract:** This paper investigates the problem of rotating formation control for multi-spacecraft systems with prescribed performance in the presence of model uncertainties. Firstly, The spacecraft dynamics containing unmodelled parts is described in a polar coordinate system, which is to solve the problem of the controllable angular velocity of rotating formation. Then, the prescribed performance control method is improved by developing new prescribed performance functions. Based on the improved prescribed performance control method, the distributed controller is designed for multi-spacecraft systems to achieve rotating formations with prescribed performance, i.e., the formations error converges to a predefined arbitrarily small residual set, with convergence time no less than a prespecified value. And an RBF neural network is used to fit the unmodelled components of the spacecraft dynamics. Compared with the existing works of literature, this paper not only solves the robust prescribed performance rotating formation control of multi-spacecraft system, but also achieves rotating formation with adjustable angular velocity. Finally, the Lyapunov approach is employed for convergence analysis, and simulation results are provided to illustrate the effectiveness of the theoretical results.

**Keywords:** rotating formation; leader-following; finit-time control; prescribed performance control; controllable angular velocity



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## 1. Introduction

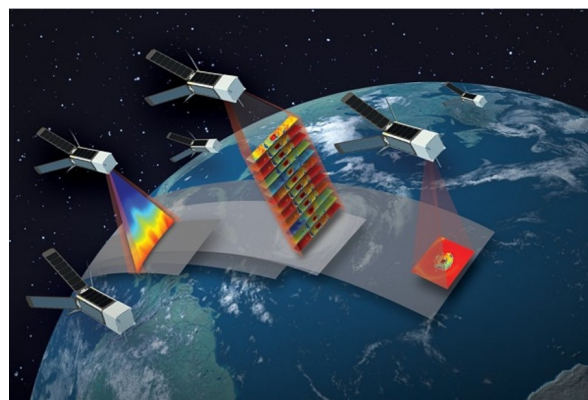
With the advancement of technology, the international community pays more and more attention to space development, consequently, the cluster of small satellites are closely followed [1–3]. For example, in 2016, NASA plans to launch a cluster of small satellites into orbit around the Earth in order to conduct more in-depth studies of the Earth's climate and weather patterns (see Figure 1), the program known as “Time-Resolved Observations of Precipitation structure and storm Intensity with a Constellation of Smallsats” (TROPICS) [4]. To this day, the project is being slowly progressed. One of the key technologies for the cluster of small satellites is rotating formation of multi-satellites, i.e, Multiple following satellites in a formation configuration rotating around the leader satellite under the influence of gravity. It is used in a wide range of tasks such as surveillance, exploration, communication, combat, etc. However, as space mission requirements became more complex, electronics and communications technologies became more advanced, and professionals were placing higher demands on the quality of rotating formation, specifically its steady-state error, overshoot, and convergence speed, as well as its robustness. Those are both essential to the task's success.

The rotating formation problem [5] refers to multi-spacecraft systems in a formation configuration rotating around the center of a circle under the action of an actuator. There

are already some research results. Lin and Jia [5] defined the concepts of rotating consensus and rotating formation in a complex coordinate system using a second-order integrator as a study firstly, and designed a control protocol to implement rotating formation motion of multi-agent systems using kinematic knowledge of circular motion. And the results were extended to three dimensions in [6]. The aforementioned rotating formation controller [5,6] was based on undirected network topology, especially, the rotating formation motion under directed network topology was studied in [7]. With mixed uncertainties and communication delay between agents, a distributed rotating formation control scheme was presented for multi-spacecraft systems based on  $H_\infty$  control theory in [8]. Further, the problem of rotating formation with nonuniform delays was investigated in [9,10]. Founded on [7], article [11] investigated the rotating consensus of multi-spacecraft systems with signed directed graphs. article [12] addressed the rotating consensus of multi-agent systems with restricted communication, and article [13] studied rotating encirclement formations of second-order multi-agent systems with communication noises.

The results of the survey indicate the outstanding contribution of scholars on rotating formations in terms of communication topology and time delay, while the quality of the rotating formation is equally important, it affects the accuracy of satellites operations such as surveillance [14–16], communications [17,18], combat [19,20], etc. Literature [21] was concerned with the Control problem of a single-link flexible-joint robotic manipulator with dynamic uncertainties, considering the problem of pre-defined steady-state error and overshoot of the control system. Article [22] investigated the multi-agent formation problem with pre-defined steady-state error and overshoot. Formation of underwater unmanned aerial vehicles with pre-determined steady-state error and overshoot was investigated in [23]. And paper [24] studied formation control problem of underactuated surface vessels (USVs) with prescribed performance. There have also been many studies [25–27] in preset the steady-state error and overshoot of control systems, but these are few considered in rotating formations problems.

In addition, robustness is also an indicator of the quality of the rotating formation. The literature [8] first considers the problem of robustness of multi-agent consensus, which takes into account the effect of external disturbances. But there are still many problems need to be solved, such as the model uncertainties of nonlinear systems. Therefore, there is an urgent requirement to study control schemes that can improve the quality of rotating formations.



**Figure 1.** TROPICS project.

In summary, the problem of prescribed performance and uncertainties from unmodelled parts in each spacecraft's dynamics in leader-following rotating formations are concerned in this paper. Prescribed performance control methods are used to ensure that the rotating formation control system meets the preset overshoot, steady-state error and convergence time. RBF neural networks are used to approximate uncertainties. The main contributions of this article are as follows:

(1) This paper deals with the prescribed performance robust rotating formation problem of multi-spacecraft systems with controllable angular velocity. To realize rotating formation motion with controlled angular velocity, spacecraft dynamics in a polar coordinate system including unmodelled parts is established. The dynamics contain the angular velocity state of the spacecraft. Compared with the existing works of [5–8] where the angular velocity is assumed to be a fixed value “1”, the formation in this paper can achieve any specified angular velocity rotating.

(2) A prescribed performance rotating formation controller is designed for the uncertain multi-spacecraft system, which can not only comply with the preset overshoot and steady-state error metrics of the closed-loop error system but also make the convergence time less than the preset value. By modifying the exponential function to a power-of-two function, a novel prescribed performance function is developed, which not only decreases computational complexity but also can provide the preset time metric. Compared with the works of [22,24,25,28–30] which can only guarantee control system stability. The proposed prescribed performance function-based controller in this paper is constructed to ensure various performance metrics for the rotating formation of the multi-spacecraft system.

(3) The neural networks-based approximator is constructed to estimate the lumped nonlinear model uncertainties of spacecraft. Meanwhile, the adaptive laws of neural network parameters are updated online. To this end, the robustness of the closed-loop error system is effectively enhanced. Compared with the works of [5–9,12] which only achieve the rotating formation of linear normal systems.

The rest of this paper is organized as follows: Section 2 introduces some preliminaries and presents the dynamics model of the spacecraft in the polar coordinate system. In Section 3, the main results of distributed rotating formation control based on the prescribed performance control method and RBF neural networks are proposed. In Section 4, simulation examples are designed to verify the effectiveness of the controller. Finally, Section 5 concludes the whole paper and discusses the future research work.

## 2. Preliminaries and Dynamics Model

In this section, some preliminary knowledges of graph theory, coordinate frame, and RBF neural networks are introduced. Then, the spacecraft dynamics model in the polar coordinate system is given.

