

*Article*



# **Looking at the Modern to Better Understand the Ancient: Is It Possible to Differentiate Mars Pigments from Archaeological Ochres?**

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**Abstract:** This article offers a discussion of the possibility of distinguishing ochres from Mars pigments. The discussion addresses technological, archaeological, and artistic aspects. Natural earth pigments such as ochres, siennas, and umbers have been widely used from the Paleolithic to the present day and still find wide application despite the development of synthetic iron oxide pigment synthesis processes, called Mars pigments. The potential ability of today's analytical techniques to distinguish between two classes of pigments of the same color with very similar chemical composition—but perhaps sufficient for reliable recognition—is also discussed. The paper begins by addressing the proper use of the terms "ochres" and "Mars pigments" and their accurate identification in artworks. It reviews the literature on the chemical–mineralogical characterization of yellow and red iron pigments and analyzes pigment catalogs to understand how companies distinguish ochres from Mars pigments. An experimental analysis using External Reflection Infrared Spectroscopy (FTIR-ER) compared painting samples made with natural ochres and Mars pigments, confirming the literature findings and suggesting future research directions. Key differences such as hematite in yellow ochres and specific spectral peaks in red ochres support the potential of FTIR-ER spectroscopy as a noninvasive tool for distinguishing pigments, especially for fragile artifacts and archaeological applications.

**Keywords:** ochres; Mars pigments; archaeometry; infrared spectroscopy; yellow pigments; red pigments

## **1. Introduction**

Color has always been one of the primary means used by humans to communicate events or states of mind. From the earliest forms of figurative expression, dating back to the Paleolithic with cave paintings, the search for coloring substances or means has been a typical activity of every civilization, from primitive ones onward. Initially, natural mineral and vegetable sources (such as charcoal, earth, or rocks) were used, and with scientific progress, suitable means for chromatic expression were synthesized [\[1–](#page-18-0)[4\]](#page-18-1). Ochres are stable pigments: resistant to light, oxidation, and corrosion, with low reactivity with other pigments and external pollutants. They are characterized by good covering power, low toxicity, and the ability to cover a wide chromatic spectrum (yellow, orange, red, brown, and violet) depending on their mineral composition, impurities, and formation conditions. Thanks to these characteristics, ochres are still widely used today; they can be found in the palettes of contemporary artists and are also used in the construction, coatings, and paint industries despite the development of the synthetic pigment industry [\[5\]](#page-18-2).

Ochre pigments are a natural mixture of iron oxides and clays, combined with other accessory components such as calcite, gypsum, quartz, or other mineral oxides. The



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presence of manganese oxide separates ochres from siennas and umbers [\[1,](#page-18-0)[3,](#page-18-3)[6\]](#page-18-4). Pigments which may be classed as ochres can be found globally. Their exact composition and, consequently, their color and working properties are highly specific to their local geological conditions. Significant producers of ochre pigments include France, Spain, Cyprus, Iran, Italy, Australia, and the USA [\[6\]](#page-18-4).

Iron oxides can be present in varying amounts, mainly as goethite (α-FeOOH), hematite (α-Fe<sub>2</sub>O<sub>3</sub>), magnetite (Fe<sub>3</sub>O<sub>4</sub>), maghemite (γ-Fe<sub>2</sub>O<sub>3</sub>), and lepidocrocite (γ-FeOOH) [\[7\]](#page-18-5). In yellow ochres, the main chromophore is goethite, while in red ochres, hematite is more prevalent. The color of these materials has been attributed to the light absorption due to ligand–metal charge transfer between the Fe $^{3+}$  ion and its  $\mathrm{O}^{2\text{-}}$  or OH<sup>-</sup> ligands, influenced by field d–d transitions [\[5,](#page-18-2)[7](#page-18-5)[,8\]](#page-18-6). Since prehistoric times, ochres have been used either alone, by pulverizing the extracted mineral, or by mixing them with other minerals to achieve a precise chromatic gradation. Additionally, they could undergo thermal treatments to further modify the color, yielding artificial red ochres  $[8,9]$  $[8,9]$ . Evidence of thermal treatment of coloring materials is not well documented, and it is still unclear whether the earliest forms of cooking yellow ochres to obtain red or brown pigments were intentional or not. At temperatures between 350 and 400  $\degree$ C, the reaction that dehydrates goethite and transforms it into disordered hematite occurs. To achieve the crystalline form of the chromophore, temperatures close to 800/900 ◦C are necessary. Mastrotheodoros and Beltsios, 2022 [\[8\]](#page-18-6), highlighted that this process was carried out due to the scarce presence of pure hematite and the difficulty encountered in grinding the hematite mineral. Later, throughout the centuries, the search for new methods of producing ochres from other iron-rich minerals began, primarily for alchemical purposes. To produce yellow ochres, methods were studied that allowed the incorporation of the hydroxyl group into iron oxides. The production of artificial red ochres remained mainly tied to the firing of yellow ochres for centuries despite the existence of alternative methods based either on the firing of ferrous sulfates or other iron salts or on the dissolution of ferrous compounds followed by drying and calcination. To produce brown and violet ochres, manganese dioxide  $(MnO<sub>2</sub>)$  or carbon black was added to further darken the color. Despite these descriptions, the processes for producing artificial iron pigments, known as Mars pigments, became widely disseminated towards the end of the 18th century.