### 2.1. Preliminaries

#### 2.1.1. Graph Theory

Let  $\mathcal{G} \in (\mathcal{V}, \mathcal{E})$  be a undirected graph with  $n$  nodes, where  $\mathcal{V} = \{v_1, \dots, v_n\}$  is the set of nodes,  $\mathcal{E} = \mathcal{V} \times \mathcal{V}$  is the set of edges, and  $A = [a_{ij}] \in \mathcal{R}^{n \times n}$ ,  $a_{ij} \geq 0$  is a adjacency matrix. The node indexes belong to a finite index set  $\mathcal{I} = \{1, \dots, n\}$ . An edge of  $\mathcal{G}$  is represented as  $e_{ij} = \{v_i, v_j\}$ . If  $e_{ij} \in \mathcal{E}$ ,  $a_{ij} > 0$ , else  $a_{ij} = 0$ . The set of neighbors of node  $v_i$  is denoted by  $\mathcal{N}_i = \{v_j | (v_i, v_j) \in \mathcal{E}\}$ . The Laplacian matrix of the undirected graph is  $L = \mathcal{D} - A \in \mathcal{R}^{n \times n}$ , where  $\mathcal{D} = \mathbf{diag}(\mathcal{D}_1, \dots, \mathcal{D}_n)$  is a diagonal matrix, and  $\mathcal{D}_i = \sum_{j \in \mathcal{N}_i} a_{ij}$ .

If there exists a path between any two nodes, the undirected  $\mathcal{G}$  is connected. And the corresponding Laplacian matrix  $L$  of graph  $\mathcal{G}$  is semi-definite [31].

**Assumption 1.** *Assuming that there are  $n$  spacecraft in multi-spacecraft systems and that the communication topology of the system is connected in this paper.*

#### 2.1.2. Coordinate Frame

For the purpose of the following analysis, two coordinate systems is defined:

Inertial coordinate system  $o_l x_l y_l$ : Multi-spacecraft systems rotation center  $o_l$  is chosen as the origin of the coordinate system. The  $x$ -axis ( $o_l x_l$ ) points in a fixed direction and the direction does not change with time. the  $y$ -axis ( $o_l y_l$ ) is perpendicular to the  $x$ -axis and satisfies the right-hand rule with  $x$ -axis.

Polar coordinate system  $o_e x_e y_e$ : The spacecraft center of mass  $o_e$  is chosen as the origin of the coordinate system.  $o_e x_e$  points from the origin of the inertial coordinate system to the origin of the polar coordinate system.  $o_e y_e$  and  $o_e x_e$  form a right-handed coordinate system. The inertial coordinate system is stationary, but the polar coordinate system is bound to the spacecraft, and moves with the spacecraft all the time, like the body coordinate system of a drone.

### 2.1.3. RBF Neural Network

The RBF neural networks are widely used to approximate uncertainties [32–35]. As indicated in [36], any continuous function  $f(x) : \mathcal{R} \rightarrow \mathcal{R}$  can be approximated by an RBF neural network over a compact set  $\Omega_x$  as

$$f(x) = W^T \boldsymbol{\varphi}(x) + \delta(x) \quad \forall x \in \Omega_x.$$

where  $W \in \mathcal{R}^s$  represents the neural networks weight,  $\delta(x) \in \mathcal{R}$  denotes the fitting error,  $|\delta(x)|$  can become arbitrarily small as the number of basis functions increases [35],  $\boldsymbol{\varphi}(x) \in \mathcal{R}^s$  indicates the basic function vector.

### 2.1.4. Notions

The following notions will be used throughout this paper. The variable subscripts  $l, e$  represent the projection of the vector in the inertial or polar coordinate system, respectively, unless otherwise specified. Let  $\mathcal{R}, \mathcal{R}^n$  and  $\mathcal{R}^{n \times n}$  be the real number set,  $n$ -tuple real vector space and  $n \times m$  real matrix space.  $\| * \|$  is the 2-norm of a vector or a matrix.  $I_n$  is an  $n \times n$  identity matrix [37,38]. The diagonal matrix is denoted as  $\text{diag}(*)$ .  $\otimes$  means the Kronecker product,  $| * |$  is the absolute value of a real number.

The concept of practical finit-time stable is defined in the following:

**Definition 1** (Practical Finit-time Stable [39]). *Consider the nonlinear system  $\dot{x} = f(x(t))$ ,  $x(0) = x_0$ , where  $f : \mathcal{L} \rightarrow \mathcal{R}^n$  is continuous on an open neighbourhood of the origin. The equilibrium  $x = 0$  of nonlinear system is practical finit-time stable if for all  $x(t_0) = x_0$ , there exists  $\epsilon > 0$  and a time  $T(\epsilon, x_0) < \infty$  to make  $\|x(t)\| < \epsilon$ , for  $\forall t \geq t_0 + T$ .*

**Lemma 1** ([40]). *For  $\forall \eta_1 \in \mathcal{R}$  and  $\forall \delta_t > 0$ , the following inequality always holds:*

$$0 \leq |\eta_1| - \eta_1 \tanh\left(\frac{\eta_1}{\delta_t}\right) \leq 0.2785\delta_t.$$

**Lemma 2.** *For  $\forall e \in \mathcal{R}^2$  and  $e \neq \mathbf{0}_2$ ,  $\mathbf{0}_2 = [0, 0]^T$ , there exists  $h \in \mathcal{R}^2$  such that*

$$e^T h = 1. \quad (1)$$

**Proof.** Let  $e = [e_1, e_2]^T$ ,  $h = [h_1, h_2]^T$ .

- (1) If  $e_1 = 0, e_2 \neq 0$ : Since  $e_1 h_1 + e_2 h_2 = 1$ , we have  $h_2 = \frac{1}{e_2}$  and  $h_1$  can take any value.
- (2) If  $e_1 \neq 0, e_2 = 0$ : Since  $e_1 h_1 + e_2 h_2 = 1$ , we have  $h_1 = \frac{1}{e_1}$  and  $h_2$  can take any value.
- (3) If  $e_1 \neq 0, e_2 \neq 0$ : We take the value of  $h_2$  to be 0. Since  $e_1 h_1 + e_2 h_2 = 1$ , we have  $h_1 = \frac{1}{e_1}$ .

Thus, Lemma 2 is proved.  $\square$

## 2.2. Dynamics Model

In many cases the formation spacecraft will operate in the same plane. Therefore, to simplify the analysis, this paper considers spacecraft dynamics in multi-spacecraft systems modeling in the two-dimensional plane, and each spacecraft is represented by a

mass point, then the dynamics model of the  $i$ th spacecraft in the polar coordinate system can be described as:

$$\begin{cases} \dot{r}_i = v_i \\ \dot{\theta}_i = w_i \\ \dot{v}_i = u_{eiv} + r_i w_i^2 + d_{eiv}(r_i, \theta_i, v_i, w_i, t) \\ r_i \dot{w}_i = u_{eiw} - 2v_i w_i + d_{eiw}(r_i, \theta_i, v_i, w_i, t) \end{cases}, \quad i \in \mathcal{I}. \tag{2}$$

The index set of an  $n$ -spacecraft system is defined as  $\mathcal{I} = \{1, \dots, n\}$ .  $r_i \in \mathcal{R}^+, v_i \in \mathcal{R}$  denote the distance of the spacecraft  $i$  from the origin of the inertial system to the spacecraft  $i$  and its derivative, respectively.  $\theta_i, w_i \in \mathcal{R}$  denotes the angle between the polar coordinate system  $x$ -axis of the spacecraft  $i$  and the inertial system  $x$ -axis and angular velocity, respectively.  $u_{eiv}, u_{eiw} \in \mathcal{R}$  denotes the projection of the control input of the spacecraft  $i$  on the  $x$ -axis and  $y$ -axis of the polar coordinate system, respectively.  $d_{eiv}(r_i, \theta_i, v_i, w_i, t), d_{eiw}(r_i, \theta_i, v_i, w_i, t)$  denote unmodelled parts of the spacecraft  $i$ .

**Remark 1.** The states  $r_i, \theta_i$  of the  $i$ th spacecraft in (2) indicates the polar axis and polar angle. The polar axis refers to the distance from the current position of the spacecraft to the center of rotation, which is the origin of the inertial system. And the polar angle refers to the angle between the line connecting the position of the spacecraft and the center of rotation and the  $x$ -axis of the inertial coordinate system. When the rotating formation is realized, although the positions of the spacecraft are changing at all times, their polar axes are fixed, and since the formation configuration is generally fixed, the polar angle difference between any spacecraft is also fixed, which transforms the rotating formation problem into the traditional consensus formation problem. Besides, the angular velocity  $w_i$  is included in the (2), thus rotating formation with controllable angular velocity will be achieved when the states in (2) can track the desired information under the action of the designed controller. In addition,  $r_i$  less than or equal to 0 has no physical meaning. So the constraint  $r_i > 0$  should be noted when designing the controller.

**Remark 2.** The dynamics of the spacecraft in this study are based on the Newtonian dynamics of a mass in a polar coordinate system. In order to study spacecraft rotating formation control problems, many researchers simplified the spacecraft dynamics to a second-order integrator [5–9], however two questions were highlighted. The inability to create controllers that produce rotating formation with controllable angular velocity is the first issue. Consequently, the simplified spacecraft dynamics model in the inertial system is transformed into the polar coordinate system in this article. The oversimplified spacecraft dynamics is the second issue. The Newtonian dynamics of a mass in the polar coordinate system cannot be easily adapted to the spacecraft dynamics model since the latter is nonlinear. And a nonlinear factor that is unknown and challenging to analyze will be present in the spacecraft dynamics when it is immediately transformed to the polar coordinate system. Consequently, this article extends the mass point dynamics under the polar coordinate system to include an uncertainties element  $d_{eiv}(r_i, \theta_i, v_i, w_i, t), d_{eiw}(r_i, \theta_i, v_i, w_i, t)$  in order to get closer to the actual spacecraft dynamics.

Define that  $p_i = [r_i, \theta_i]^T, q_i = [v_i, w_i]^T$ , then (2) can be simplified as the following compact form:

$$\begin{aligned} \dot{p}_i &= q_i, \\ B_i \dot{q}_i &= u_i + f_i + d_i. \end{aligned} \tag{3}$$

where

$$B_i = \begin{bmatrix} 1 & 0 \\ 0 & r_i \end{bmatrix}, f_i = \begin{bmatrix} r_i w_i^2 \\ -2v_i w_i \end{bmatrix}, u_i = \begin{bmatrix} u_{eiv} \\ u_{eiw} \end{bmatrix}, d_i = \begin{bmatrix} d_{eiv}(r_i, \theta_i, v_i, w_i, t) \\ d_{eiw}(r_i, \theta_i, v_i, w_i, t) \end{bmatrix}.$$

Since  $r_i > 0$ ,  $B_i$  is positive definite.

### 2.3. Problem Description

There are  $n$  spacecraft and a leader, some of the spacecraft can receive information from the leader's states, but most of them just know the relative desired state between themselves and their neighbors. Our main objective is to design a distributed controller to make multi-spacecraft systems form a certain configuration to do the circular motion with desired angular velocity in the presence of uncertainties, and the rotating formation control system satisfies various preset performance indicators, such as overshoot, steady state error, convergence time, etc. Before giving the main results, we need to make the following definitions.

**Definition 2.** *multi-spacecraft systems is said to achieve rotating formation with prescribed performance if the value of  $T, \Delta_1, \Delta_2$  is predetermined by the user and the states of all the spacecraft satisfy that*

$$\begin{aligned} \lim_{t \rightarrow T} |r_i - r_j - r_{ijd}| &\leq \Delta_1, \\ \lim_{t \rightarrow T} |\theta_i - \theta_j - \theta_{ijd}| &\leq \Delta_2, \\ \lim_{t \rightarrow \infty} |v_i| &\leq \Delta_3, \\ \lim_{t \rightarrow \infty} |w_i - w_j| &\leq \Delta_4. \end{aligned} \quad (4)$$

where  $r_{ijd}, \theta_{ijd}$  are the deviations of desired relative state between spacecraft  $i$  and spacecraft  $j$ .  $r_{id}$  denotes the desired length of spacecraft  $i$ , then  $r_{ijd} = r_{id} - r_{jd}$ , we have  $r_{ijd} = -r_{jia}$ . Similarly, the desired angle of the spacecraft  $i$  is  $\theta_{id}$ , then  $\theta_{ijd} = \theta_{id} - \theta_{jd}$ .  $w_{id}$  denotes the desired angular velocity of the spacecraft  $i$ , and  $T$  represents the preset convergence time, and  $\Delta_1, \Delta_2$  represent the preset metrics of steady-state error,  $\Delta_3, \Delta_4$  indicates very small positive numbers.

**Remark 3.** *The physical meaning of (4) is that the difference between any two spacecraft's polar axes  $r_i - r_j$  and their polar angles  $\theta_i - \theta_j$ , as time tends to a preset value  $T$ , tends to the desired difference  $r_{ijd}$  and  $\theta_{ijd}$ , and the error between the actual difference and desired difference  $|r_i - r_j - r_{ijd}|, |\theta_i - \theta_j - \theta_{ijd}|$  is smaller than the preset value  $\Delta_1, \Delta_2$ . Additionally, the derivative of the polar axes  $v_i$  and the difference of their angular velocities  $w_i - w_j$ , when time approaches to infinity, tends to a neighborhood containing 0. Only the performance of the errors about the polar axis and the polar angle are preset in Definition 2 because for actual systems, we commonly expect the spacecraft to reach a given position in a preset time  $T$ . If the performance of the angular velocity was further prescribed in the Definition 2, it would be more challenging to build the controller.*

The conditions in (4) mean that multi-spacecraft systems will achieve rotating formation in a preset time, and the formation error is less than the preset value finally. In contrast to the literature [5], the definition in this paper does not require exponential and complex calculations. As shown in Figure 2, the three spacecraft and a leader have realized a rotating formation, and the interrelationship between the spacecraft can be visually represented.

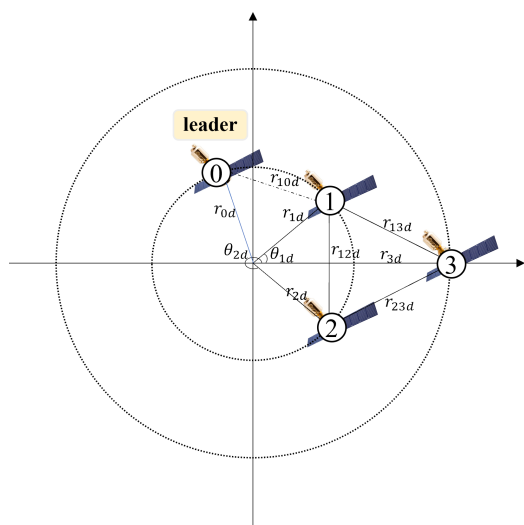


Figure 2. The rotating formation illustration of multi-spacecraft systems.