From that time to now, the use of Mars pigments has been documented in painters' palettes. Fragoso et al. (2016) [\[10\]](#page-18-8) found the use of Mars pigments (yellow, red, and black) in the painting *The Hermit* by Antonio Dacosta (1914–1990). Townsend (1993) [\[11\]](#page-18-9) characterized the pigments of J.M.W. Turner (1775–1851) found in his studio after his death and highlighted the presence of Mars pigments. Kampasakali et al. (2007) [\[12\]](#page-18-10) studied the pigments used by some Russian avant-garde painters. Their analyses revealed the presence of yellow, red, and black iron-based pigments, but they could not determine whether they were ochres or Mars pigments. In many other articles consulted, the terms *ochre* and *Mars pigments* were often used as synonyms, making it more challenging to understand the type of material chosen by the artist. The choice of one term over the other is not explained at all. Moreover, it is surprising how few articles report the recognition of Mars pigments even in artworks created during a period when they were presumably present in many painters' palettes. The literature, therefore, indicates the urgency of clarifying the issue. The differentiation between ochres and Mars pigments goes beyond the correct use of terminology, as it implicitly indicates the technology of pigment preparation and provides valuable information about the period of execution of the artwork. Additionally, it could provide information about the authenticity of an archaeological artwork or the presence of undocumented restoration or conservation interventions, which are useful for reconstructing the conservation history of the work of art and the history of past restoration techniques [\[13–](#page-18-11)[15\]](#page-18-12).

The paper looks at the modern to better understand the ancient: Starting from an in-depth literature review focused on the use of the term *Mars pigment* and the chemicalmineralogical characterization techniques used, we sought to learn the criteria adopted by

researchers to distinguish Mars pigments from ochres. Beyond the purely technological, historical, and artistic aspects, the potential capability of analytical techniques to distinguish between two classes of pigments of the same color with a very similar chemical composition—but perhaps sufficient for reliable recognition—is discussed herein. The aim is the research of analytical ways to recognize restorative interventions or forgeries in archaeological artifacts where modern synthetic pigments have been used to integrate or imitate natural ones, respectively. The study begins with a survey of the correct use of the terms *ochres* and *Mars pigments* and their accurate identification in artworks, from archaeological objects to modern paintings, a period that coincides with the spread of Mars pigments. Our research conducted in the catalogs of the main companies selling pigments is laid out to elucidate the specifications used by companies to label a product as *ochre* rather than *Mars pigment*. Finally, the results obtained from a comparison between Mars pigments and ochres, both as such and applied on a substrate, are discussed. The analysis, conducted by infrared spectroscopy, is intended to experimentally verify the results of the literature search and to propose possible avenues of investigation.

#### *1.1. Natural Ochres: Analytical Characterization*

Articles discussing the chemical characterizations of archaeological ochres and ochres used in more recent times were reviewed.

Aura Tortosa et al., 2021 [\[16\]](#page-18-13), analyzed some ochres dating from the Paleolithic-Mesolithic period in Spain. The research highlighted that the composition of archaeological ochres was so varied that major chromophores are often present even in much lower concentrations than other minerals. The same was observed in the study carried out on archaeological findings from the Hoabinhian complex (Vietnam) [\[17\]](#page-18-14) and on Neolithic ochres from Clearwell Caves and Çatalhöyük [\[18\]](#page-18-15). The presence of maghemite was detected in decorated Neolithic ceramics in addition to the accessory minerals [\[19\]](#page-18-16). The authors concluded that its presence could be associated with the production of red ochre by firing yellow ochres rather than using natural red. The literature reports that the most common accessory minerals in ochres are kaolinite [\[7,](#page-18-5)[17,](#page-18-14)[20](#page-18-17)[–22\]](#page-18-18), quartz [\[7,](#page-18-5)[17](#page-18-14)[–23\]](#page-18-19), and calcite [\[7](#page-18-5)[,16,](#page-18-13)[17](#page-18-14)[,19,](#page-18-16)[20,](#page-18-17)[23](#page-18-19)[–25\]](#page-18-20) and more generally silicates [\[18](#page-18-15)[,21](#page-18-21)[,23](#page-18-19)[,25\]](#page-18-20), clays [\[20](#page-18-17)[,22,](#page-18-18)[24\]](#page-18-22), and sulfates [\[21\]](#page-18-21). The most frequently used techniques for the identification of ochres were Raman and FTIR spectroscopy [\[17](#page-18-14)[,19–](#page-18-16)[24\]](#page-18-22), XRD [\[17](#page-18-14)[–19,](#page-18-16)[21–](#page-18-21)[23\]](#page-18-19), and XRF [\[17](#page-18-14)[,23\]](#page-18-19). Infrared spectroscopy was used to a much lesser extent.

#### *1.2. Mars Pigments: Historical Sources*

At the end of the 18th century, new and efficient methods for producing iron pigments with high covering power and uniform particle size, which were also much purer than ochres, were introduced [\[11](#page-18-9)[,26\]](#page-18-23). Harley, 1892 [\[27\]](#page-18-24), highlighted that Mars pigments derive their name from the Latin *crocus martis*, used by alchemists since ancient times to refer to yellow iron oxide. The term *crocus* was associated with the saffron yellow color, and *martis* was internationally used to refer to iron (due to Mars, the God of War) [\[28\]](#page-18-25). Although alchemists began working on synthetic yellow iron pigments several centuries earlier, there are not many references to artists using these pigments. Harley (1892) [\[27\]](#page-18-24) believed this might be due to the high availability of yellow ochre, making a synthetic version unnecessary. Consequently, their subsequent spread might be more due to the development of the chemical industry rather than an actual need for synthetic materials like ochres. The first reference to Mars yellow was found in a patent from 1780, where it was called *oker*, and later, in a 1794 specification, it was referred to as *crocus martis* or *saffron of Mars*.

The first recipe for Mars pigments synthesis, written by Jacques Blockx, dates to 1865 [\[6,](#page-18-4)[27\]](#page-18-24). Generally, they are obtained by precipitating a solution containing a soluble iron salt and alum (to obtain a light yellow) with an alkali, such as soda or lime; the product is then dried at low temperature to avoid altering its color. Extenders such as carbonates, gypsum, or barite or other organic and inorganic pigments are often added to the pigment to enhance its intensity. Currently, the preferred synthesis method requires

the precipitation and hydrolysis of iron salts with subsequent oxidation with an aromatic nitrogen compound in the presence of hydrolyzable multivalent salts [\[6,](#page-18-4)[29–](#page-18-26)[31\]](#page-19-0).