### 3. Main Results

#### 3.1. Controller Design

Since only certain spacecraft can obtain the leader’s states  $p_0 = [r_0, \theta_0]$  and the majority of spacecraft only know the desired relative state between themselves and their neighbors, the rotating formation error is defined as follows:

$$\begin{aligned}
 e_{pi} &= c_i(p_i - p_0 - p_{i0d}) + \sum_{j=1}^n a_{ij}(p_i - p_j - p_{ijd}), \\
 e_{qi} &= c_i(q_i - q_0 - q_{i0d}) + \sum_{j=1}^n a_{ij}(q_i - q_j - q_{ijd}).
 \end{aligned}
 \tag{5}$$

where  $c_i$  indicates whether spacecraft  $i$  can receive the leader’s states, and if it can,  $c_i = 1$ , otherwise  $c_i = 0$ .  $p_{ijd} = [r_{ijd}, \theta_{ijd}]^T$ ,  $q_{ijd} = \dot{p}_{ijd}$ . And assume the leader is moving along the desired path, the following dynamic equation is satisfied by the leader’s states:

$$\dot{r}_0 = 0, \dot{\theta}_0 = w_{0d}, r_0(0) = r_{0d}, \theta_0(0) = 0.
 \tag{6}$$

where  $w_{0d}$  denotes desired rotating formation’s angular velocity, and  $r_{0d}$  represents the desired radius of rotation.

As a convenience for the stability proof following, define desired state of spacecraft  $i$  as  $p_{id} = [r_{id}, \theta_{id}]^T$  and its derivative is  $q_{id} = \dot{p}_{id}$ , then, there have  $p_{ijd} = p_{id} - p_{jd}$ ,  $q_{ijd} = q_{id} - q_{jd}$ . Since the leader is now following the planned trajectory, i.e.,  $p_0 = p_{0d}$ ,  $q_0 = q_{0d}$ , there have  $p_{id} = p_{i0d} + p_{0d} = p_{i0d} + p_0$  and  $q_{id} = q_{i0d} + q_{0d} = q_{i0d} + q_0$ . Define that

$$e_p = [e_{p1}^T, \dots, e_{pn}^T]^T, e_q = [e_{q1}^T, \dots, e_{qn}^T]^T, p = [p_1^T, \dots, p_n^T]^T, q = [q_1^T, \dots, q_n^T]^T,$$

$$p_d = [p_{1d}^T, \dots, p_{nd}^T]^T, q_d = [q_{1d}^T, \dots, q_{nd}^T]^T, C = \text{diag}(c_1, \dots, c_n), H = (C + L) \otimes I_2.$$

where  $I_2 \in \mathcal{R}^{2 \times 2}$  denotes the unit matrix. Then

$$\begin{aligned}
 e_p &= H(p - p_d), \\
 e_q &= H(q - q_d).
 \end{aligned}
 \tag{7}$$

According to Assumption 1,  $H$  is positive definite, so  $\|p - p_d\|, \|q - q_d\|$  will converge to 0 when  $\|e_p\|, \|e_q\|$  converges to 0, then the Equation (4) will hold, therefore the system will realize the rotating formation. But due to the problems of uncertainties, calculation accuracy, and actuator output accuracy, it is difficult for  $\|e_p\|, \|e_q\|$  to converge to 0 exactly,

so if  $\|e_p\|, \|e_q\|$  can finally satisfy the preset metrics of steady-state error, it is also said that the system achieves the rotating formation.

To achieve the rotating formation motion with prescribed performance, according to the prescribed performance control theory [41], the inequality describing the performance of the rotating formation error is defined as:

$$\begin{aligned}
 & -v_{1i,m}p(t) < e_{pi,m} < v_{2i,m}p(t), m = 1, 2, \\
 & v_{1i,m} = \frac{r+1}{2} + \text{sign}(e_{pi,m}(0))\frac{r-1}{2} \in \mathcal{R}, \\
 & v_{2i,m} = \frac{r+1}{2} + \text{sign}(e_{pi,m}(0))\frac{1-r}{2} \in \mathcal{R}.
 \end{aligned}
 \tag{8}$$

where  $e_{pi,m} \in \mathcal{R}$  denotes the  $m$ th element of the vector  $e_{pi}$  and  $p(t) \in \mathcal{R}$  denotes the performance function that constrains the transient-state convergence process and steady-state error of the control system, which is designed as  $p(t) = (p_0 - p_\infty)e^{-lt} + p_\infty$  in [22,25,26,42–44].  $r \in [0, 1]$  is a constant that constrains the maximum overshoot of the error.  $p_\infty$  can constrain the steady-state error of the rotating formation error. And  $e^{-lt}$  can constrain the transient-state convergence process of the rotating formation error.  $\text{sign}(x)$  is the signum function, which described as:

$$\text{sign}(x) = \begin{cases} 1, & x \geq 0 \\ -1, & x < 0 \end{cases}$$

The conditions in (8) can be graphically depicted by Figure 3.  $e(0) \in \mathcal{R}$  is greater than 0, then  $v_{1i,m} = r, v_{2i,m} = 1$ . If (8) is satisfied,  $e(t)$  will be contained within the function envelope consisting of  $p(t)$  and  $-rp(t)$  over the entire time domain. Then the whole response process of  $e(t)$  will satisfy the user-set performance metrics such as overshoot, convergence speed, steady-state error, etc.

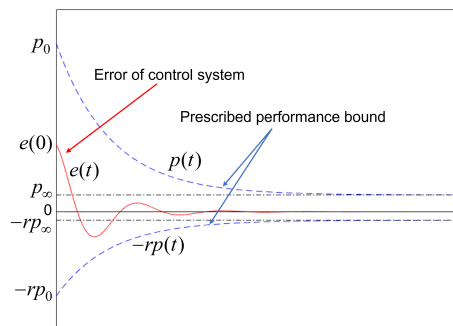


Figure 3. curve of prescribed performance function.

From the formula of  $p(t)$  in [22,25,26,42–44], we have  $p(t) \rightarrow p_\infty$  when  $t \rightarrow \infty$ . Then only when  $t \rightarrow \infty$ , there is  $\|e(t)\| < p_\infty$ . According to Definition 1, the control system is not finite time convergence. To solve this problem, a new prescribed performance function in this paper is described as:

$$p(t) = \begin{cases} p_0(1 - \frac{t}{T})^2 + p_\infty, & t < T \\ p_\infty, & t \geq T. \end{cases}
 \tag{9}$$

The performance function defined by (9) is  $p(t) \rightarrow p_\infty$  when  $t \rightarrow T$ , then when  $t \rightarrow T$ , there is  $\|e(t)\| < p_\infty$ . According to Definition 1,  $e(t)$  is the practical finit-time stable.

For the convenience of controller design, the constraint of the rotating formation error defined by (8) is transformed into an unconstrained form by an error transformation. The error transformation function is defined as:

$$e_{pi,m}(t) = p(t)f(\varepsilon_{pi,m}).
 \tag{10}$$



Let  $\Lambda_{pi,m} = \frac{e_{pi,m}(t)}{p(t)} \in \mathcal{R}$ , then the inverse of  $f(e_{pi,m})$  can be written as:

$$e_{pi,m} = f^{-1}\left(\frac{e_{pi,m}(t)}{p(t)}\right) = f^{-1}(\Lambda_{pi,m}). \tag{11}$$

$f^{-1}(\cdot)$  is taken to be in logarithmic functional form and is written as:

$$e_{pi,m} = \ln\left(\frac{v_{2i,m}(v_{1i,m} + \Lambda_{pi,m})}{v_{1i,m}(v_{2i,m} - \Lambda_{pi,m})}\right) \in \mathcal{R}, m = 1, 2. \tag{12}$$

Its definition domain is:

$$\Theta := (-v_{1i,m}, v_{2i,m}). \tag{13}$$

The equation in (12) shows that if  $e_{pi,m}$  is bounded, then  $\Lambda_{pi,m}$  will satisfies:

$$-v_{1i,m} < \Lambda_{pi,m} < v_{2i,m}.$$

Consequently,  $e_{pi,m}$  satisfies (8), and the problem is then transformed into designing the controller so that  $e_{pi,m}$  is bounded.

Then, we can obtain the derivative of (11) as:

$$\dot{e}_{pi,m} = \frac{de_{pi,m}}{d\Lambda_{pi,m}} \frac{d\Lambda_{pi,m}}{dt} = \zeta_{pi,m}(\dot{e}_{pi,m}(t) + \eta e_{pi,m}(t)). \tag{14}$$

where

$$\begin{aligned} 0 \leq \eta &= -\frac{\dot{p}(t)}{p(t)} < \frac{2}{T}, \\ 0 < \zeta_{pi,m} &= \frac{v_{1i,m} + v_{2i,m}}{(v_{1i,m} + \Lambda_{pi,m})(v_{2i,m} - \Lambda_{pi,m})} \frac{1}{p(t)}. \end{aligned} \tag{15}$$

Define  $\epsilon_{pi} = [\epsilon_{pi,1}, \epsilon_{pi,2}]^T$  as the transformed rotating formation error, the derivative of which is:

$$\dot{\epsilon}_{pi} = \zeta_{pi}(\dot{\epsilon}_{pi} + \eta \epsilon_{pi}). \tag{16}$$

where  $\zeta_{pi} = \mathbf{diag}(\zeta_{pi,1}, \zeta_{pi,2})$ .

From Lemma 2, there exists  $h_i \in \mathcal{R}^2$ , we get  $e_{pi}^T h_i = 1$  when  $e_{pi} \neq 0$ . For the purpose of controller design, the vector function is defined as:

$$g_i = -k_\epsilon \epsilon_{pi} + \zeta_{pi} \epsilon_{pi} + B_i^{-1} d_i - \dot{q}_{id} + \eta \epsilon_{pi}^T \zeta_{pi} \epsilon_{pi} h_i. \tag{17}$$

Fitting  $\|g_i\|$  using the RBF neural networks:

$$\begin{aligned} \|g_i\| &= W_i^T \varphi_i + \delta_i, \\ \varphi_i &= e^{-\frac{(\|e_{pi}\|_{1s-\mu})^2}{2} - \frac{(\|e_{qi}\|_{1s-\mu})^2}{2} - \frac{(\|e_{pi}\|_{1s-\mu})^2}{2}} \in \mathcal{R}^s. \end{aligned} \tag{18}$$

where  $\mu$  is the center of the receptive field, and  $1_s$  denotes s-dimensional 1-vector.

To achieve rotating formation, the controller of the spacecraft  $i$  is designed as:

$$\begin{aligned} u_i &= -B_i(k_\epsilon \epsilon_{pi} + k_q e_{qi} + u_{ti}) - f_i, \\ u_{ti} &= \hat{W}_i^T \varphi_i \tanh\left(\frac{\hat{W}_i^T \varphi_i e_{qi}}{\delta_i}\right), \\ \dot{\hat{W}}_i &= \gamma \|e_{qi}\| \varphi_i. \end{aligned} \tag{19}$$

where  $k_\varepsilon > 0, k_q > \frac{1}{2}, \delta_t > 0, \gamma > 0$ .

**Remark 4.**  $B_i, f_i$  in  $u_i$  are used to linearize the spacecraft dynamics model. The nonlinear terms containing the unmodelled parts are fitted using the RBF neural networks  $u_{ii}$ . The consensus control term  $k_\varepsilon \varepsilon_{pi} + k_q e_{qi}$  is used to achieve the rotating formation motion with the prescribed performance and controllable angular velocity.

Since  $\varphi_i \geq 0$ , then  $\dot{W}_i \geq 0$ , and the initial value of the weight  $\hat{W}_i$  in this paper is vector  $0_s$ , thus, we have

$$\hat{W}_i \varphi_i \geq 0. \tag{20}$$

### 3.2. Stability Analysis

**Theorem 1.** Consider multi-spacecraft systems of  $n$  spacecraft with a connected communication topology. With the controller (19), the performance metric defined by (8) can be satisfied for all time, and multi-spacecraft systems will achieve rotating formation with prescribed performance according to Definition 2, i.e., the states of each spacecraft satisfy (4) with (21):

$$\begin{aligned} \Delta_1 &= 4np_\infty \|\mathbf{H}^{-1}\|, \\ \Delta_2 &= 4np_\infty \|\mathbf{H}^{-1}\|, \\ \Delta_3 &= \|\mathbf{H}^{-1}\| \sqrt{\frac{0.557n\delta_t + \sum_{i=1}^n \frac{1}{2}\delta_i^2}{k_q - 0.5}}, \\ \Delta_4 &= 2\|\mathbf{H}^{-1}\| \sqrt{\frac{0.557n\delta_t + \sum_{i=1}^n \frac{1}{2}\delta_i^2}{k_q - 0.5}}. \end{aligned} \tag{21}$$

**Proof:** According to (3) and (7), we have

$$\begin{aligned} \dot{e}_p &= e_q, \\ \dot{e}_q &= \mathbf{H}(\dot{q} - \dot{q}_d) = \mathbf{H}(\mathbf{B}^{-1}(f + u + d) - \dot{q}_d). \end{aligned} \tag{22}$$

where

$$\begin{aligned} f &= [f_1^T, \dots, f_n^T]^T, d = [d_1^T, \dots, d_n^T]^T, \mathbf{B} = \mathbf{diag}(\mathbf{B}_1, \dots, \mathbf{B}_n), \\ u_t &= [u_{t1}^T, \dots, u_{tn}^T]^T, u = [u_1^T, \dots, u_n^T]^T = -\mathbf{B}(k_\varepsilon \varepsilon_p + k_q e_q + u_t) - f. \end{aligned}$$

Define that  $\varepsilon_p = [\varepsilon_{p1}^T, \dots, \varepsilon_{pn}^T]^T$ , then  $\varepsilon_p = \xi_p(e_q + \eta e_p)$ , where  $\xi_p = \mathbf{diag}(\xi_{p1}, \dots, \xi_{pn})$ . The Lyapunov function is written

$$V = \frac{1}{2} \varepsilon_p^T \varepsilon_p + \frac{1}{2} e_q^T \mathbf{H}^{-1} e_q + \frac{1}{2\gamma} \sum_{i=1}^n \tilde{W}_i^T \tilde{W}_i \tag{23}$$