Mars yellow, FeOOH, the first pigment in this category to be synthesized, can exist in a variety of distinct crystalline forms, with the most used in painting being goethite and lepidocrocite. Goethite,  $\alpha$ -FeOOH, is the most common form that iron oxide can take and belongs to the jasper group; it is a dark brown or black mineral from which a yellow powder of various shades can be obtained depending on particle size, and it takes its name from the German writer Goethe [\[6\]](#page-18-4). In 1996, Cornell and Schwertmann [\[1\]](#page-18-0) described several synthesis routes, the most common of which involve iron sulfide or ferric nitrate as the starting reagent [\[6\]](#page-18-4). Lepidocrocite, γ-FeOOH, from the Greek λεπ*ις*´ κρoκη, meaning *saffron-colored flakes*, is instead yellow-orange, with an orthorhombic layered crystalline structure held together by hydrogen bonds. It is commonly found in rust and soils but is much less abundant than goethite and hematite. Several synthesis routes have also been found for this mineral, one of which involves the precipitation of ferrous hydroxide followed by oxidation.

Red Mars can be obtained by calcination of Mars yellow. Mars red is an iron oxide with the formula Fe<sub>2</sub>O<sub>3</sub>. This brick red pigment's shade depends on granularity, calcination, and the possible presence of other pigments with which it is mixed. It ensures stability in light and has significant chemical inertia  $[3]$ . Here too, different crystals with the same empirical formula can be distinguished. Hematite ( $\alpha$ -Fe<sub>2</sub>O<sub>3</sub>), the most common, is a mineral that produces a bright-red powder with a rhombohedral crystalline structure. It can be synthesized in multiple ways, although calcination or dehydration of Mars yellow is generally preferred, a process that converts iron oxide-hydroxide into iron (III) oxide. Heating goethite to a high temperature, around  $350-400\degree C$ , leads to the formation of a disordered form of hematite that then evolves into ferric oxide once about 900  $\degree$ C is reached, with the migration of cations present in the mineral to their sites in the structure. Other synthesis methods described in 1996 by Cornell and Schwertmann include reacting a hydrochloric acid solution with ferric nitrate enneahydrate at 98 ◦C or heating iron sulfate hydrate alone or with calcium oxide [\[6\]](#page-18-4). Calcining lepidocrocite, on the other hand, initially yields maghemite ( $\gamma$ -Fe<sub>2</sub>O<sub>3</sub>), a dark red iron oxide, which, with further temperature increase, recrystallizes as hematite. Maghemite can be also obtained by firing Mars yellow (or yellow ochres) in a reducing atmosphere.

#### **2. Literature Review**

Eastaugh et al. proposed useful definitions of the two classes of pigments [\[6\]](#page-18-4):

Ochres are defined as a "*variably coloured rocks and soils primarily composed of oxides and hydroxides of iron. Colours can vary through various shades of purple, red, orange and yellow. Ochres are secondary deposits occurring either as soils or 'gossans' (the weathered, highly oxidised surface outcrops of ore deposits) which have become enriched in the colour-bearing constituent, usually iron oxides or iron hydroxides. The presence of such minerals promotes a strong and permanent colour, which due to the micron-scale grain size of the particles is not reduced on grinding".*

On the contrary, Mars pigments are "*a group of synthetically produced iron oxide pigments with colours in the yellow-red-violet-black range, the term typically qualified by a colour descriptor (such as 'Mars red'). The mars colours emerged in the eighteenth century"*.

Based on these definitions, characterization studies of pigments used to produce artistic artifacts were searched for in the literature. Articles in which the presence of Mars pigments was identified were selected, and the methodology followed by the authors to differentiate such pigment from ochres was critically reviewed. In addition, it was evaluated whether the use of the term Mars pigment was truthful or the result of possible misunderstanding of the term. The selected articles are shown in Table [1.](#page-6-0)



**Table 1.** Bibliography reporting on the recognition of Mars pigments in works of art. For each reference, the work of art analyzed and the date it was made, the ochres and/or Mars pigments recognized, the analytical techniques used, and the rationale on the use of the two terms are given.

**Table 1.** *Cont.*



**Table 1.** *Cont.*



<span id="page-6-0"></span>Legenda: GC/MS = gas chromatography-mass spectrometry (Py-GC/MS = pyrolysis GC/MS); XRF = X-ray fluorescence; FTIR = Fourier-transform infrared spectroscopy; SEM-EDX = scanning electron microscopy with energy dispersive X-ray spectroscopy; PLM = polarized light microscopy; TG-DSC = thermogravimetric analysis and differential scanning calorimetry; LIPS = laser-induced plasma emission spectroscopy; MA-XRF = macro X-ray fluorescence; HPLC-DAD-ESI-Q-ToF = high-performance liquid chromatography–diode array detector–electrospray ionization–quadrupole–time of flight.

From the review of the articles shown in Table [1,](#page-6-0) it was first observed that the analytical techniques most used by the authors for element analyses were XRF, LIPS, and SEM-EDX. Optical microscopy and SEM (on polished sections, thin sections, or painted surfaces) were employed to investigate the morphology and possible stratigraphy of the painting layer. Raman spectroscopy and FTIR spectroscopy instead identified the molecular composition. References were also found for the use of TG-DSC and XRD for qualitative and quantitative analysis of molecules and detection of crystalline compounds, respectively. Microchemical tests and chromatographic techniques (GC-MS, Py-GC-MS, TLC, and HPLC-DAD-ESI-Q-ToF) were used for the study of binders.