where  $\tilde{W}_i = W_i - \hat{W}_i$ , then we have

$$\begin{aligned}
 \dot{V} &= \boldsymbol{\varepsilon}_p^T \dot{\boldsymbol{\varepsilon}}_p + \mathbf{e}_q^T \mathbf{B}^{-1}(\mathbf{f} + \mathbf{u} + \mathbf{d}) - \mathbf{e}_q^T \dot{\mathbf{q}}_d - \frac{1}{\gamma} \sum_{i=1}^n \tilde{\mathbf{W}}_i^T \dot{\hat{\mathbf{W}}}_i \\
 &= \boldsymbol{\varepsilon}_p^T \boldsymbol{\zeta}_p(\mathbf{e}_q + \eta \mathbf{e}_p) - \mathbf{e}_q^T (k_\varepsilon \boldsymbol{\varepsilon}_p + k_q \mathbf{e}_q + \mathbf{u}_t) + \mathbf{e}_q^T \mathbf{B}^{-1} \mathbf{d} - \mathbf{e}_q^T \dot{\mathbf{q}}_d - \frac{1}{\gamma} \sum_{i=1}^n \tilde{\mathbf{W}}_i^T \dot{\hat{\mathbf{W}}}_i \\
 &= -k_q \mathbf{e}_q^T \mathbf{e}_q + \mathbf{e}_q^T (-k_\varepsilon \boldsymbol{\varepsilon}_p + \boldsymbol{\zeta}_p \boldsymbol{\varepsilon}_p + \mathbf{B}^{-1} \mathbf{d} - \dot{\mathbf{q}}_d) - \mathbf{e}_q^T \mathbf{u}_t + \eta \boldsymbol{\varepsilon}_p^T \boldsymbol{\zeta}_p \mathbf{e}_p - \frac{1}{\gamma} \sum_{i=1}^n \tilde{\mathbf{W}}_i^T \dot{\hat{\mathbf{W}}}_i \\
 &= \sum_{i=1}^n \mathbf{e}_{qi}^T (-k_\varepsilon \boldsymbol{\varepsilon}_{pi} + \boldsymbol{\zeta}_{pi} \boldsymbol{\varepsilon}_{pi} + \mathbf{B}_i^{-1} \mathbf{d}_i - \dot{\mathbf{q}}_{id}) + \sum_{i=1}^n \eta \boldsymbol{\varepsilon}_{pi}^T \boldsymbol{\zeta}_{pi} \mathbf{e}_{pi} - \frac{1}{\gamma} \sum_{i=1}^n \tilde{\mathbf{W}}_i^T \dot{\hat{\mathbf{W}}}_i \\
 &\quad - k_q \sum_{i=1}^n \mathbf{e}_{qi}^T \mathbf{e}_{qi} - \sum_{i=1}^n \mathbf{e}_{qi}^T \mathbf{u}_{ti}.
 \end{aligned} \tag{24}$$

Substituting (1), (17) and (19) into (24), we have

$$\begin{aligned}
 \dot{V} &= -\frac{1}{\gamma} \sum_{i=1}^n \tilde{\mathbf{W}}_i^T \dot{\hat{\mathbf{W}}}_i + \sum_{i=1}^n \mathbf{e}_{qi}^T (-k_\varepsilon \boldsymbol{\varepsilon}_{pi} + \boldsymbol{\zeta}_{pi} \boldsymbol{\varepsilon}_{pi} + \mathbf{B}_i^{-1} \mathbf{d}_i - \dot{\mathbf{q}}_{id} + \eta \boldsymbol{\varepsilon}_{pi}^T \boldsymbol{\zeta}_{pi} \mathbf{e}_{pi} \mathbf{h}_i) \\
 &\quad - k_q \sum_{i=1}^n \|\mathbf{e}_{qi}\|^2 - \sum_{i=1}^n \tilde{\mathbf{W}}_i^T \boldsymbol{\varphi}_i \mathbf{e}_{qi}^T \tanh\left(\frac{\tilde{\mathbf{W}}_i^T \boldsymbol{\varphi}_i \mathbf{e}_{qi}}{\delta_t}\right) \\
 &= -\sum_{i=1}^n \tilde{\mathbf{W}}_i^T \boldsymbol{\varphi}_i \mathbf{e}_{qi}^T \tanh\left(\frac{\tilde{\mathbf{W}}_i^T \boldsymbol{\varphi}_i \mathbf{e}_{qi}}{\delta_t}\right) - \frac{1}{\gamma} \sum_{i=1}^n \tilde{\mathbf{W}}_i^T \dot{\hat{\mathbf{W}}}_i + \sum_{i=1}^n \mathbf{e}_{qi}^T \mathbf{g}_i - k_q \sum_{i=1}^n \|\mathbf{e}_{qi}\|^2.
 \end{aligned} \tag{25}$$

Under Lemma 1 and (20), (25) is simplified as

$$\begin{aligned}
 \dot{V} &\leq -\sum_{i=1}^n \tilde{\mathbf{W}}_i^T \boldsymbol{\varphi}_i \sum_{j=1}^2 |e_{qi,j}| + 0.2785 \cdot 2 \cdot n \delta_t + \sum_{i=1}^n \|\mathbf{e}_{qi}\| \|\mathbf{g}_i\| \\
 &\quad - \frac{1}{\gamma} \sum_{i=1}^n \tilde{\mathbf{W}}_i^T \dot{\hat{\mathbf{W}}}_i - k_q \sum_{i=1}^n \|\mathbf{e}_{qi}\|^2.
 \end{aligned} \tag{26}$$

From Young-inequality and (18), (26) can be reduced to:

$$\begin{aligned}
 \dot{V} &\leq -\sum_{i=1}^n \tilde{\mathbf{W}}_i^T \boldsymbol{\varphi}_i \|\mathbf{e}_{qi}\| + 0.2785 \cdot 2 \cdot n \delta_t \\
 &\quad + \sum_{i=1}^n \|\mathbf{e}_{qi}\| (\mathbf{W}_i^T \boldsymbol{\varphi}_i + \delta_i) - \sum_{i=1}^n \tilde{\mathbf{W}}_i^T \|\mathbf{e}_{qi}\| \boldsymbol{\varphi}_i - k_q \sum_{i=1}^n \|\mathbf{e}_{qi}\|^2 \\
 &= \sum_{i=1}^n \tilde{\mathbf{W}}_i^T \boldsymbol{\varphi}_i \|\mathbf{e}_{qi}\| - \sum_{i=1}^n \tilde{\mathbf{W}}_i^T \|\mathbf{e}_{qi}\| \boldsymbol{\varphi}_i + 0.2785 \cdot 2 \cdot n \delta_t + \sum_{i=1}^n \|\mathbf{e}_{qi}\| \delta_i - k_q \sum_{i=1}^n \|\mathbf{e}_{qi}\|^2 \\
 &\leq -\left(k_q - \frac{1}{2}\right) \sum_{i=1}^n \|\mathbf{e}_{qi}\|^2 + 0.2785 \cdot 2 \cdot n \delta_t + \sum_{i=1}^n \frac{1}{2} \delta_i^2.
 \end{aligned} \tag{27}$$

If (28) holds,

$$\left(k_q - \frac{1}{2}\right) \sum_{i=1}^n \|\mathbf{e}_{qi}\|^2 \leq 0.2785 \cdot 2 \cdot n \delta_t + \sum_{i=1}^n \frac{1}{2} \delta_i^2. \tag{28}$$

We may have  $\dot{V} > 0$ , then  $\boldsymbol{\varepsilon}, \mathbf{e}_q$  will become unstable and increase with time, if  $\mathbf{e}_{qi}$  keeps increasing until (29) holds,

$$\left(k_q - \frac{1}{2}\right) \sum_{i=1}^n \|\mathbf{e}_{qi}\|^2 > 0.2785 \cdot 2 \cdot n \delta_t + \sum_{i=1}^n \frac{1}{2} \delta_i^2. \tag{29}$$

Then  $\dot{V}$  becomes negative definite and  $\varepsilon, e_q$  will decrease with time until (28) holds. Therefore, as time tends to infinity,  $\varepsilon, e_q$  will converge within the bounded region and the (30) holds ( $k_q > \frac{1}{2}$ ).

$$\begin{aligned} & \lim_{t \rightarrow \infty} \sqrt{\sum_{i=1}^n ((v_i - v_{id})^2 + (w_i - w_{id})^2)} \\ &= \lim_{t \rightarrow \infty} \|\mathbf{q} - \mathbf{q}_d\| \leq \lim_{t \rightarrow \infty} \|\mathbf{H}^{-1}\| \|\mathbf{H}(\mathbf{q} - \mathbf{q}_d)\| \\ &= \lim_{t \rightarrow \infty} \|\mathbf{H}^{-1}\| \sqrt{\sum_{i=1}^n \|e_{qi}\|^2} \leq \|\mathbf{H}^{-1}\| \sqrt{\frac{0.557n\delta_t + \sum_{i=1}^n \frac{1}{2}\delta_i^2}{k_q - 0.5}}. \end{aligned} \tag{30}$$

Consequently,

$$\begin{aligned} \lim_{t \rightarrow \infty} |v_i - v_{id}| &\leq \|\mathbf{H}^{-1}\| \sqrt{\frac{0.557n\delta_t + \sum_{i=1}^n \frac{1}{2}\delta_i^2}{k_q - 0.5}}, \\ \lim_{t \rightarrow \infty} |w_i - w_{id}| &\leq \|\mathbf{H}^{-1}\| \sqrt{\frac{0.557n\delta_t + \sum_{i=1}^n \frac{1}{2}\delta_i^2}{k_q - 0.5}}. \end{aligned}$$

Since  $v_{id} = 0$ , we have

$$\lim_{t \rightarrow \infty} |v_i| \leq \|\mathbf{H}^{-1}\| \sqrt{\frac{0.557n\delta_t + \sum_{i=1}^n \frac{1}{2}\delta_i^2}{k_q - 0.5}},$$

$$\lim_{t \rightarrow \infty} |w_i - w_j - w_{ijd}| \leq \lim_{t \rightarrow \infty} (|w_i - w_{id}| + |w_j - w_{jd}|) \leq 2\|\mathbf{H}^{-1}\| \sqrt{\frac{0.557n\delta_t + \sum_{i=1}^n \frac{1}{2}\delta_i^2}{k_q - 0.5}}.$$

Since  $\varepsilon$  converges to within the bounded region, then (8) holds, so we have  $\lim_{t \rightarrow T} |e_{pi,m}| < p_\infty$ . And then  $\lim_{t \rightarrow T} \|e_p\| < 2np_\infty$ , i.e., when the time  $t \rightarrow T$ , the rotating formation error  $e_p$  converges to within the preset steady-state error region. Since  $e_p = \mathbf{H}(\mathbf{p} - \mathbf{p}_d)$ , then there must be:

$$\begin{aligned} & \lim_{t \rightarrow T} \sqrt{\sum_{i=1}^n ((r_i - r_{id})^2 + (\theta_i - \theta_{id})^2)} \\ &= \lim_{t \rightarrow T} \|\mathbf{p} - \mathbf{p}_d\| \\ &\leq \lim_{t \rightarrow T} \|\mathbf{H}^{-1}\| \|\mathbf{H}(\mathbf{p} - \mathbf{p}_d)\| \leq 2np_\infty \|\mathbf{H}^{-1}\|. \end{aligned} \tag{31}$$

Consequently,

$$\begin{aligned} \lim_{t \rightarrow T} |r_i - r_{id}| &\leq 2np_\infty \|\mathbf{H}^{-1}\|, \\ \lim_{t \rightarrow T} |\theta_i - \theta_{id}| &\leq 2np_\infty \|\mathbf{H}^{-1}\|. \end{aligned}$$

Then, we have

$$\lim_{t \rightarrow T} |r_i - r_j - r_{ijd}| \leq \lim_{t \rightarrow T} (|r_i - r_{id}| + |r_j - r_{jd}|) \leq 4np_\infty \|\mathbf{H}^{-1}\|,$$

$$\lim_{t \rightarrow T} |\theta_i - \theta_j - \theta_{ijd}| \leq \lim_{t \rightarrow T} (|\theta_i - \theta_{id}| + |\theta_j - \theta_{jd}|) \leq 4np_\infty \|\mathbf{H}^{-1}\|.$$

In summary, the following inequality holds:

$$\begin{aligned} \lim_{t \rightarrow \infty} |v_i| &\leq \|\mathbf{H}^{-1}\| \sqrt{\frac{0.557n\delta_t + \sum_{i=1}^n \frac{1}{2}\delta_i^2}{k_q - 0.5}}, \\ \lim_{t \rightarrow \infty} |w_i - w_j - w_{ijd}| &\leq 2\|\mathbf{H}^{-1}\| \sqrt{\frac{0.557n\delta_t + \sum_{i=1}^n \frac{1}{2}\delta_i^2}{k_q - 0.5}}, \\ \lim_{t \rightarrow T} |r_i - r_j - r_{ijd}| &\leq 4np_\infty \|\mathbf{H}^{-1}\|, \\ \lim_{t \rightarrow T} |\theta_i - \theta_j - \theta_{ijd}| &\leq 4np_\infty \|\mathbf{H}^{-1}\|. \end{aligned} \tag{32}$$

The relationship between  $4np_\infty \|\mathbf{H}^{-1}\|$  and  $p_\infty$  allows us to infer that  $4np_\infty \|\mathbf{H}^{-1}\|$  represents the steady-state error that has been prescribed. As a result, the multi-spacecraft system can accomplish rotating formations with prescribed performance.

The proof of Theorem 1 is complete.

#### 4. Simulation Results

To illustrate the obtained theoretical results, numerical simulations will be given in this section. Figure 4 is a communication topology for multi-spacecraft systems that have four spacecraft and a leader, where only spacecraft 1 can receive the leader’s states.

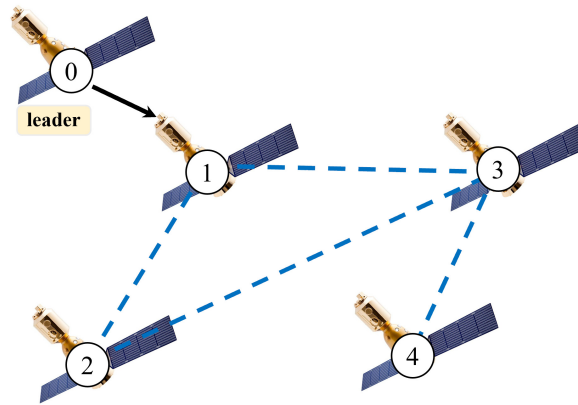


Figure 4. The communication topology among the multi-spacecraft system.

The angular velocity of the leader is  $w_{0d} = 0.4$ , and radius of rotation of the leader is  $r_{0d} = 7$ . Four spacecraft form an equilateral triangle configuration for rotating formation around the center of the circle, then, the system’s parameter are shown in Table 1.

Table 1. parameter of multi-spacecraft system.

Spacecraft Number	Initial States ( $p_i(0)$ )	Derivative of the Initial States ( $q_i(0)$ )	Desired States ( $p_{id}(0)$ )
1	$[6, \frac{\pi}{3}]^T$	$[3, 0]^T$	$[4, 0.4t]^T$
2	$[8, -\frac{\pi}{3}]^T$	$[1, 7]^T$	$[4\sqrt{3}, 0.4t + \frac{\pi}{6}]^T$
3	$[6, 0]^T$	$[4, -2]^T$	$[4, 0.4t + \frac{\pi}{3}]^T$
4	$[10, \frac{\pi}{6}]^T$	$[0, 5]^T$	$[2\sqrt{3}, 0.4t + \frac{\pi}{6}]^T$

The desired difference  $p_{ijd}$  and  $q_{ijd}$  can be calculated from the desired states  $p_{id}$  and  $q_{id}$ .