Franquelo et al. [\[24\]](#page-18-22) and Franquelo and Rodriguez [\[33\]](#page-19-10) highlighted that the presence of minerals naturally associated with ochres (such as gypsum, kaolinite, calcite, etc.) could be a primary discriminator between natural and synthetic iron pigments, as Mars pigments contain fillers in much smaller quantities than ochres. According to the authors, the IR signal at 608–609 cm<sup>-1</sup> can also be indicative, as it is more intense for Mars pigments than for hematite. Finally, the presence of elements or minerals related to the raw materials used to produce Mars pigments can be further proof of the type of pigment [\[16\]](#page-18-13). Townsend, 1993 [\[11\]](#page-18-9), highlighted the importance of pigment grain size in differentiating the two materials. Ochres are obtained by grinding the extracted mineral, resulting in a wide range of particle sizes. In contrast, Mars pigments, obtained by synthesis, are characterized by a more uniform and fine particle size distribution.

Cukovska et al. [[32\]](#page-19-11) were unable to determine whether Dicho Zograph used ochres or Mars pigments in their studies but suggested the use of the Raman band at 660 cm<sup>-1</sup> as a possible discriminator since it is not observed in Mars pigments but is common in ochres.

Regarding the use of the term, the present study first ensured that the possible presence of Mars pigments was consistent with their dissemination (late 18th century) [\[8,](#page-18-6)[31\]](#page-19-0) or that their presence in an earlier artwork was due to restoration interventions. This is the case with the *Sacred Family with Angels* (Leonardo Glores, 17th century, Peru) [\[36\]](#page-19-12). The presence of Mars pigments detected by the authors might be due to the restoration interventions the canvas underwent. The same was seen by Ionescu et al.  $[34]$ , who found Mars red in a 13th-century painting, and by Franquelo et al. [\[24\]](#page-18-22), where the same pigment was observed on a 14th-century polychrome sculpture. In both cases, the authors believed that Mars red was used during the restoration work.

#### **3. Review of Commercial Mars Pigment and Ochre**

To evaluate what is currently meant by the terms *ochre* and *Mars pigments*, the catalogs of various companies in the sector were consulted. For each Mars pigment, iron pigment, or ochre found, the technical data sheet and the safety data sheet were reviewed. Commercially available pigments are listed in Table [2.](#page-9-0)

To ensure the anonymity of the companies, the full names of the pigments are not included in Table [2,](#page-9-0) and only the description of the pigment class to which they belong is provided. Numbers were added to differentiate them. For this reason, many pigments are listed with the same name but different declared compositions. If a company classified a synthetic iron oxide as ochre, the term "ochre" was retained. Companies typically name materials according to their class (yellow iron oxide, yellow ochre, Mars pigment, etc.), and they often add specific color characteristics, such as brightness, hue, or particle size, to differentiate them.

Many of the pigments studied share the same name (or a very similar name) and a similarly declared composition. The differences between these materials could be due to variations in particle size or slight color differences caused by the possible presence of additives not listed in the technical and safety data sheets. Natural ochres and earth pigments are in fact characterized by a certain variability in composition, which may not always be disclosed by the manufacturer.



**Table 2.** List of the found commercial Mars pigment and ochres. The capital letter preceding the name of the pigment is the acronym chosen to indicate the company that produces it.

## <span id="page-9-0"></span>**Table 2.** *Cont.*



From the study of iron pigments marketed by companies, it was observed that the trade name does not always coincide with what the pigment should be.

A pigment called red ochre was found to be made from synthetic iron oxides (B—Red ochre). One yellow ochre, according to the data sheets, is composed of chromium antimony titanium rutile (D—Yellow ochre 2), one golden ochre is composed of synthetic zinc ferrite (D—Gold ochre), and finally, one Mars yellow was found to be composed not of synthetic iron oxides but of an azo pigment and natural earth (D—Mars yellow). In some cases, companies reported the addition of other pigments to change the color gradation. Ochres containing natural iron oxides to which are added synthetic iron oxides and zinc oxides, mica, carbon black, or phthalocyanine green were also found. Finally, a red ochre obtained by calcining yellow ochre was found commercially (D—Red ochre); it would thus be an artificial pigment. The addition of any extenders such as barite or calcite was never declared.

In general, consistency between the trade name of the material and its composition was observed for most pigments. However, cases of great difference between the trade name and the chemical composition were also observed. Powder pigments are widely used as conservation materials for historical artifacts. In cases where the trade name does not match the pigment content, this could be a problem. The use of materials that are not philologically correct and not compatible with historic materials increases the risk of unpredictable degradation reactions.

#### **4. Experimental Part**

The possibility of differentiating Mars pigments from natural ochres was investigated using external reflection infrared spectroscopy (FTIR-ER) both on different samples of Mars yellow and red in powder form and on fragments of red ochre and yellow ochre. In parallel, paint layers on paper, one with Mars yellow and another one with yellow ochre, were studied to see if it is possible to differentiate the samples when applied to a surface. The paint layers were prepared in the laboratory 10 years ago, thus showing natural aging of the binder, whose expected duration can range from a few months [\[41](#page-19-14)[–44\]](#page-19-15) to years [\[45\]](#page-19-16). In addition, the paper substrate allowed for noninvasive analysis using FTIR-ATR spectroscopy, which is generally destructive. Previous analyses did in fact show that the mode of analysis did not irreversibly damage the specimen. Thus, the results obtained by a relatively recent technique such FTIR-ER could be compared with those obtained by an established one.

FTIR-ER is increasingly used in the field of cultural heritage because it is portable, noninvasive, and has a wide acquisition range (375–7500 cm $^{-1}$ ) [\[23,](#page-18-19)[46](#page-19-17)[–48\]](#page-19-18), and it has also found wide use in the characterization of archaeological objects [\[49](#page-19-19)[–51\]](#page-19-20). The instrument can provide many indications of surface molecular composition. The main difficulty with the technique is the interpretation of the spectra due to the complexity and variability of pigment and binder mixtures and surface morphology. This difficulty is due to the nature of the technique, which combines different types of reflection and leads to spectral distortions that are absent in more traditional techniques. This mode of IR analysis was chosen to assess its potential in distinguishing ochres from Mars pigments on real artworks since noninvasive techniques are increasingly preferred over traditional ones to investigate archaeological objects that are too valuable or in such a critical state of preservation that they do not even allow the collection of micro-fragments to perform the analysis.

The FTIR-ATR technique, which is also based on a reflection phenomenon, has long been employed in the field of cultural heritage and, more specifically, in archaeology [\[51–](#page-19-20)[53\]](#page-19-21). Its acquisition range varies depending on the type of crystal used, though it is generally limited to the mid-IR region. The spectra obtained with this method are simpler compared to those from FTIR-ER, as they typically do not present distortions and closely resemble the spectra produced by traditional transmission infrared spectroscopy. However, unlike FTIR-ER, this technique requires the collection of at least a small sample, which must be placed in close contact with the analysis crystal by applying pressure to achieve high-quality spectra. This method is therefore unsuitable for archaeological objects from which samples cannot be taken or for materials that are too fragile to withstand the applied pressure without damage. Compared to FTIR-ER, FTIR-ATR generally allows for more superficial analyses without probing deeper layers and is better suited to rough and uneven surfaces, which pose greater challenges for radiation reflection.

#### *4.1. Materials and Methods*

## 4.1.1. Materials

To carry out FTIR-ER analysis of the Mars pigments provided by Kremer Pigmente [\[54\]](#page-19-22), anhydrous KBr and the pigment under investigation were ground and mixed in an agate mortar and pressed. The resulting mixture was subjected to the action of a press until a thin pellet was obtained. The pellet was deposited on a stand and placed in front of the analysis window.

The paint layers were prepared on Fabriano [\[55\]](#page-19-23) cotton paper, using Zecchi [\[56\]](#page-19-24) pigments and bleached crude linseed oil as a binder.

As for the analysis of raw materials and pictorial layers on paper, it was sufficient to place the material subject of study in front of the window. For FTIR-ATR analysis, the paint layers on paper were placed in close contact with the analysis crystal by applying gentle pressure. Table [3](#page-11-0) shows the materials analyzed.

**Powder Pigments and Raw Materials Characterized by FTIR-ER Yellow Pigments Red Pigments** Pigment Description **Pigment** Description **Pigment** Description MY1: Iron Oxide Yellow, MY1: Iron Oxide Yellow, Synthetic iron oxide MR1: Iron Oxide Red 110 M,<br>maize yellow, #48001 Synthetic iron oxide  $\frac{\text{MR1:}}{\text{light, #48100}}$ <br>MP2: Iron Oxide Red 120 M Red pigment  $101.77491$  Fe<sub>2</sub>O<sub>3</sub> MV2: Iron Oxide Vellow 415 MY2: Iron Oxide Yellow 415, Yellow pigment 42 (FeOOH) MR2: Iron Oxide Red 120 M,<br>greenish, #48020  $\frac{101.77491 \text{ Fe}_2\text{O}_3}{148120}$  Red pigment 101.77491 Fe<sub>2</sub>O<sub>3</sub> MY3: Iron Oxide Yellow 940, dark, #48040 C.I. Yellow pigment 42.77492 (FeOOH) red pigment 101.77491 (Fe<sub>2</sub>O<sub>3</sub> ) MR3: Iron Oxide Red 130 B, medium, #48150 Red pigment  $101.77491 \text{ Fe}_2\text{O}_3$ MY4: Iron Oxide Yellow 930, dark, #48045 Yellow pigment 42.77492, FeO(OH) MR4: Iron Oxide Red, clinker red, #48151 red, #48151 MY5: Iron Oxide Yellow, MY5: Iron Oxide Yellow, Iron oxide pigment synthetic MR5: Iron Oxide Red 130 M,<br>maize yellow, #48001 Iron oxide pigment synthetic medium, #48200 medium, #48200 Red pigment  $101.77491 \text{ Fe}_2\text{O}_3$ MY5: Iron Oxide Yellow-Orange 943, Gamma, #48050 FeO(OH); yellow pigment 42.77492 MR6: Iron Oxide Red 160 M,  $\frac{1}{448210}$  Red pigment 101.77491 Fe<sub>2</sub>O<sub>3</sub> MR7: Caput Mortuum Synthetic 180 M,  $#48220$  Red pigment 101.77491 Fe<sub>2</sub>O<sub>3</sub> MR8: Iron Oxide Red 222, dark, #48250 Red pigment  $101.77491 \text{ Fe}_2\text{O}_3$ MR9: Iron Oxide Red, micronized, #48289 Iron oxide (II), 98–100% **Raw materials** YO1: Yellow Moroccan ochre, in pieces, #1164205 Raw materials such as minerals and earth RO1: Red Moroccan ochre, in Red ochre<br>pieces, #12450 Red ochre RO2: Red bole in pieces, #40520 Natural mixture of hematite, quartz, and feldspars. Characterized paper pictorial layers via FTIR-ATR and FTIR-ER MYP: Mars yellow in oil on paper YOP: Yellow ochre in oil on paper

<span id="page-11-0"></span>**Table 3.** Materials analyzed in experimental part.

Figure [1](#page-12-0) shows optical microscope images of the two pictorial layers on paper. The images made it possible to observe the effect of the presence of oil on the surface appearance and the different color tone.