The number of basis functions is  $s = 80$ , select  $\mu$  evenly from  $[-5, 5]$ , and Both all spacecraft's parameters in the controller (19) are chosen as follows:

$$k_e = 8, k_q = 15, \delta_t = 0.1, \gamma = 0.5.$$

The parameter in (8) is defined as  $r = 1$ . The parameters of (9) are chosen as

$$p_0 = 20, p_\infty = 1.5 \times 10^{-4}, T = 10. \tag{33}$$

According to (33), the control system's maximum steady-state error should be less than  $1.5 \times 10^{-4}$ , its convergence time should be less than 10, and its maximum overshoot should be less than 100%.

Assuming that the unmodelled components of the spacecraft  $i$  is

$$d_i = \begin{bmatrix} 0.2i \cos(0.1t) + 5w_i \cdot \sin\left(0.01t + \frac{i\pi}{5}\right) + 0.05i \sin(0.3t + 2.5) \\ w_i \cdot \cos\left(0.05t + \frac{i\pi}{6}\right) + (5 + 0.1i) \sin(0.02t) + v_i \cdot \sin(0.1t + 1.5) \end{bmatrix}.$$

Simulation results are shown in Figures 5–7. Figure 5 shows the trajectory of all spacecraft in 30 s, from which it is clear that multi-spacecraft systems eventually forms desired configuration to do circular motion around the origin. Figure 6 shows the time response curves of the rotating formation error  $e_{pi}$  for all the spacecraft. The response process of this error reflects various performances of the system such as steady-error and convergence time of the formation control system, the convergence time and steady-error of rotating formation error are set in this paper at (33), and it can be seen from the Figure 6 that the system can achieve the rotating formation motion within 10 seconds and the steady-error of the rotating formation is less than  $10^{-4}$ , which fully satisfies the performance preset above, and there is no overshoot in the response process of the error. Figure 7 shows the angular velocity response curves of all spacecraft, from which it can be seen that the final convergence error of angular velocity is less than  $10^{-5}$ . For perspective, the performance metrics of the rotating formations of the multi-spacecraft system are summarized in Table 2. Therefore, the controller designed in this paper not only resists uncertainties but also can meet various preset performance indicators, and finally achieve rotating formation with high precision and controlled angular velocity in preset time.

**Table 2.** performance indicators of multi-spacecraft system.

Spacecraft Number	Limited Overshot	Practical Overshot	Desired Steady-State Error
1	100%	0.01%	$1.5 \times 10^{-4}$
2	100%	0.01%	$1.5 \times 10^{-4}$
3	100%	0.0075%	$1.5 \times 10^{-4}$
4	100%	0.175%	$1.5 \times 10^{-4}$
Spacecraft Number	Practical Steady-State Error	Desired Convergence Time(s)	Practical Convergence Time(s)
1	$1.2 \times 10^{-5}$	10	9.82
2	$1.4 \times 10^{-5}$	10	9.88
3	$9.2 \times 10^{-6}$	10	9.83
4	$8 \times 10^{-6}$	10	9.86

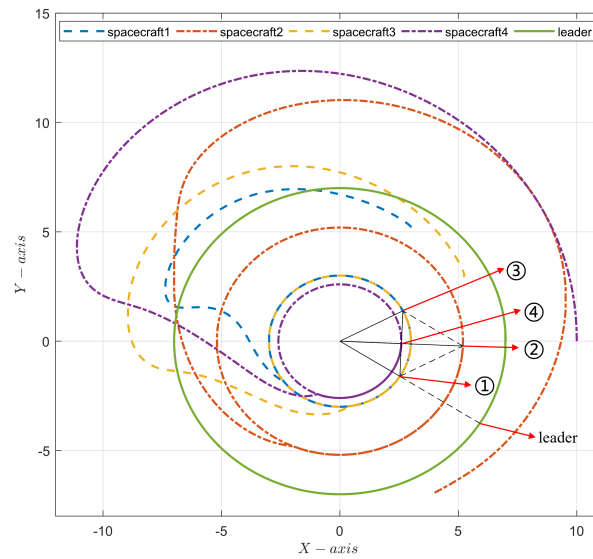


Figure 5. Trajectory of all spacecraft under the inertial coordinate system.

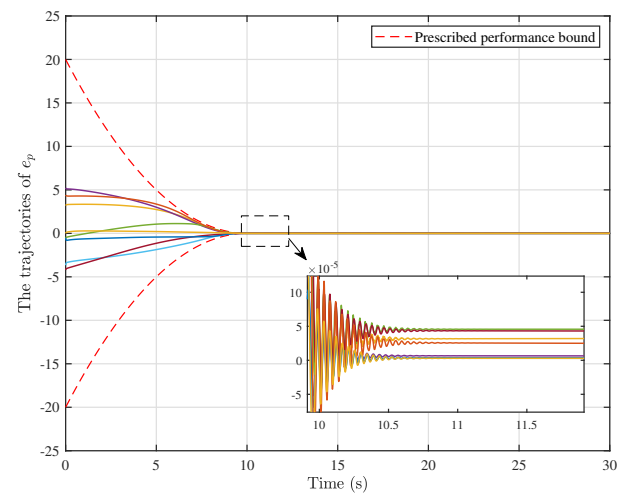


Figure 6. Rotating formation error time response curve of all spacecraft under the controller (19).

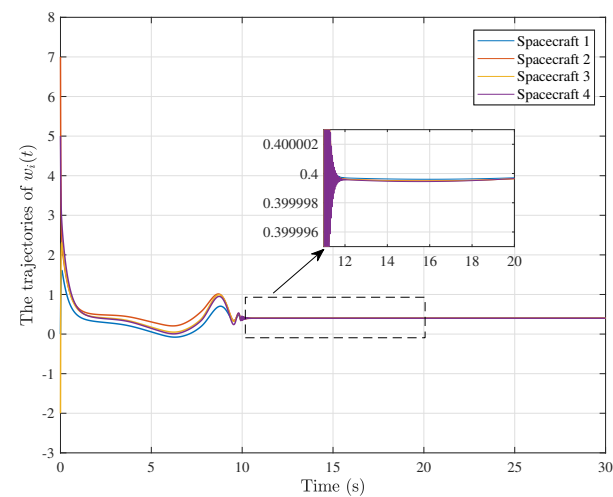


Figure 7. Angular velocity response curves for all spacecraft.

## 5. Conclusions

This paper studied the problem of rotating formation of multi-spacecraft systems with prescribed performance and controllable angular velocity in the presence of model uncertainties. The spacecraft dynamics is presented in a polar coordinate system for the realization of rotating formation with controlled angular velocity. The prescribed performance control method is used to solve the problem of rotating formation with prescribed performance, i.e., the rotating formation error converges to a preset region in a preset time and the transient and steady-state response curves meet preset performance. And an RBF neural network is used to approximate model uncertainties. Finally, simulation results demonstrate the effectiveness of the controller.

In future work, we will consider the problem of controlled angular velocity rotating encirclement control, such as, In the presence of non-matching interference, the controller is designed to make multi-spacecraft systems form a configuration to rotate around a moving target for surveillance or attack.

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