<span id="page-12-0"></span>appearance and the different color tone.

Figure 1. Light microscope image of MYP (a) and YOP (b). The images show the morphology of the pigments when oil is used as a binder and their chromatic gradation. pigments when oil is used as a binder and their chromatic gradation.

#### 4.1.2. Methods

Pigment powders and raw materials were studied by external reflection infrared spectroscopy (FTIR-ER). A Bruker Alpha FTIR instrument with total external reflection module and DTGS detector (spectral range:  $7500-375$  cm<sup>-1</sup>; resolution:  $4 \text{ cm}^{-1}$ ; scans: 200) was used. The spectra were then edited with OPUS software (version 7.2).

mas used: The spectra were then edited with OTGS software (version 7.2).<br>Pictorial layers on paper were characterized by comparing spectra obtained by FTIR-ER and by FTIR-ATR. A Thermo Scientific Nicolet iS10 FTIR instrument in ATR (Attenuated Total Reflection) mode with diamond crystal (spectral range: 4000–600 cm<sup>-1</sup>; resolution:  $4 \text{ cm}^{-1}$ ; scans: 32) was used for this purpose. The spectra were then edited with Omnic software (version 8.3.103).

Omnic software (version 8.5.105).<br>FTIR-ER spectra, as previously discussed, can be difficult to interpret due to distortions or inversions in the shape of the bands and absorption frequencies. This complexity was particularly evident in paint layers on paper, where the combination of support, pigments, and binder complicated the spectra. FTIR-ATR analysis of the same areas facilitated the interpretation. In contrast, this approach was not necessary for the characterization of powder pigments, where the presence of a single component made the spectra significantly<br>cosier to interpret easier to interpret.

Samples on paper were observed with a Maozua USB001 MicroCapture Plus handheld digital microscope. In contrast, this approach was not necessary for the interpretation. In contrast, the interpretation of the interp

#### $c_1$ *4.2. Results*

1.2. *Results*<br>4.2.1. Characterization of Raw Materials Mars Yellow and Yellow Ochre

Figure [2](#page-13-0) shows the FTIR-ER spectra of Mars yellow powder and yellow ochre—raw material. From an initial observation of the FTIR-ER spectra of the analyzed materials, the presence of accessory minerals in the ochres was immediately clear, as the spectra of the ochres showed many more signals than the Mars pigments.

The IR band at about 3100  $cm^{-1}$  observed in all MY spectra is associated with the Fe<sup>2+</sup>-OH vibration of goethite [\[57–](#page-19-25)[59\]](#page-20-0). The other characteristic peaks of goethite observed in all spectra of MY were between 1660 and 1614 cm<sup>-1</sup> (bending of hydroxyl groups [\[57,](#page-19-25)[60](#page-20-1)[,61\]](#page-20-2)), at about 900 and 795 cm<sup>-1</sup> (vibrational deformation of  $Fe^{2+}$ -OH bond [\[57](#page-19-25)[,60](#page-20-1)[,61\]](#page-20-2)), and, finally, Fe-O stretching at 460 cm<sup>-1</sup> [\[28](#page-18-25)[,57](#page-19-25)[,60\]](#page-20-1). The signal at 670 cm<sup>-1</sup>, which is also characteristic of goethite [\[60\]](#page-20-1), was observed only in the spectrum of MY2 and MY3. In MY5 the peaks at about 900 and 795  $cm^{-1}$  were not very intense. In contrast, peaks characteristic of lepidocrocite fell at 1150, 1021, 754, and 616 cm<sup>-1</sup>[\[62–](#page-20-3)[65\]](#page-20-4). These signals, all more or less intense, were observed in all FTIR-ER spectra of Mars pigments. MY5 is the only *4.2. Results* 

pigment to show all four very intense bands, consistent with what is described in the data sheet provided by the manufacturers, which states the use of lepidocrocite to produce the pigment. The peaks at 1150 and 616 cm−<sup>1</sup> were also observed in the spectra of MY3 and MY4. The peak at 1021 cm<sup>-1</sup> was seen in all spectra of MY. Patterns characteristic of calcite (1431 and 875 cm<sup>-1</sup> [\[22\]](#page-18-18)) in MY3 and barite (981 and 613 cm<sup>-1</sup> [\[28\]](#page-18-25)) in MY3 and MY4, which were probably added as fillers, were observed.

<span id="page-13-0"></span>

**Figure 2.** FTIR-ER spectra of yellow pigments. Gray-filled areas mark regions containing the most **Figure 2.** FTIR-ER spectra of yellow pigments. Gray-filled areas mark regions containing the most prominent signals for identifying goethite. Solid lines represent goethite's primary peaks, while prominent signals for identifying goethite. Solid lines represent goethite's primary peaks, while dashed lines indicate lepidocrocite's characteristic peaks. dashed lines indicate lepidocrocite's characteristic peaks.

of signals in the overtones region (4000–7500 cm<sup>-1</sup>) absent in the spectra of MY was first observed. These bands were attributed to kaolinite  $[66,67]$  $[66,67]$  (7168, 7111, 7963, and  $[56,67]$ bending or overtones of iron oxides were observed [\[66,](#page-20-5)[68\]](#page-20-7), which have never been observed in Mars pigments. The presence of kaolinite was also confirmed by peaks at 3650, 3624, 3433, 1130 (inverted), 1050 (inverted), 911 (inverted), and 704 cm<sup>-1</sup> [62]. The inverted FTIR-ER signals are due to a spectral distortion typical of the technique. The band at 3100 cm<sup>-1</sup> of iron oxyhydroxide is not clearly distinguishable from the OH stretches of kaolinite, but<br>the harding of the bordganal against 1647 mm<sup>-1</sup> share tagistic of goathite genericalized of the characteristic doublet of  $\alpha$ -FeOOH, the signal at 795 cm<sup>-1</sup> (inverted) is still clearly distinguishable, while the peak at 900 cm<sup>-1</sup> is partially hidden by kaolinite, so a shoulder can be observed, probably identifiable as this characteristic peak. Finally, the two peaks related to Fe-O stretching at 670 and 460 cm<sup>-1</sup> are clearly observable. The lepidocrocite marteristic peaks, present in all mais yenows to a greater or resserved in, are not present in the FTIR-ER spectrum of ochre. The signal at 1333 cm<sup>-1</sup> could be attributed to the presence of barite. From the observation of yellow ochre in the form of raw material (YO1), the presence  $4527 \text{ cm}^{-1}$ ). Between  $4000$  and  $500 \text{ cm}^{-1}$ , peaks due to combinations of stretching and the bending of the hydroxyl group at 1647 cm<sup>-1</sup> characteristic of goethite remains visible. characteristic peaks, present in all Mars yellows to a greater or lesser extent, are not present

#### Mars Red and Red Ochre Figure 3 shows the FTIR-ER spectra of powdered Mars reds and red ochres—raw

<span id="page-14-0"></span>Figure [3](#page-14-0) shows the FTIR-ER spectra of powdered Mars reds and red ochres—raw material.



**Figure 3.** FTIR-ER spectra of red pigments. The solid lines indicate the hematite, goethite, and **Figure 3.** FTIR-ER spectra of red pigments. The solid lines indicate the hematite, goethite, and maghemite major peaks, and the dashed lines indicate the signals of lepidocrocite. maghemite major peaks, and the dashed lines indicate the signals of lepidocrocite.

the ochre spectrum than in the spectra of the Mars pigments, indicating the presence of the occurs of the ochress of the spectra of the spectrum than in the presence of the presence The FTIR-ER spectra show, as a first difference, the presence of many more signals in accessory minerals.

accessory minerals. OH stretching of the maghemite [\[69](#page-20-8)[,70\]](#page-20-9). The very broad signal at about 3430 cm<sup>-1</sup> is due to the presence of hydroxyl groups. Some of the characteristic peaks of goethite (900 and  $(795 \text{ cm}^{-1})$  and lepidocrocite (1150 and 1021 cm<sup>-1</sup>) have been found in all MR, while the peaks at 670 and 460 cm−1 for goethite and 754 and 616 cm−1 for leptdocrocite nave never<br>been observed. The signal at 1660 cm<sup>-1</sup> seen in the goethite spectra shifted to lower values  $\frac{3}{2}$  in the MR, at about 1630 cm<sup>−1</sup>. Its presence is always related to the bending of goethite hydroxyl groups [70]. Three very intense peaks have always been observed in the FTIR In all FTIR-ER spectra of MR, except for MR3, a weak band was observed at about 3740 cm $^{-1}$ , which was absent in the two characterized red ochres. The peak represents the peaks at 670 and  $\overline{460}$  cm $^{-1}$  for goethite and 754 and 616 cm $^{-1}$  for lepidocrocite have never spectra of the MR: a doublet at about 580 and 550  $\text{cm}^{-1}$  and a single signal at 481/479  $\text{cm}^{-1}$ . The peaks at 550 and 480 cm−<sup>1</sup> are due to structural vibrations of hematite [\[62,](#page-20-3)[71](#page-20-10)[–73\]](#page-20-11). The peak at about 580 cm<sup>-1</sup> could be due to stretching of the maghemite [\[62\]](#page-20-3). Another signal characteristic of hematite fell at 643 cm−<sup>1</sup> [\[62\]](#page-20-3) but was not observed in any spectrum. In all samples except MR4 and MR5, peaks were seen that could be attributed to the addition of carbonates such as calcite (1415 cm−<sup>1</sup> [\[22\]](#page-18-18)).

> Also, in the case of the red ochres, the presence of kaolinite was observed due to its characteristic signals between 3600 and 3400 cm<sup>-1</sup> and the overtones listed above. In RO1, the characteristic peaks of hematite at 559 and 473 cm $^{-1}$  (both inverted) are clearly

distinguishable, and the absorbance at 643 cm<sup>-1</sup>, which is absent in the IR spectra of Mars reds, was also observed. The bands indicating the presence of goethite clearly observed in the spectrum are at 1630, 897, and 810 cm<sup>-1</sup>. Peaks characteristic of lepidocrocite and maghemite were not found. Thus, the RO1 pigment appears to be composed of hematite, which were not found. Thus, the RO1 pigment appears to be composed of hematite, goethite, kaolinite, and quartz (plus other silicates and aluminates not clearly identifiable).  $\mu$  Bochine, and quariz (plus once sincates and annihilates not clearly identifiable).<br>As for RO2, characteristic signals of hematite were observed at 642, 559, and 480 cm<sup>-1</sup>.

while the peak at 580 cm<sup>-1</sup> of maghemite is absent. Again, typical goethite peaks were present, although they are much less intense than in RO1, at 904 and 791 cm<sup>-1</sup>. In addition to iron oxides, kaolinite, quartz, and, in this case, also a carbonate (probably calcite) were present.

## <span id="page-15-0"></span>4.2.2. Characterization of Pictorial Layers

The FTIR spectra (ATR and External Reflection modality) of the two pictorial layers on paper are shown in Figures [4](#page-15-0) and [5.](#page-16-0)



**Figure 4.** FTIR-ATR spectra of MYP and YOP. Solid lines represent goethite's primary peaks, while declear lines indicate lanidographic? characteristic peaks dashed lines indicate lepidocrocite's characteristic peaks. dashed lines indicate lepidocrocite's characteristic peaks.

Starting from the MYP spectra, the FTIR-ATR signals attributable to goethite are at 1662, 902, 792, and 666 cm $^{-1}$  (Figure [4\)](#page-15-0), while the FTIR-ER peaks are at 1647 (gray band), 924, 820, and 670  $\text{cm}^{-1}$  (Figure [5\)](#page-16-0). The characteristic peaks of lepidocrocite at 1162 and 1032 cm−<sup>1</sup> (FTIR-ATR) and 1040 cm−<sup>1</sup> (FTIR-ER) are also clearly observable. The other bands were attributed to the binder [\[47\]](#page-19-26) and paper [\[48,](#page-19-18)[74\]](#page-20-12). For yellow ochre (YOP), ATR signals at 908, 795, and 664 cm $^{-1}$  (reflection mode at 1629, 894 and 443 cm $^{-1}$ ) were attributed to goethite (Figure [4\)](#page-15-0). The characteristic bands of hematite were also seen at 634 cm $^{-1}$  (in ATR) and 578 and 630  $cm^{-1}$  (in FTIR-ER) (Figure [5\)](#page-16-0). In addition to the chromophore peaks attributable to the presence of kaolinite, gypsum and quartz were observed.

<span id="page-16-0"></span>

**Figure 5.** FTIR-ER spectra of MYP and YOP. Gray-filled areas mark regions containing the most regions containing the most prominent signals for identifying goethite. Solid lines represent goethite's primary peaks, while<br>deebed lines indicate lenide westle's abave teristic peaks dashed lines indicate lepidocrocite's characteristic peaks. dashed lines indicate lepidocrocite's characteristic peaks.

of accessory minerals and the absence of characteristic indicators of lepidocrocite. Although the natural presence of lepidocrocite in ochres is widely mentioned in the literature, no<br>the natural presence found among the studies identifies the minoral in the characterization of natural ochres. In addition, in ochres, it is more common to find the presence of accessory minerals, which are present in lower amounts in Mars pigments. Finally, the presence of hematite in ochres was observed, which is absent in MY because it is obtained by firing goethite. The presence of hematite, a mineral that can be found naturally associated with problem in the setting of The present study revealed that yellow ochres differ from MY in terms of the presence references were found among the studies identifying the mineral in the characterization of goethite in ochres, could be the means of distinguishing synthetic from natural iron yellows.

goethic firectics, collid be the means of distinguishing synthetic from natural from yenows.<br>As for red pigments, in contrast to the barite found predominantly in MY, goethite, lepidocrocite, maghemite, and hematite were almost always found in this group of pigments, unlike in the red ochres, where the chromophores present were solely goethite The present study revealed that years are very study revealed that years of the my in terms of the my in ter and hematite.

#### $p_{\text{r}}$  and the absence of absence of  $q_{\text{r}}$  and the absence of characteristic indicators of  $q_{\text{r}}$ **5. Conclusions**

This study aimed to evaluate methods for recognizing ochres and Mars pigments in works of art. Regarding the use of the terms, great confusion has been observed in the literature, as very often, the term Mars pigment is used as a synonym for ochre. The review of pigments on the market showed that the trade name did not always coincide with what the pigment should be, replicating the confusion of terms that emerged from the literature.

In a pigment should be, repheating the contasion of terms that emerged from the increase.<br>The literature review highlighted different ways of distinction between natural and synthetic iron pigments, in many cases on restored surfaces. The presence of accessory minerals is the most-discussed way. Another indicative characteristic is the morphology of pigment grains. By Raman spectroscopy, it seems to be possible to differentiate pigments by observing some characteristic bands.

by observing some characteristic bands.<br>Finally, the experimental part showed that it might be possible to distinguish ochres from Mars pigments with FTIR-ER spectroscopy, too. The presence of accessory minerals in high concentrations and of different types could indicate that the pigment is an ochre. Another way of distinction is the presence of hematite in yellow ochres, which is absent in MY. In contrast, goethite was found in both MR and red ochres. Finally, a peak, at  $642-643$  cm<sup>-1</sup>, assigned to the hematite that is absent in Mars reds, was observed from the study of red ochres. It is possible to study this issue further by analyzing more red ochres and MR to define whether this could be significant in differentiating the two pigments by FTIR-ER, which has the advantage of being portable and noninvasive and thus can be used<br>that fracile shields at directly at archaeological sites with fragile objects or directly at archaeological sites.

The differentiation between Mars pigments and ochres therefore is very complex, but there are some details that seem to make it possible. Techniques such as FTIR-ER have shown some differences between the two types of pigments, and the literature has shown that even with other techniques, it is possible to differentiate the materials either by study that even with other teentiques, it is possible to unferentiate the materials entier by study of grain size or by observation of chemical composition.

In the future, this study could be further developed by expanding the analytical techniques used for the experimental part or by using chemometric methods such as PCA to further investigate the differentiation.

A schematic description of the results of the evaluation proposed in this study is A schematic description of the results of the evaluation proposed in this study is presented in Figure [6.](#page-17-0) presented in Figure 6.1 and the figure 6.1 and the

<span id="page-17-0"></span>

**Figure 6.** Schematic representation of the findings.

analysis, M.C.C.; investigation, C.C.; resources, M.C.C.; data curation, C.C.; writing—original draft for preparation, M.C.C., C.C. and E.R., Writing—feview and editing, C.C. and E.R., Visualization, C.C.<br>supervision, L.R.; project administration, L.R. All authors have read and agreed to the published  $\frac{1}{\sqrt{2}}$  version of the manuscript. **Author Contributions:** Conceptualization, L.R.; methodology, L.R.; validation, C.C. and L.R.; formal preparation, M.C.C., C.C. and L.R.; writing—review and editing, C.C. and L.R.; visualization, C.C.;

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corresponding author. **Data Availability Statement:** The data presented in this study are available on request from the

**Conflicts of Interest:** Authors declare no conflicts of interest.

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